

# New Results in Neutrino Measurements

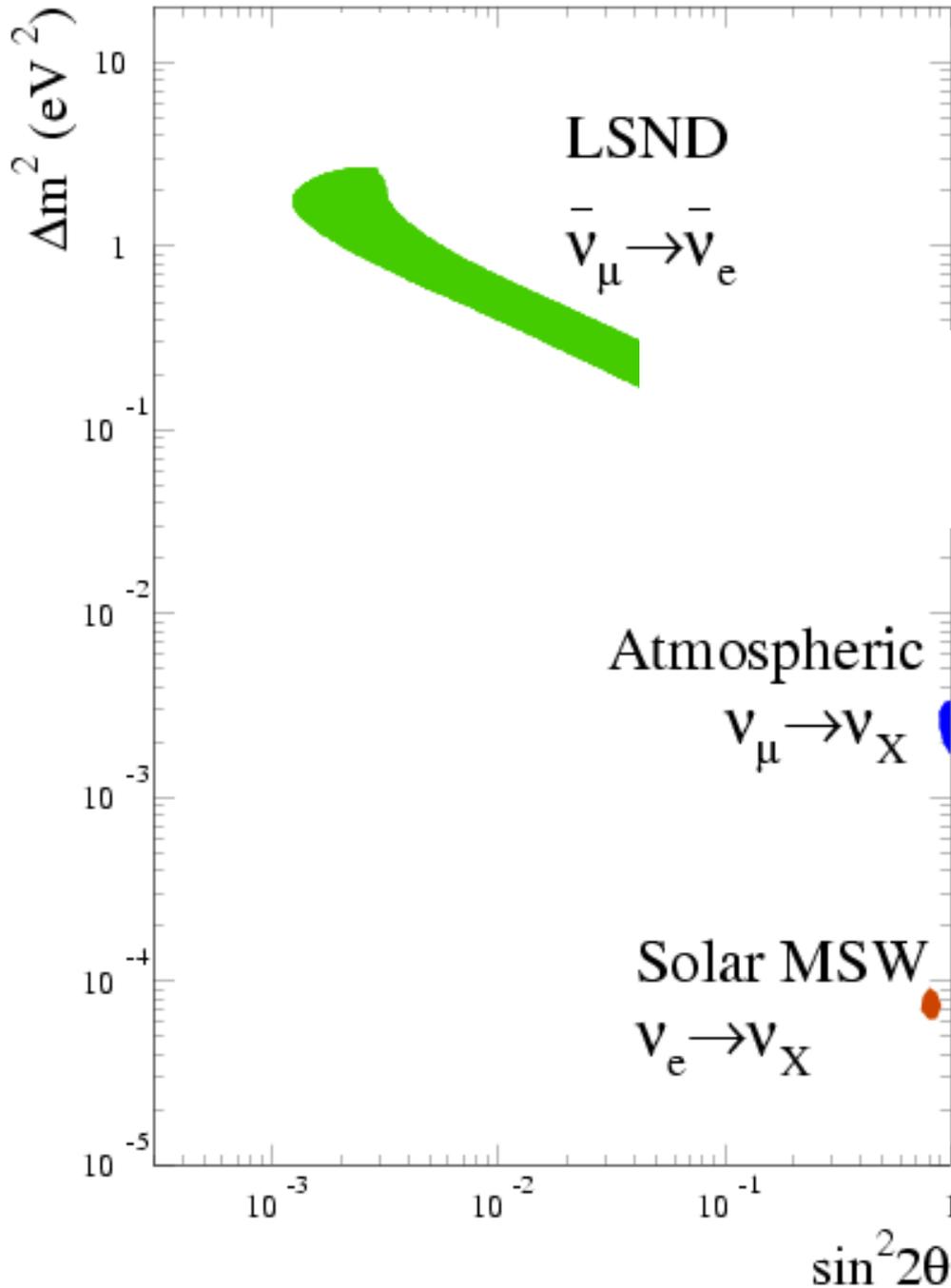
W.C. Louis, LANL

- **Neutrino Past:** Solar & Atmospheric  $\nu$  Oscillations!  
Large  $\Delta m^2$  Neutrino Oscillations?
- **Neutrino Future:** Tritium  $\beta$  Decay, Double  $\beta$  Decay,  
Reactor  $\theta_{13}$ , Accelerator  $\theta_{13}$ , Solar  $\nu$ , &  $\nu$  Cross  
Sections
- **Neutrino Present:** MINOS, MiniBooNE, & K2K  $\sigma$  s
- **Conclusions**

# Neutrino Past

- **“Solar” Neutrino Oscillations!** (Homestake, Kamiokande, SAGE, GALLEX, GNO, SNO, KamLAND)
- **“Atmospheric” Neutrino Oscillations!** (IMB, Kamiokande, SOUDAN, Super-K, MACRO, K2K)
- **Large  $\Delta m^2$  Neutrino Oscillations?** (LSND)

# Current Neutrino Oscillation Signals

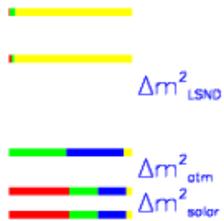


Expt.	Type	$\Delta m^2 \text{ (eV}^2\text{)}$	$\sin^2 2\theta$
LSND	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	$\sim 1$	$\sim 3 \times 10^{-3}$
Atm.	$\nu_\mu \rightarrow \nu_\tau$	$\sim 2 \times 10^{-3}$	$\sim 1$
Solar	$\nu_e \rightarrow \nu_{\mu,\tau}$	$\sim 8 \times 10^{-5}$	$\sim 0.8$

# If the LSND Signal is Confirmed: Physics Beyond the Standard Model & Connections with Astrophysics!

For example:

## 3+2 Sterile Neutrinos



Sorel, Conrad, & Shaevitz (hep-ph/0305255)

**Explain Pulsar Kicks?**

**Explain R-Process in Supernovae?**

**Sterile Neutrinos Explain Dark Matter?**

## MaVaNs & 3+1 Sterile Neutrino

Hung (hep-ph/0010126)

Kaplan, Nelson, & Weiner (hep-ph/0401099)

**Explain Dark Energy?**

## CPT Violation & 3+1 Sterile Neutrino

Barger, Marfatia, & Whisnant (hep-ph/0308299)

**Explain Baryon Asymmetry in the Universe?**

## Quantum Decoherence

Barenboim & Mavromatos (hep-ph/0406035)

## Lorentz Violation

Kostelecky & Mewes (hep-ph/0406255)

## Extra Dimensions

Pas, Pakvasa, & Weiler (hep-ph/0504096)

## Sterile Neutrino Decay

Palomares-Ruiz, Pascoli, & Schwetz (hep-ph/0505216)

# Probability of Neutrino Oscillations

$$P_{\alpha\beta} = \delta_{\alpha\beta} - 4 \sum_i \sum_j |U_{\alpha i} U_{\beta i}^* U_{\alpha j}^* U_{\beta j}| \sin^2(1.27 \Delta m_{ij}^2 L/E_\nu)$$

As  $N$  increases, the formalism gets rapidly more complicated!

$N$	$\#\Delta m_{ij}^2$	$\#\theta_{ij}$	$\#\text{CP Phases}$
<b>2</b>	<b>1</b>	<b>1</b>	<b>0</b>
<b>3</b>	<b>2</b>	<b>3</b>	<b>1</b>
<b>6</b>	<b>5</b>	<b>15</b>	<b>10</b>

# Fundamental Questions to be Answered

(Neutrinos are still largely unknown!)

- What are the neutrino masses & hierarchy?
- What are all of the neutrino mixings?
- Are neutrinos Dirac or Majorana?
- Is CP conserved in the lepton sector?
- Is CPT conserved in the lepton sector?
- Do light, sterile neutrinos exist?
- What is the resolution of the 3- $\Delta m^2$  paradox?

# Neutrino Future

(Neutrino Satellite Meeting on Saturday & Sunday!)

- **Tritium  $\beta$  Decay:** KATRIN
- **Double  $\beta$  Decay:** CUORE, EXO, GERDA, MAJORANA, MOON, Super NEMO
- **Reactor  $\theta_{13}$ :** Double CHOOZ, Braidwood, Daya Bay
- **Accelerator  $\theta_{13}$ :** T2K, NOvA, BNL, KASKA
- **Solar  $\nu$ :** Borexino, KamLAND/7Be, CLEAN, HERON
- **Large  $\Delta m^2$  Oscillations:** BooNE, OscSNS
- **Cross Sections:**  $\nu$  SNS, MINERvA, FINeSSE, J-PARC

# Neutrino Present

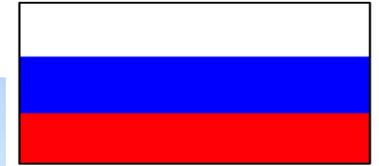
- **MINOS:** A Definitive Measurement of Atmospheric Neutrino Oscillations (Parallel Session Talk by Tzanakos)
- **MiniBooNE:** A Definitive Test of the LSND Neutrino Oscillation Evidence (Parallel Session Talks by Conrad, Djurcic, Wascko, Nelson, & Aguilar-Arevalo)
- **K2K:** Measurement of Neutrino Cross Sections (Has already confirmed atmospheric oscillations)

# MINOS: A Definitive Measurement of Atmospheric Oscillations

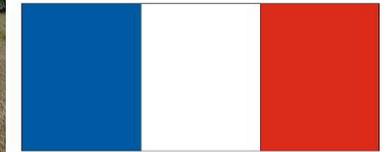
$$\text{Measure } \nu_{\mu} \rightarrow \nu_{x}$$



6 nations



175 physicists



32 institutions



Argonne – Athens – Benedictine – Brookhaven – Caltech – Cambridge – Campinas – Fermilab – College de France – Harvard – IIT – Indiana – ITEP Moscow – Lebedev – Livermore – Minnesota, Twin Cities – Minnesota, Duluth – Oxford – Pittsburgh – Protvino – Rutherford Appleton – Sao Paulo – South Carolina – Stanford – Sussex – Texas A&M – Texas-Austin – Tufts – Univ. College London – Western Washington – William & Mary - Wisconsin

# NuMI/MINOS:

735 km Baseline

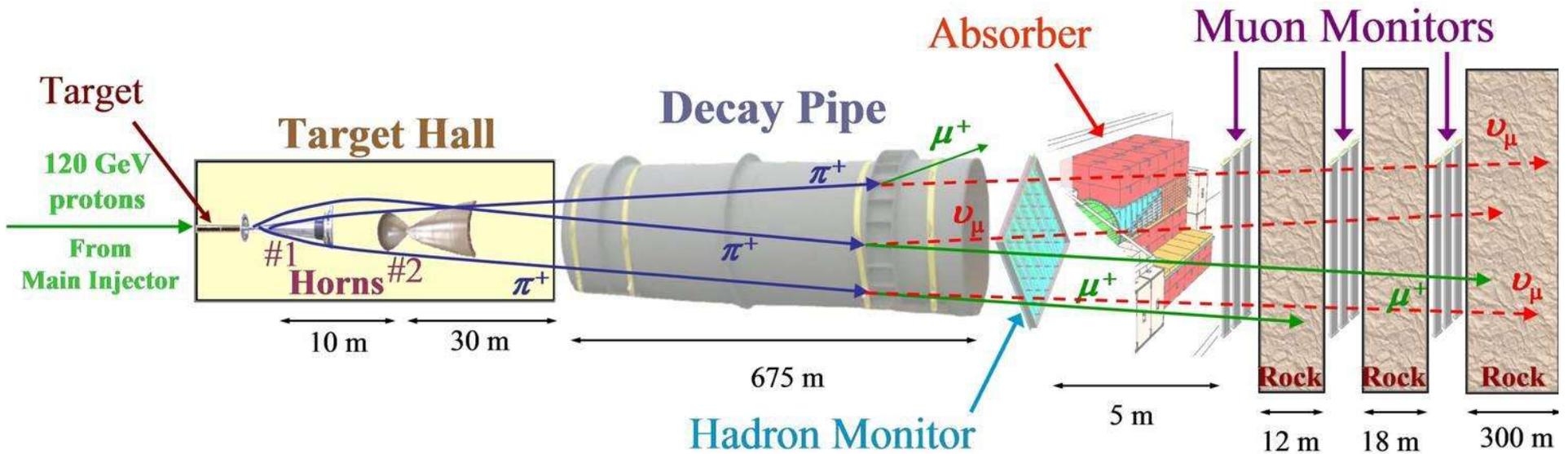
$\nu_{\mu}$  Disappearance

$\nu_{\mu} \rightarrow \nu_e$  Appearance

Sterile Neutrino Search

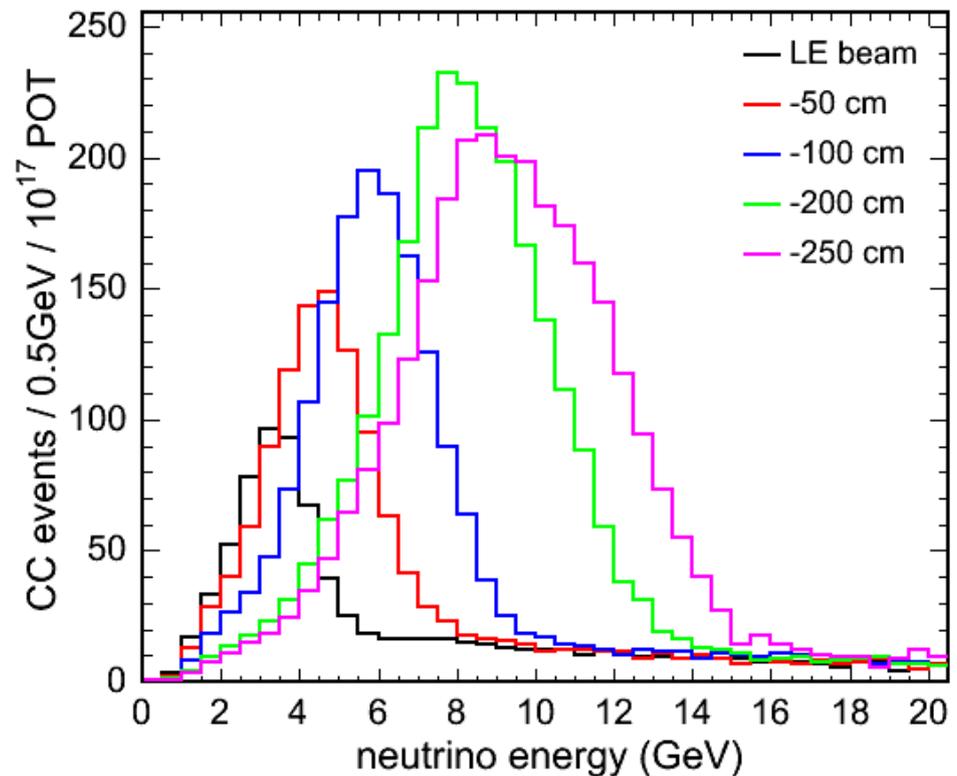
Atmospheric Neutrinos





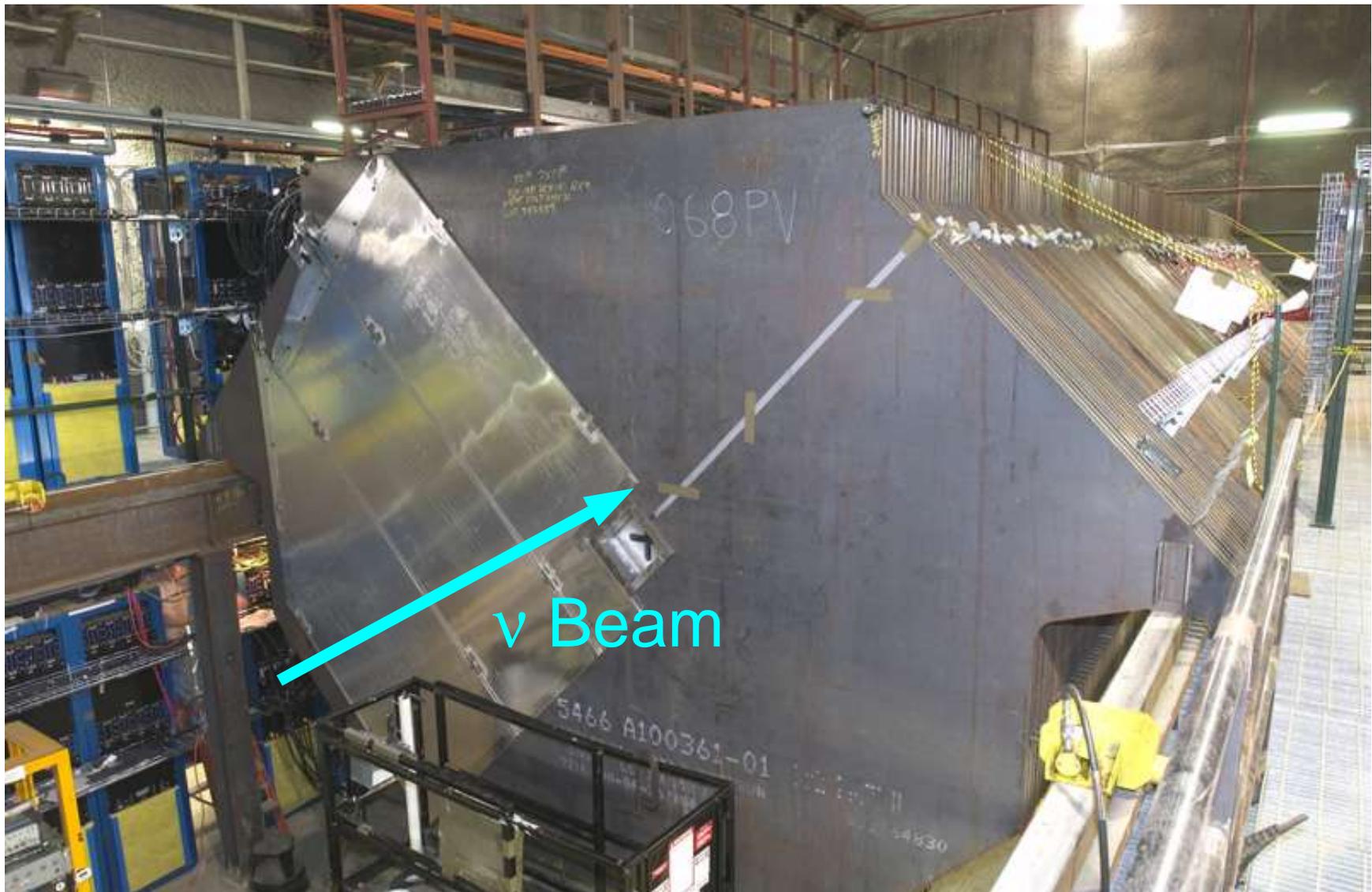
## NuMI Beamline

Neutrino energy can be adjusted by varying target/horn distances



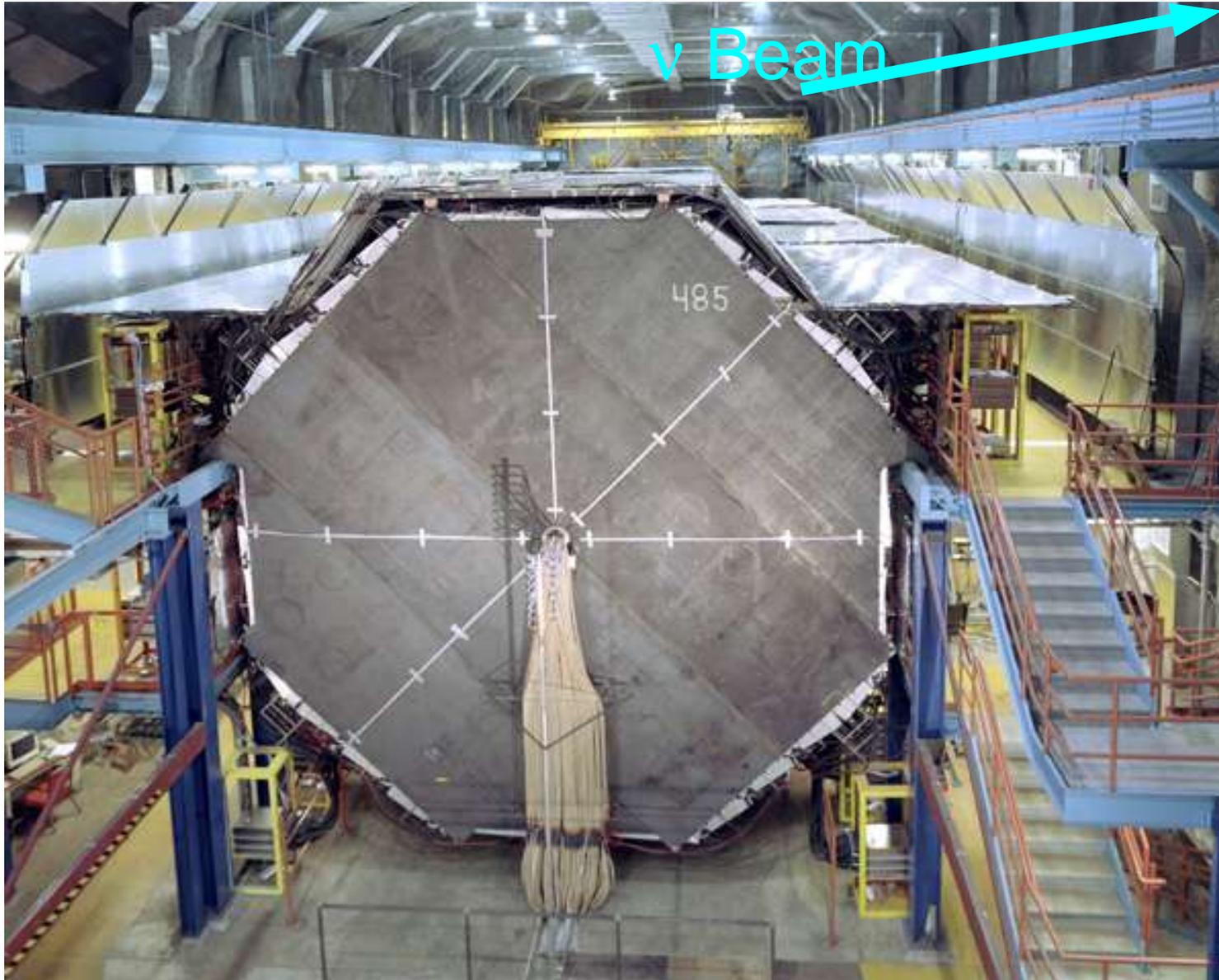
# MINOS Near Detector

- 0.98 kton steel & scintillator
- 1 km from target
- 292 steel planes (B=1.2 T)



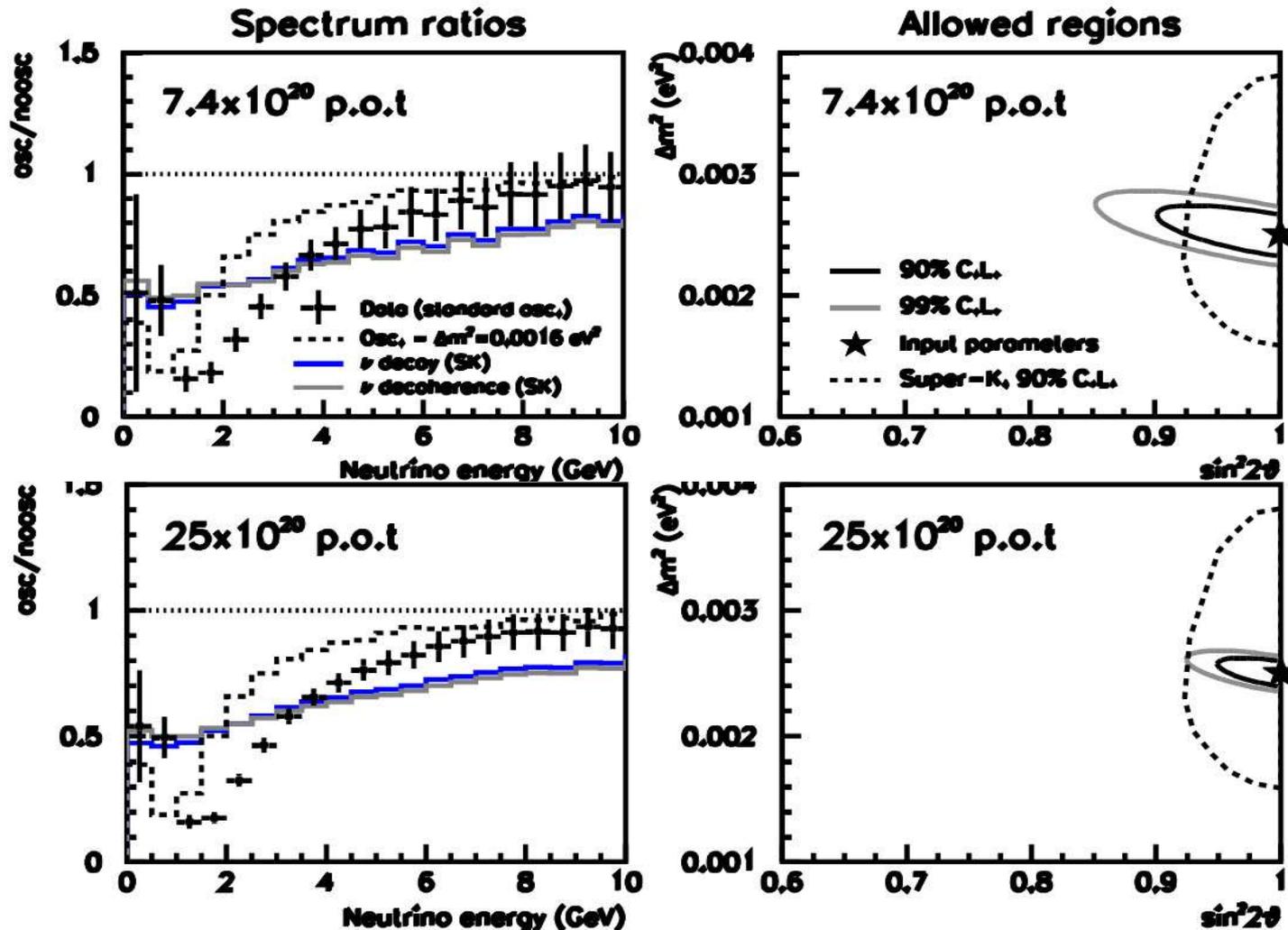
# MINOS Far Detector

- 5.4 kton steel & scintillator
- 735 km from target
- 486 steel planes (B=1.3 T)



# Muon Neutrino Disappearance

MC Prediction for  $\Delta m^2 = 0.0025 \text{ eV}^2$ ,  $\sin^2 2\theta = 1.0$



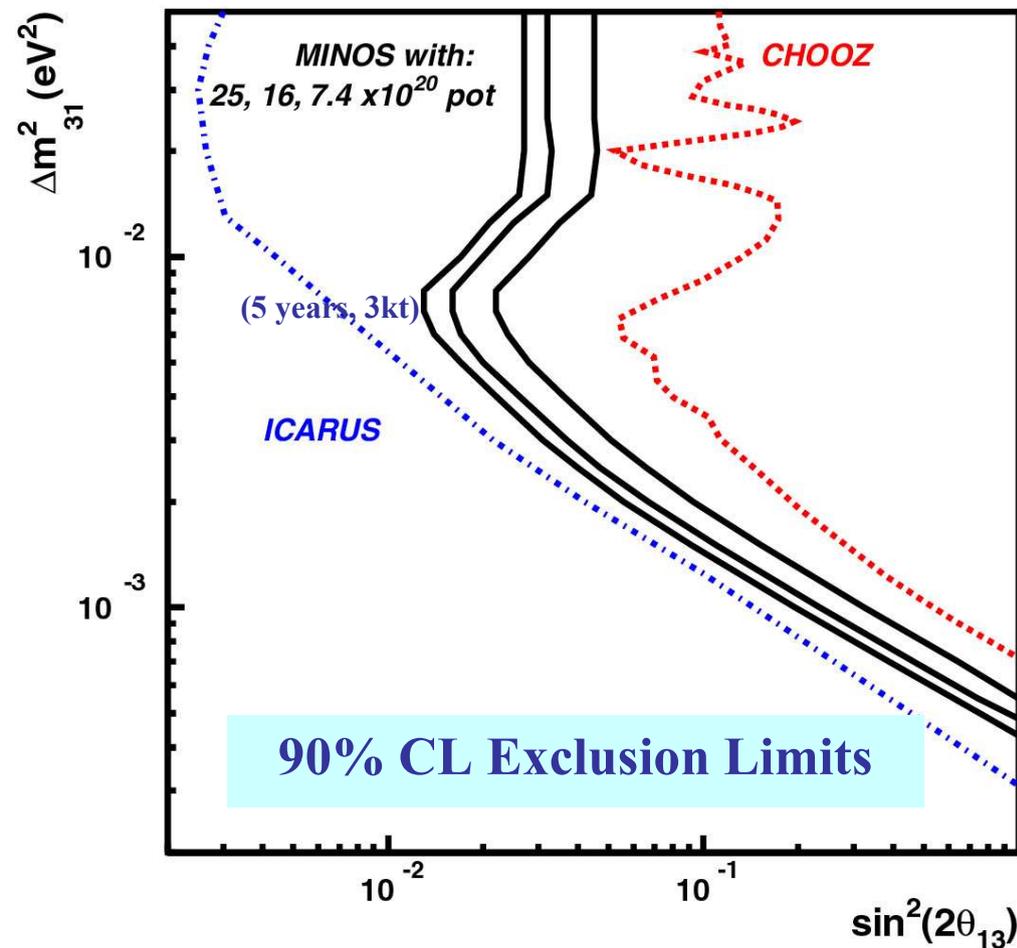
Oscillated/unoscillated ratio of number of  $\nu_\mu$  CC events in the far detector vs  $E_{obs}$

MINOS 90% and 99% CL allowed oscillation parameter space.

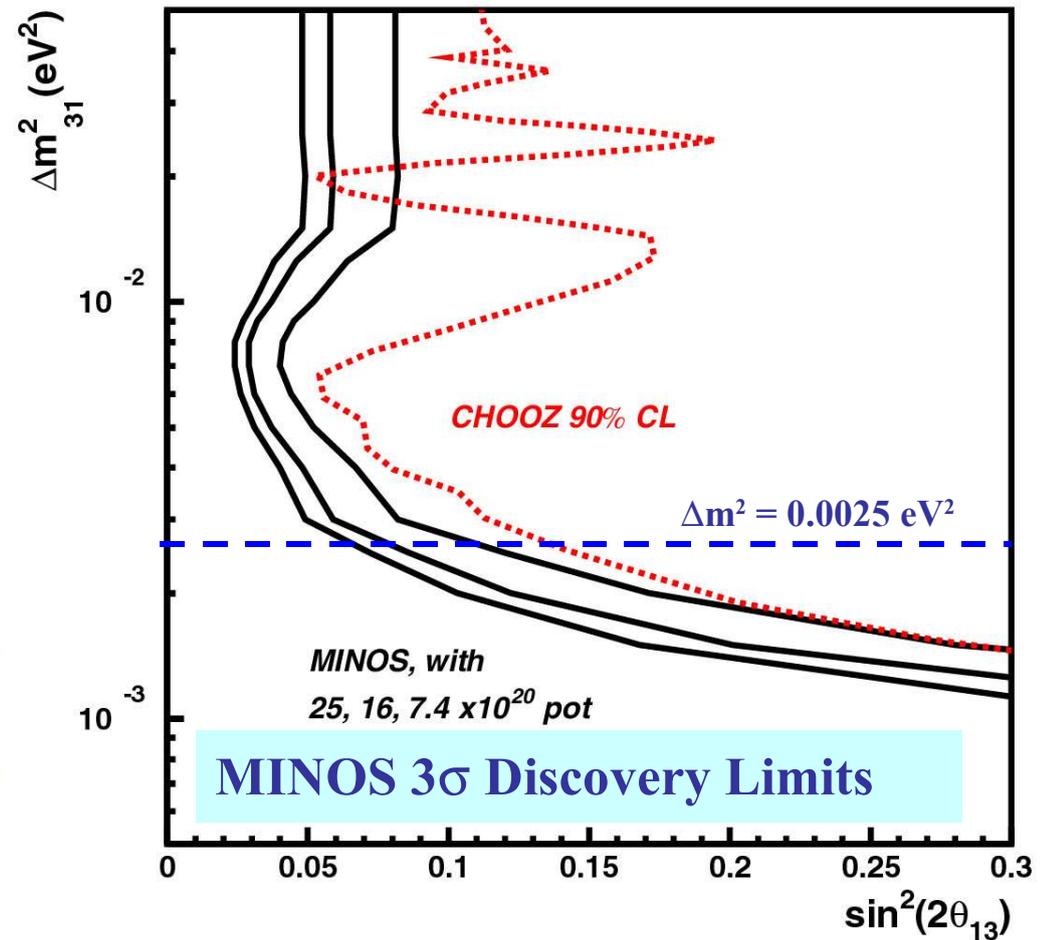
# Electron Neutrino Appearance

- MINOS sensitivities based on varying numbers of protons on target

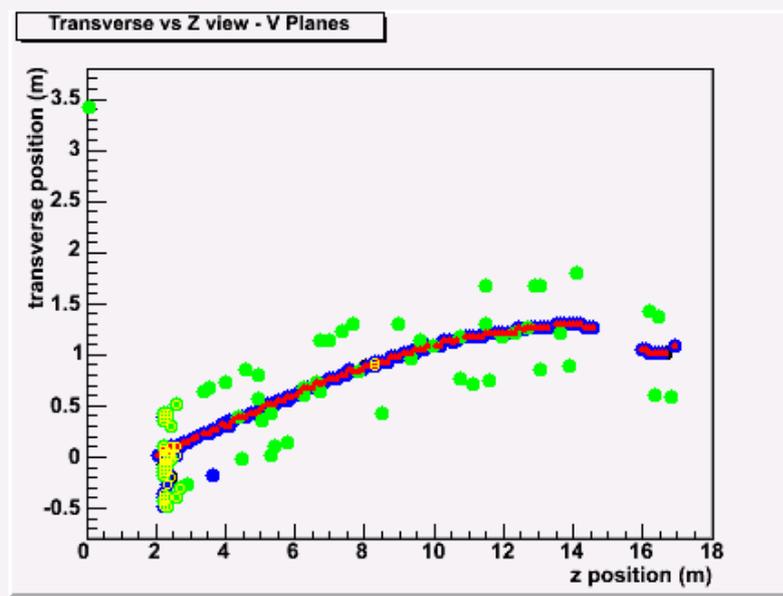
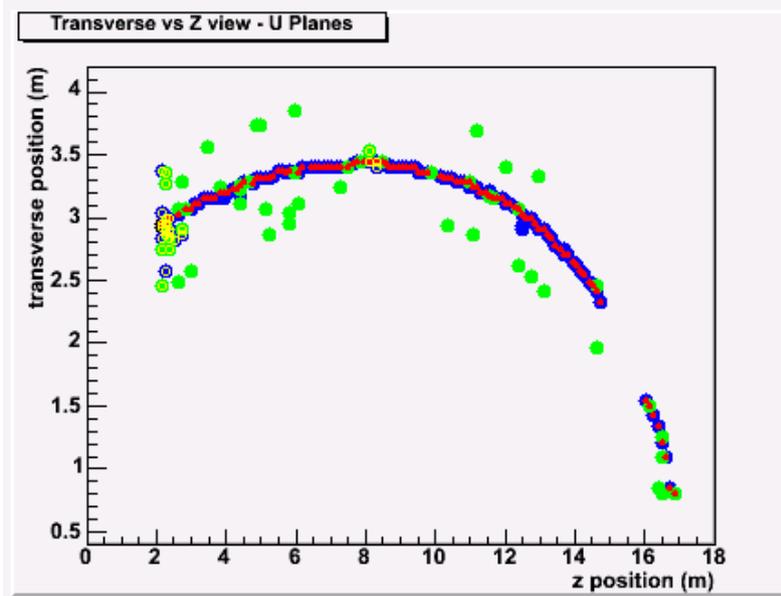
## 90% CL Exclusion



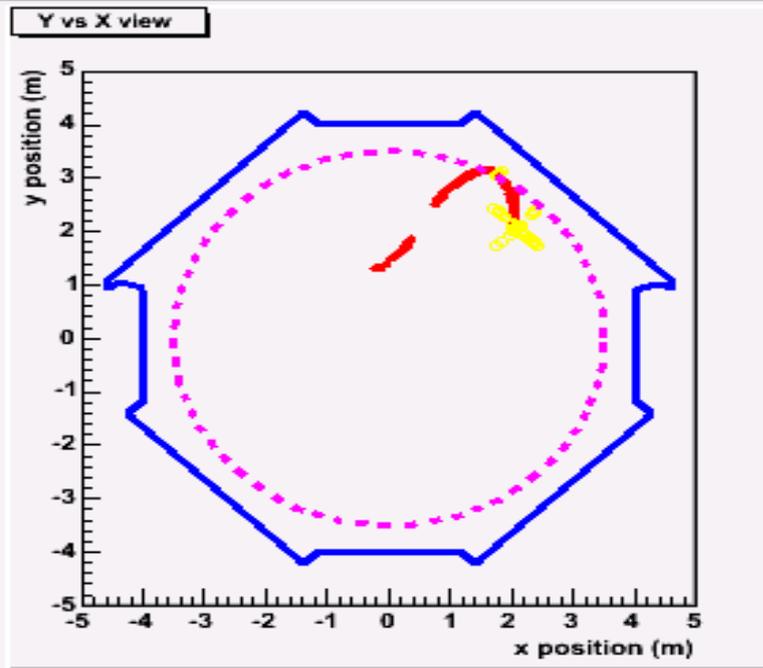
## 3 $\sigma$ Contours



# Far Detector: Numu CC Event

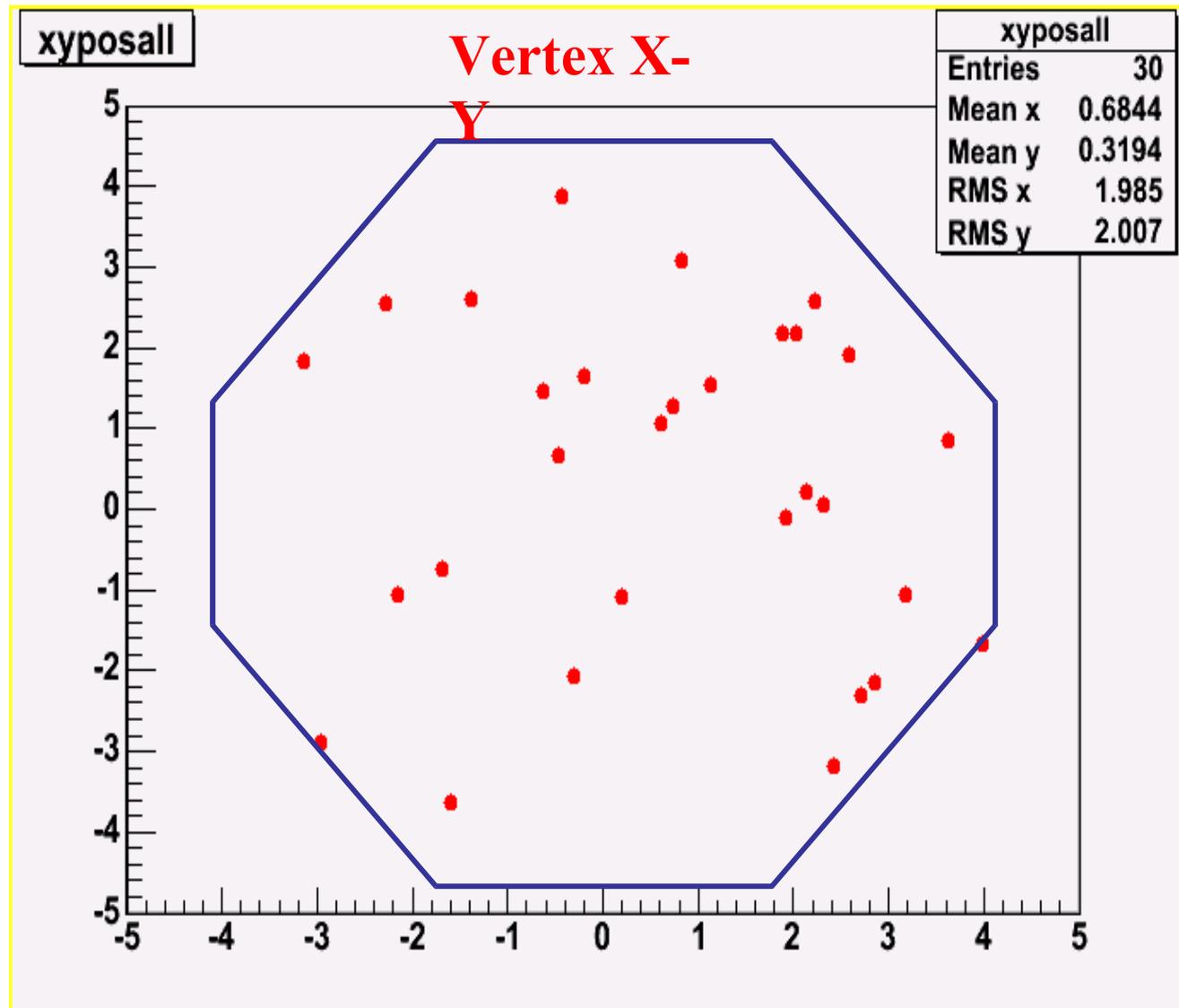


An Example:  
14.7 GeV Neutrino  
interaction (HE beam run)

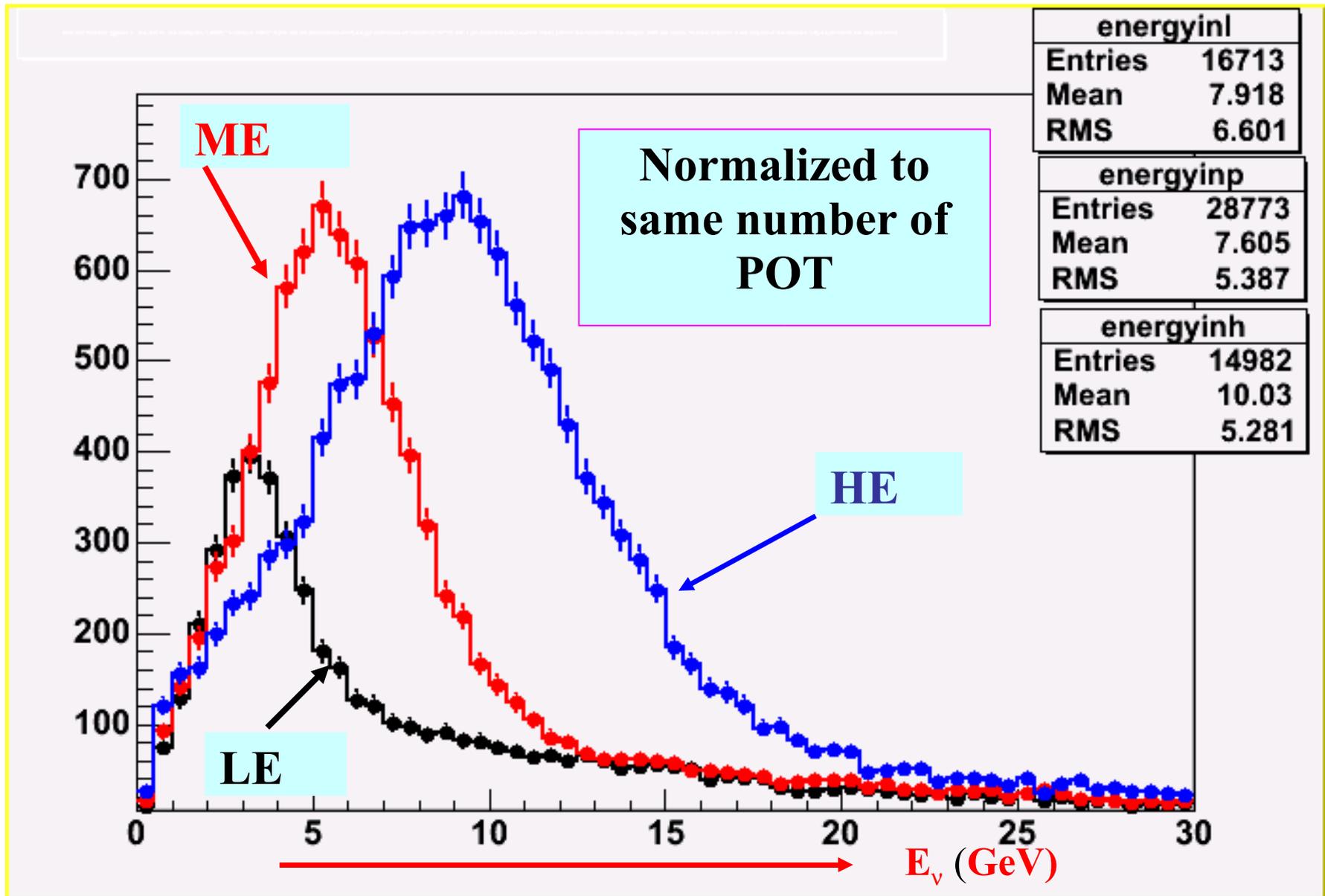


# Far Detector: $\nu$ -event Analysis

- Event characteristics in agreement with expectations
- Blind Analysis employed in the Far Detector Data

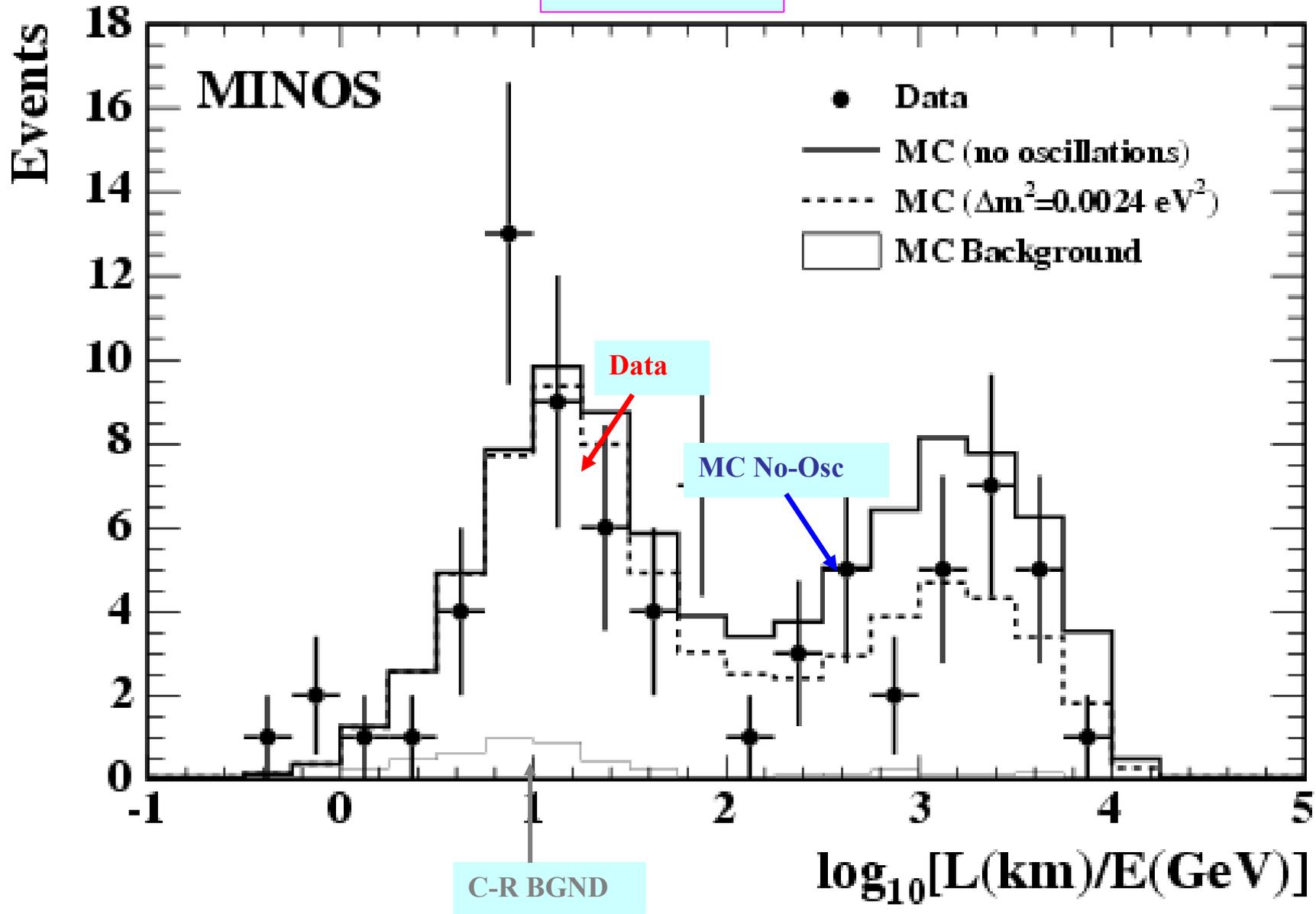


# Near Detector Data: Energy Scan

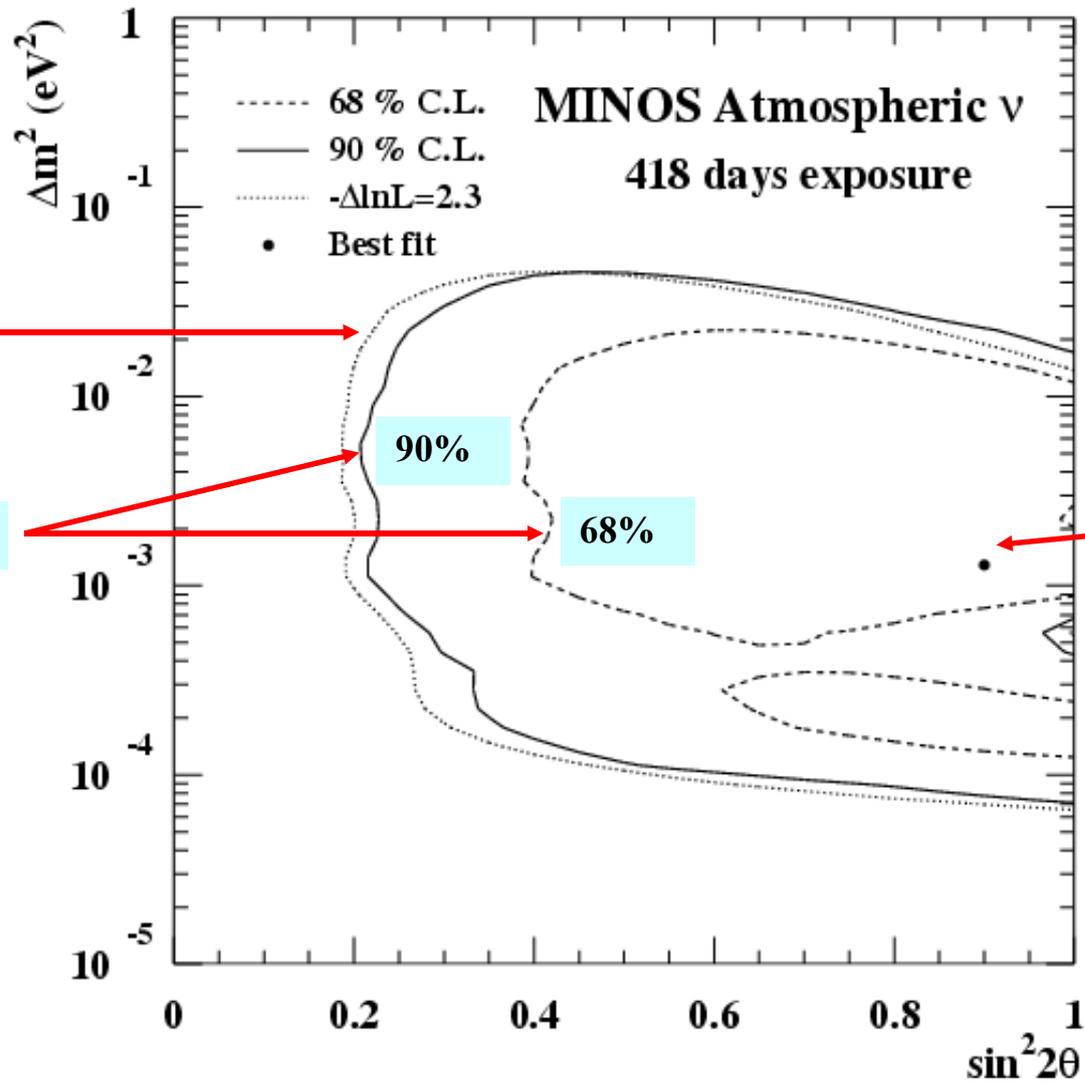


# Atmos FC + PC: Oscillation Analysis

L/E Plot



# MINOS ATMOS: Oscil Limits



$-\Delta\ln L = 2.3$   
90% CL

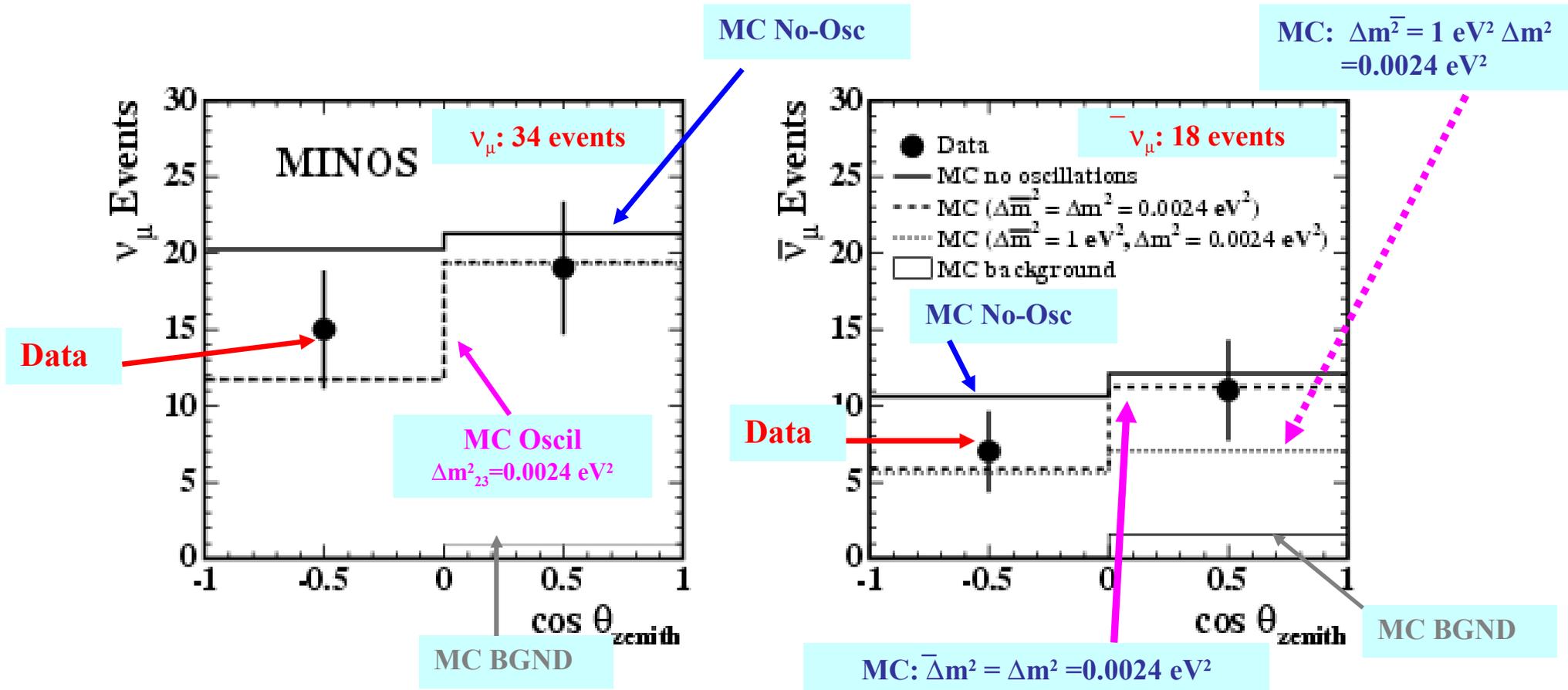
Feldman - Cousins

90%

68%

Best Fit  
 $\nu_\mu \leftrightarrow \nu_\tau$   
 $\sin^2 2\theta_{23} = 0.90$   
 $\Delta m^2_{23} = 0.0013 \text{ eV}^2$

# ATMOS: Charge Separated Up/down Distributions



**Data consistent with  $\nu_\mu$  and  $\bar{\nu}_\mu$  oscillating with same parameters.**

**CPT violating scenarios with large  $\Delta m^2_{23}$  not excluded with current data**

# MINOS Schedule

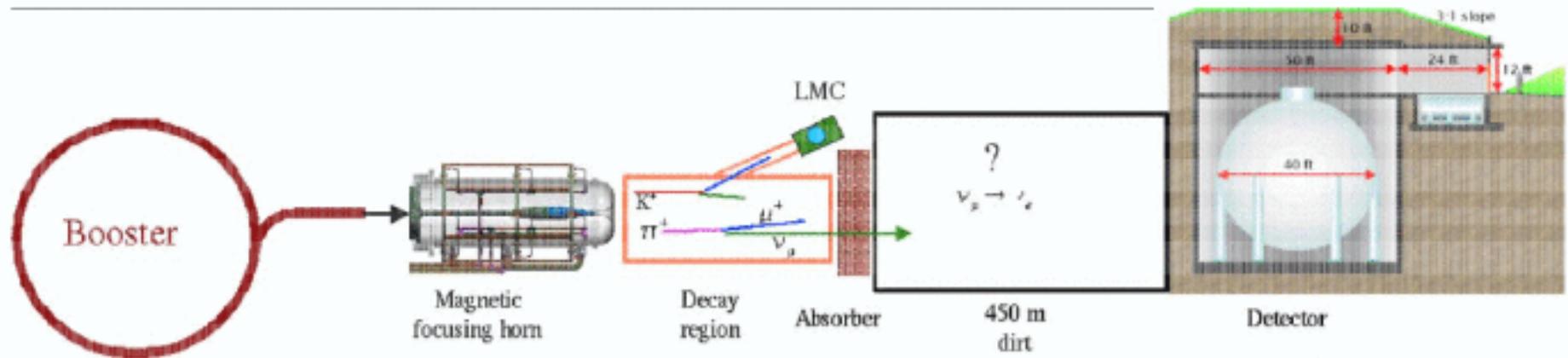
- **The MINOS Detectors and NuMI Beam construction and Commissioning have been successfully completed.**
- **Collecting Atmospheric Neutrino data since July 2003**
- **Collecting Accelerator Neutrino Data since March 2005**
- **First Results of FC and PC atmospheric neutrinos: Expect preprint in hep server soon.**
- **NuMI beam intensity is continuously improving, expect to have  $1 \times 10^{20}$  POT by the end of 2005.**
- **Both MINOS detectors operating satisfactorily**
- **Expect first physics results from NuMI beam neutrinos in 2006**

# MiniBooNE: A Definitive Test of the LSND Evidence for Oscillations: Search for $\nu_{\mu} \rightarrow \nu_e$



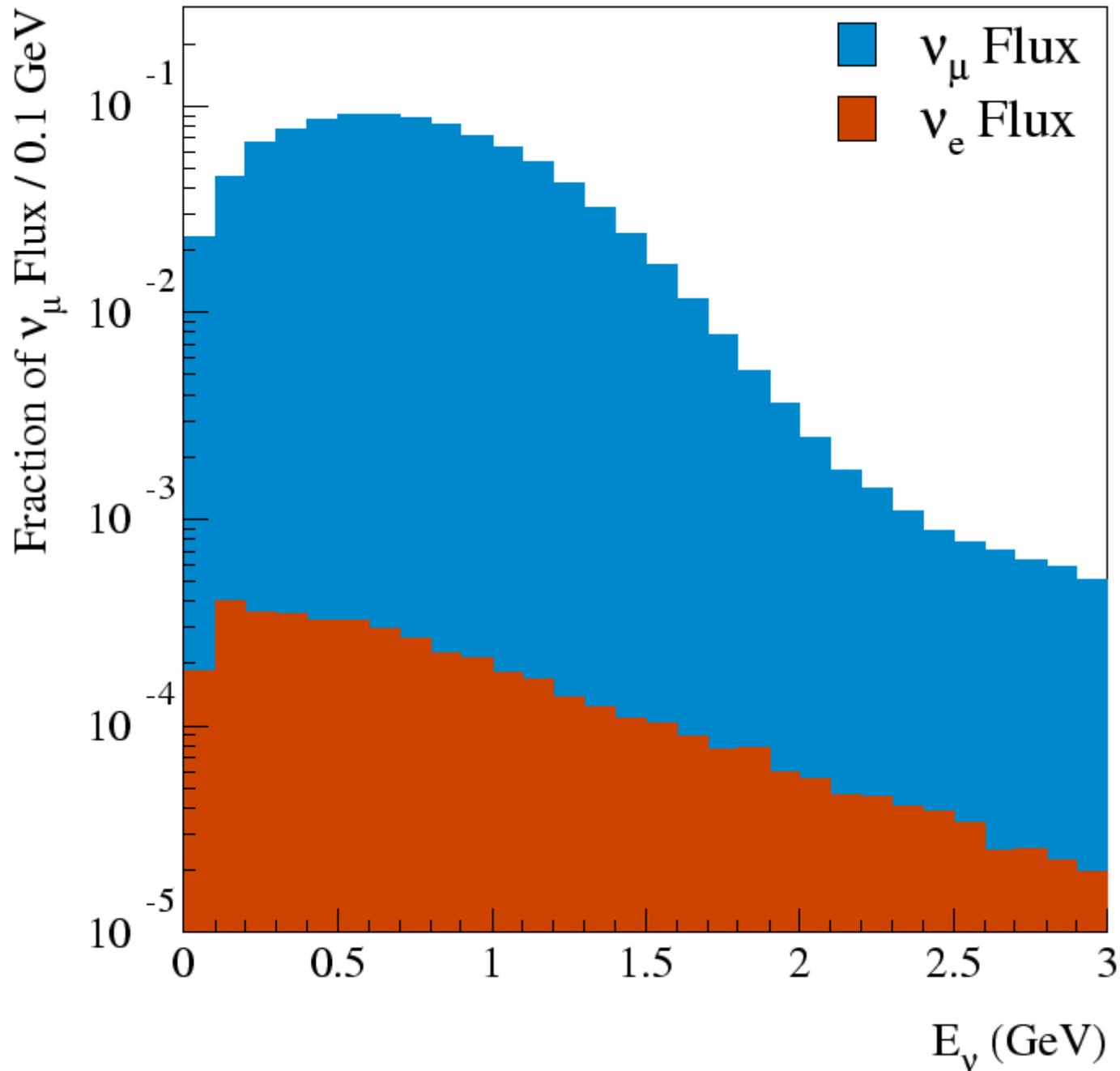
**Alabama, Bucknell, Cincinnati, Colorado, Columbia, Embry-Riddle,  
Fermilab, Indiana, Los Alamos, LSU, Michigan, Princeton, St. Mary's,  
Western Illinois, Yale**

# MiniBooNE - A Definitive Test of the LSND Evidence for $\nu$ Oscillations



- **Booster** - 8 GeV proton beam ( $5 \times 10^{20}$  POT/y)
- **Target** - 71 cm Be
- **Horn** - 5 Hz, 170 kA, 143  $\mu$ s, 2.5 kV,  $10^8$  pulses/y
- **Decay Pipe** - 50 m (adjustable to 25 m)
- **Neutrino Distance** -  $\sim 0.5$  km
- $\langle E_\nu \rangle \sim 1$  GeV
- $(\nu_e / \nu_\mu) \sim 5 \times 10^{-3}$
- **Detector** - 40' diameter spherical tank
- **Mass** - 800 (450) tons of mineral oil
- **PMTs** - 1280 detector + 240 veto, 8" diameter

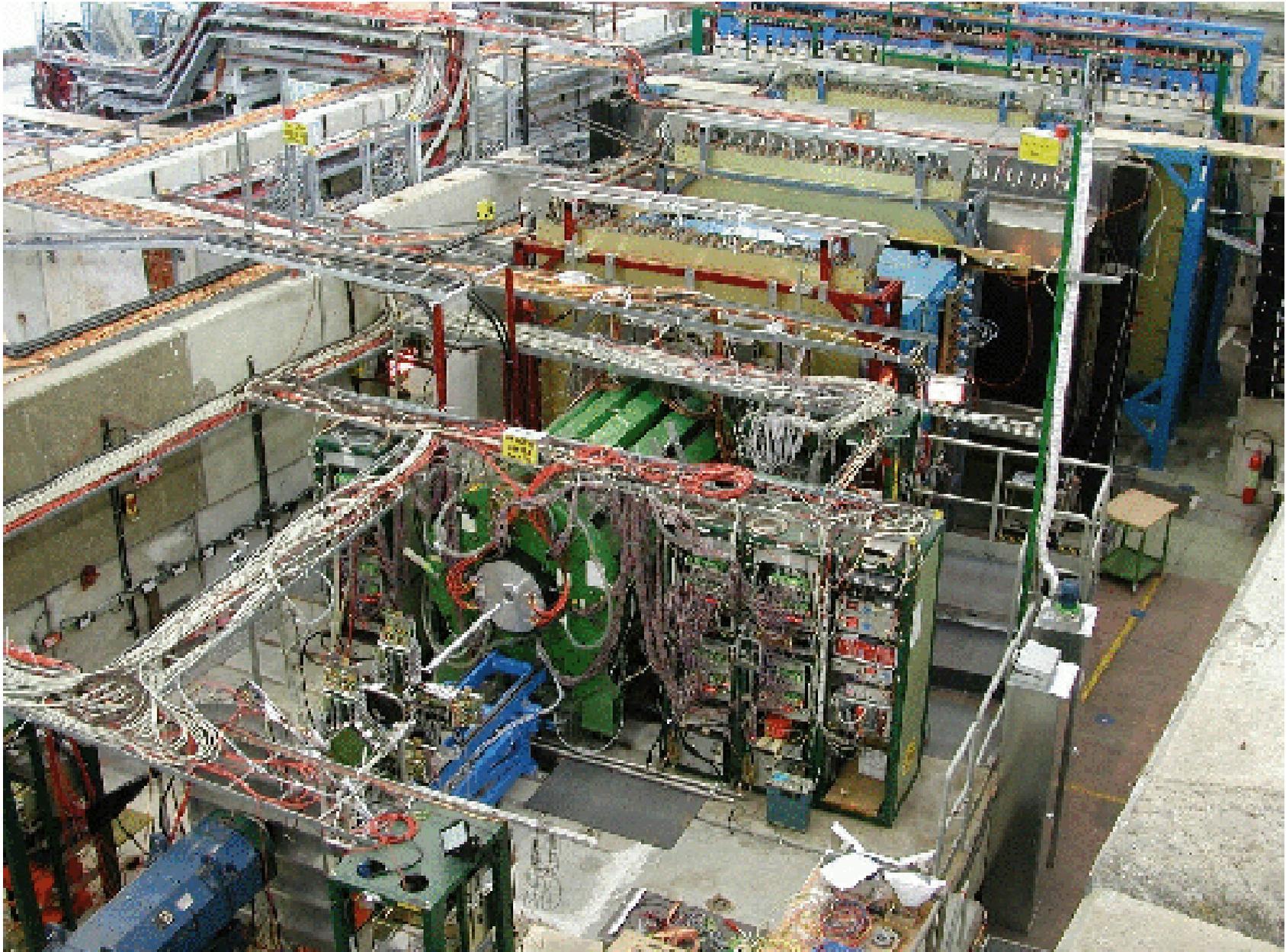
# MiniBooNE Neutrino Flux



**E910 at BNL**  
**HARP at CERN**

# HARP Experiment at CERN

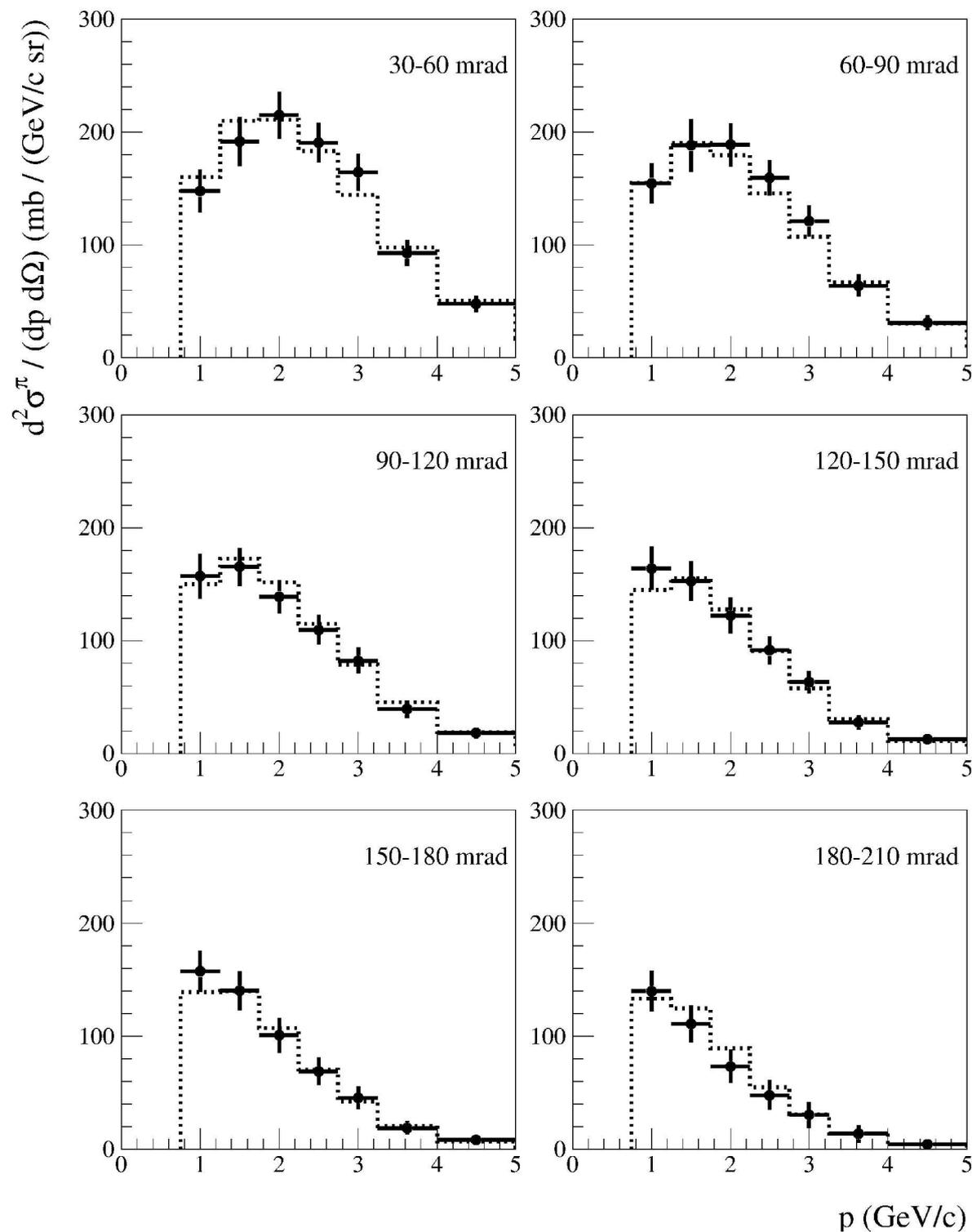
(parallel session talk by Borghi)



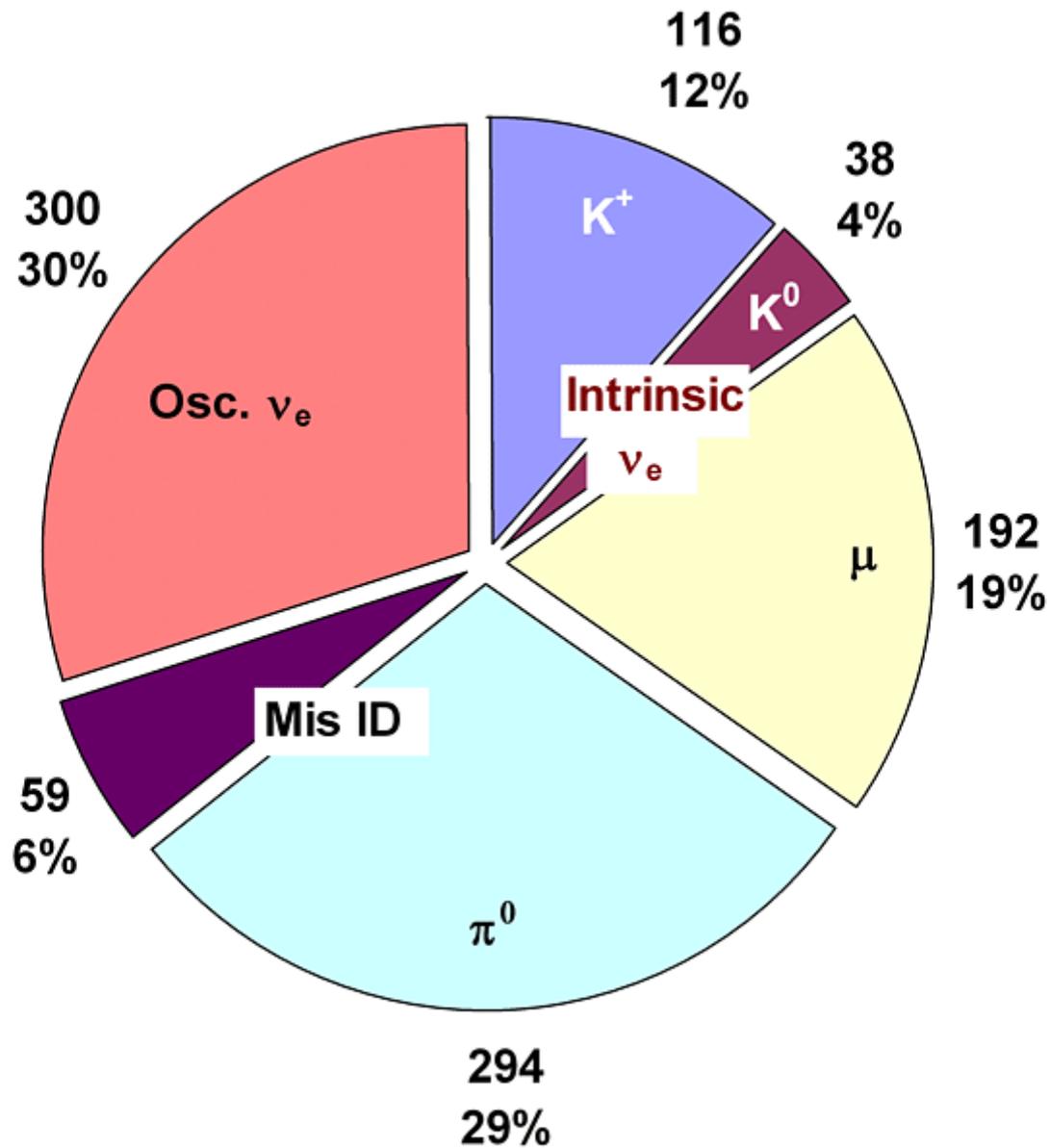
# Preliminary HARP results for 8.9 GeV/c Be data: Important for MiniBooNE

Fit to Sanford-Wang  
parameterization of production  
cross section (dotted line)

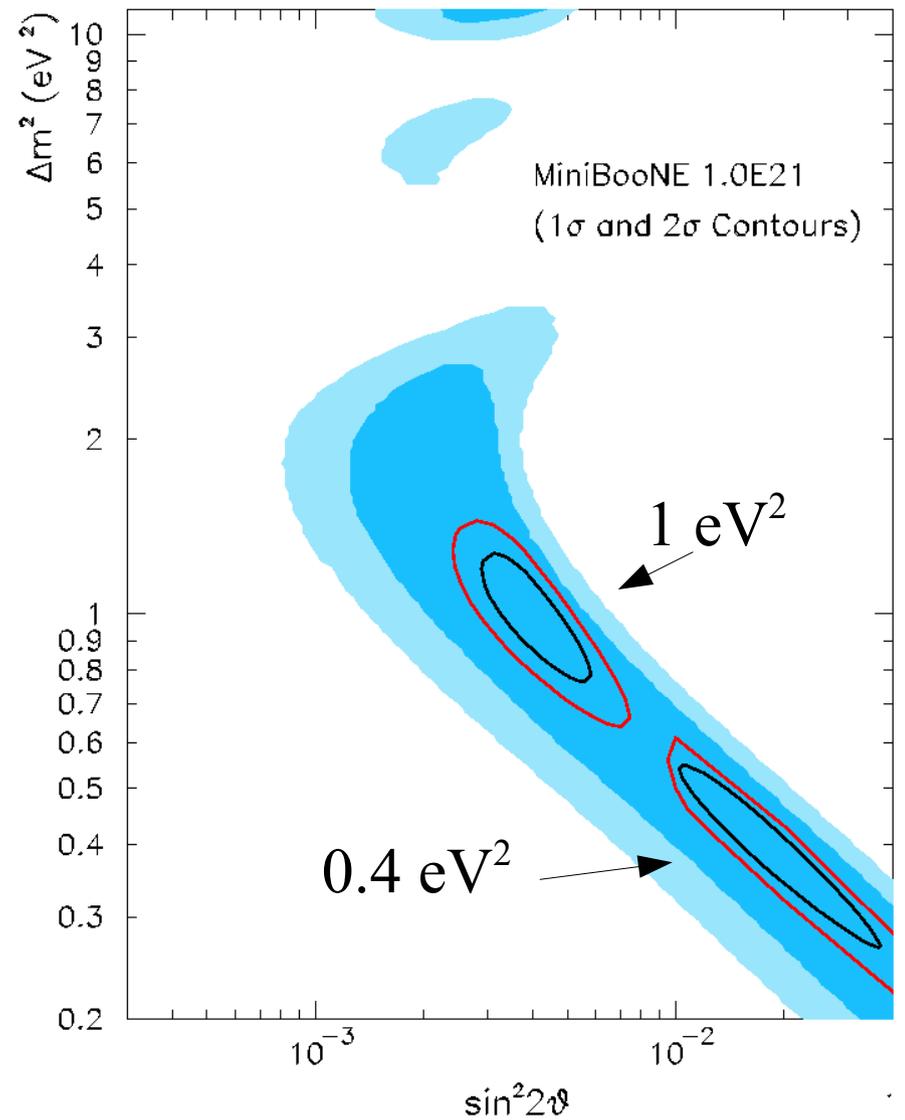
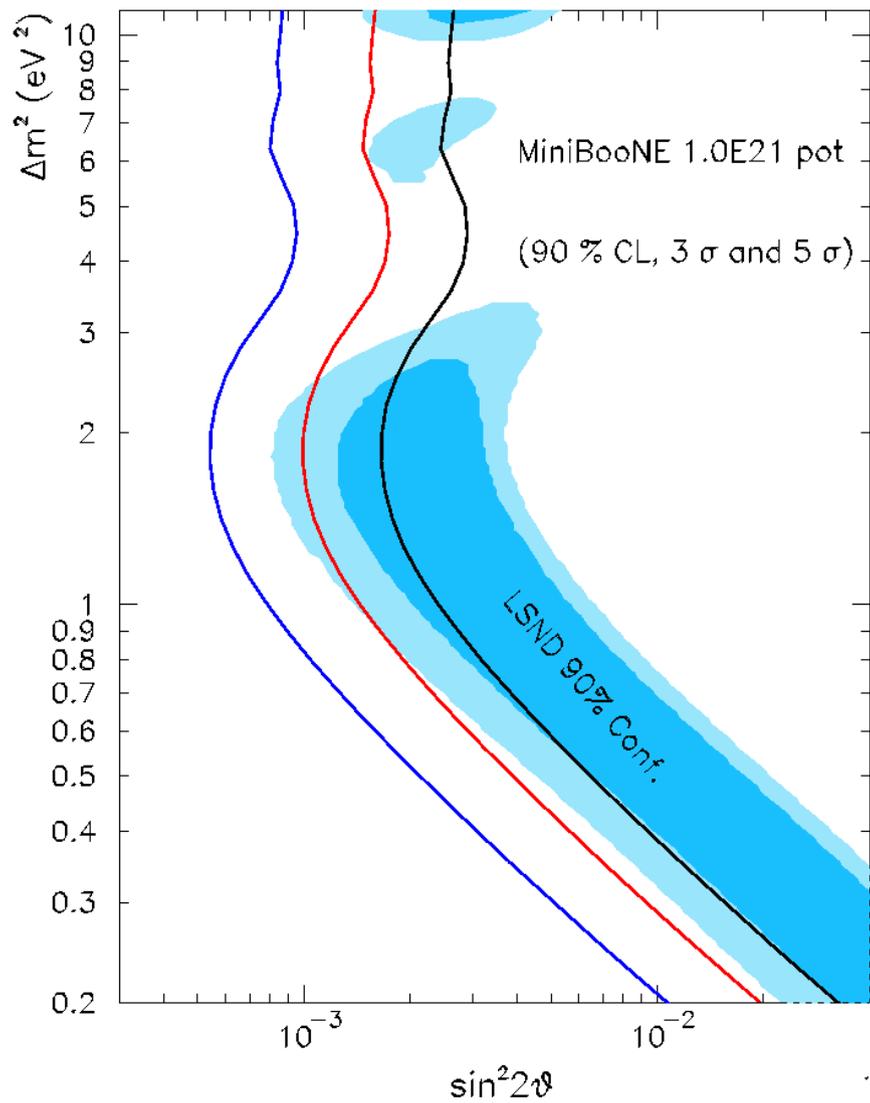
The preliminary p-Be  $\pi^+$  production  
cross sections are consistent at  
the 10-15% level with previous  
p-Be  $\pi^+$  production  
measurements



# NUANCE Prediction After PID Selection



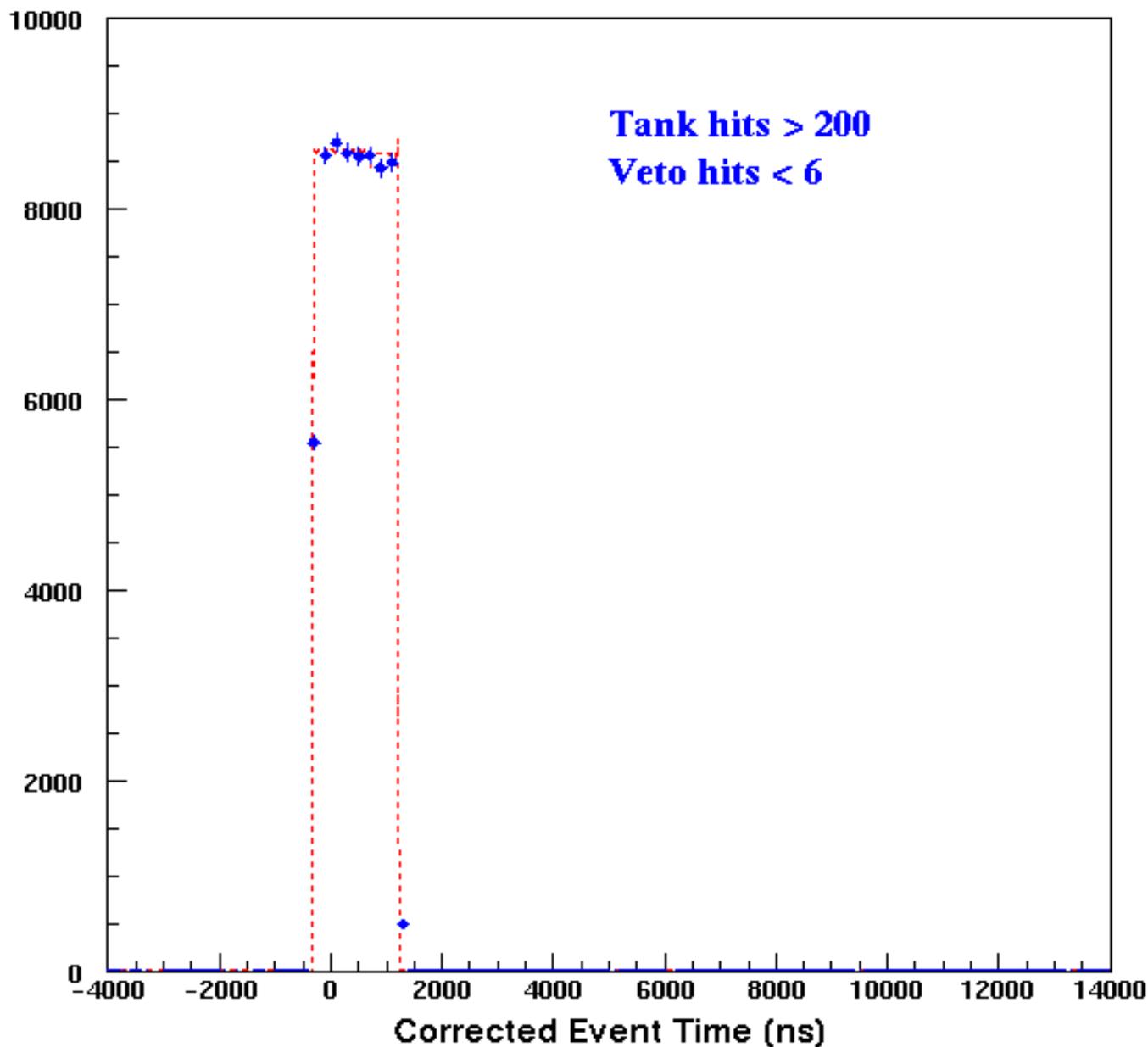
# Expected MiniBooNE Sensitivity



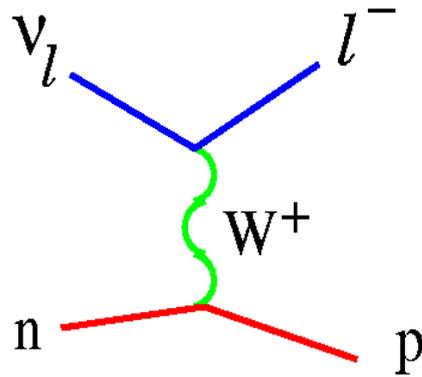
# MiniBooNE Neutrino Events Are Very Clean!

Neutrino Signal to Cosmic-Ray Background  $\sim 5000$  to 1!

Have collected a total of  $\sim 700\text{K}$  neutrino events from  $\sim 6.5\text{E}20$  POT



# I. CCQE Events $\nu_{\mu}^{12}\text{C} \rightarrow \mu^{-} \text{p}^{11}\text{C}^*$



About 48% of the MiniBooNE events are **CCQE**

MiniBooNE now has one of the largest sample of **CCQE** events!

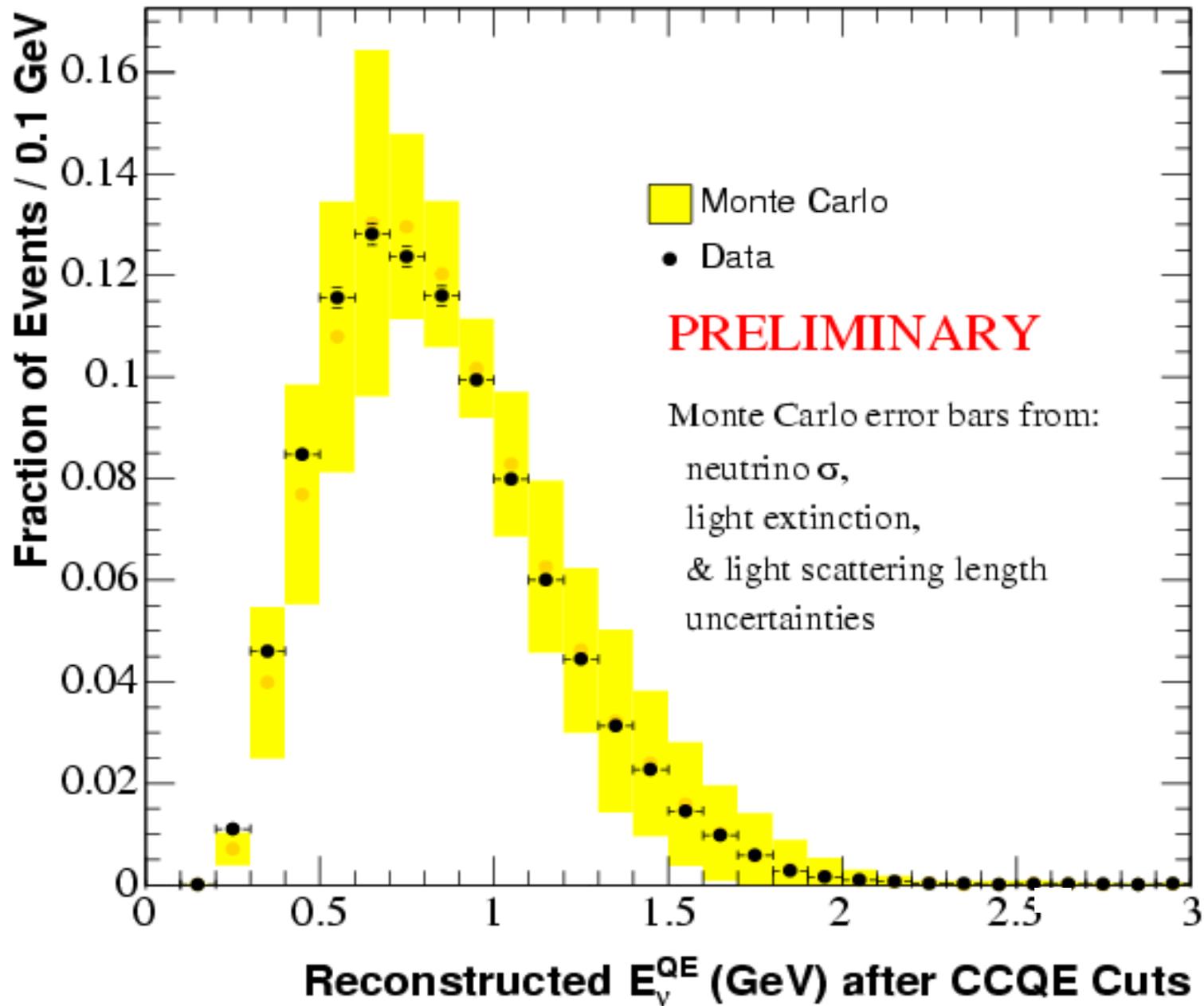
Important for  $\nu_{\mu}$  Disappearance &  $\nu_e$  Appearance Background

Serves as the “Golden Mode” for Measuring Cross Sections

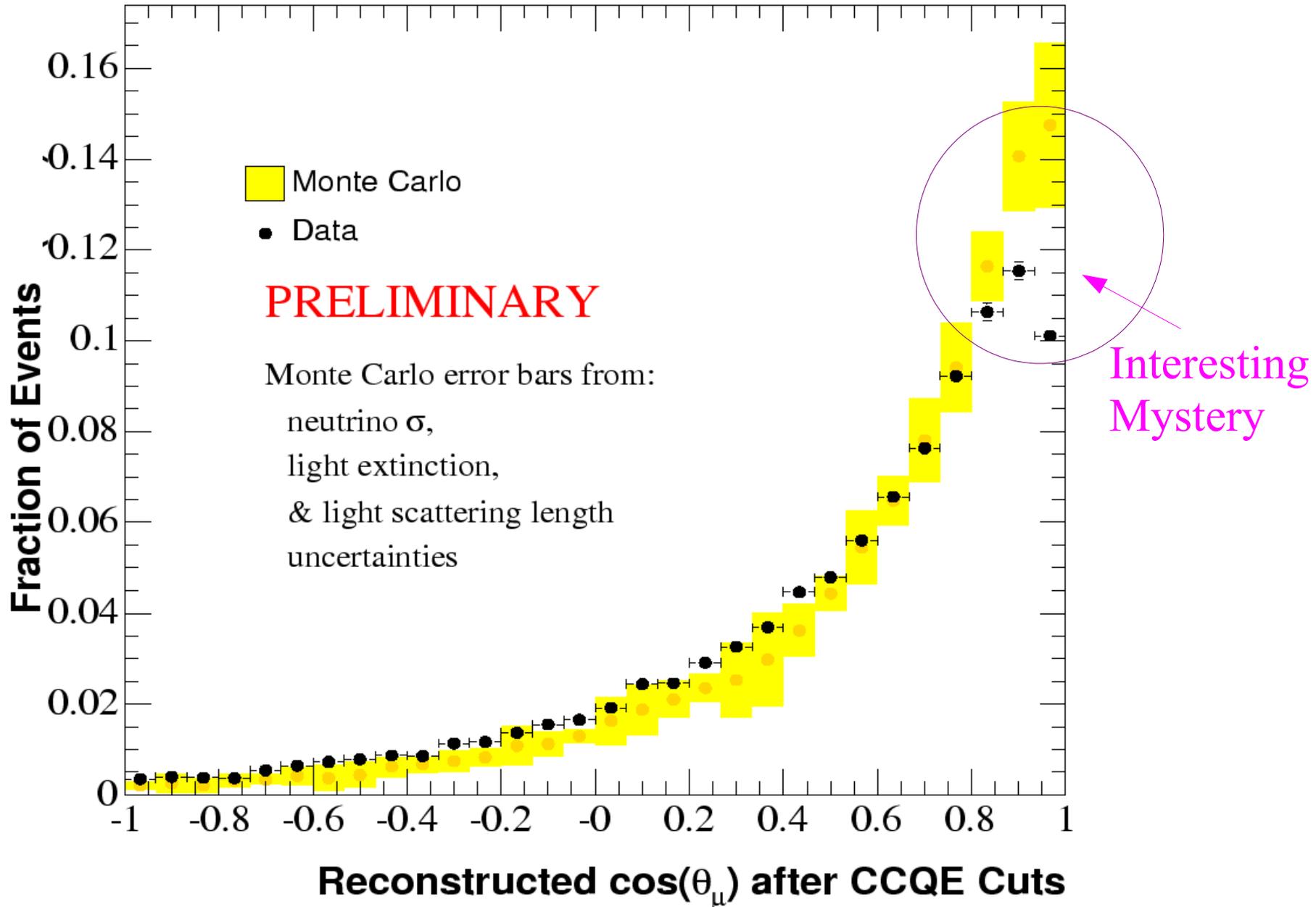
Selection: Single ring event, consistent with a muon

# CCQE Energy Distribution

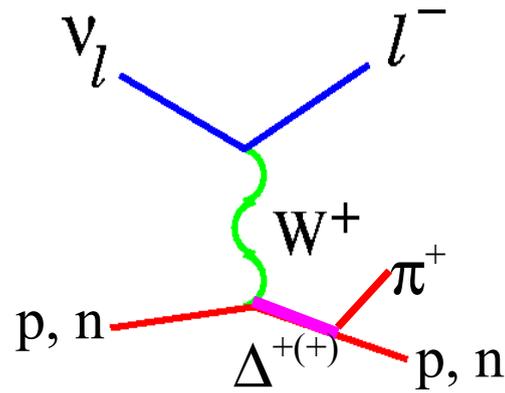
~10% energy resolution



# CCQE Angular Distribution



## II. CC $\pi^+$ Events $\nu_{\mu} {}^{12}\text{C} \rightarrow \mu^{-} \pi^{+} \text{X}$



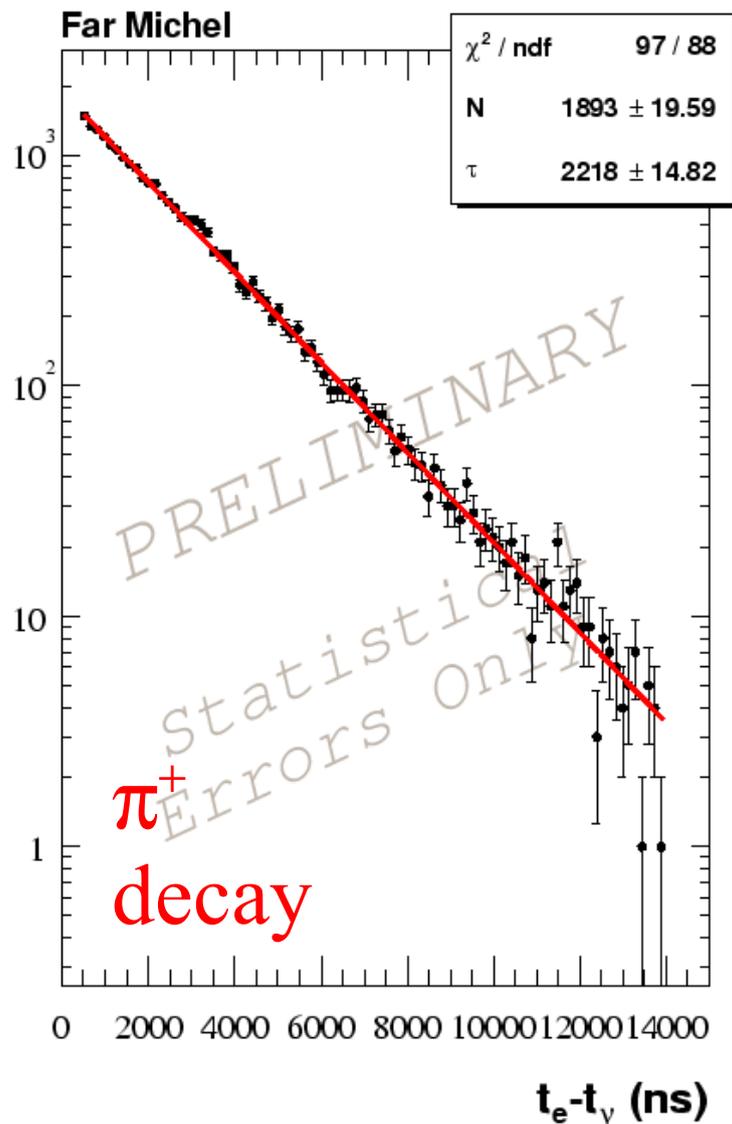
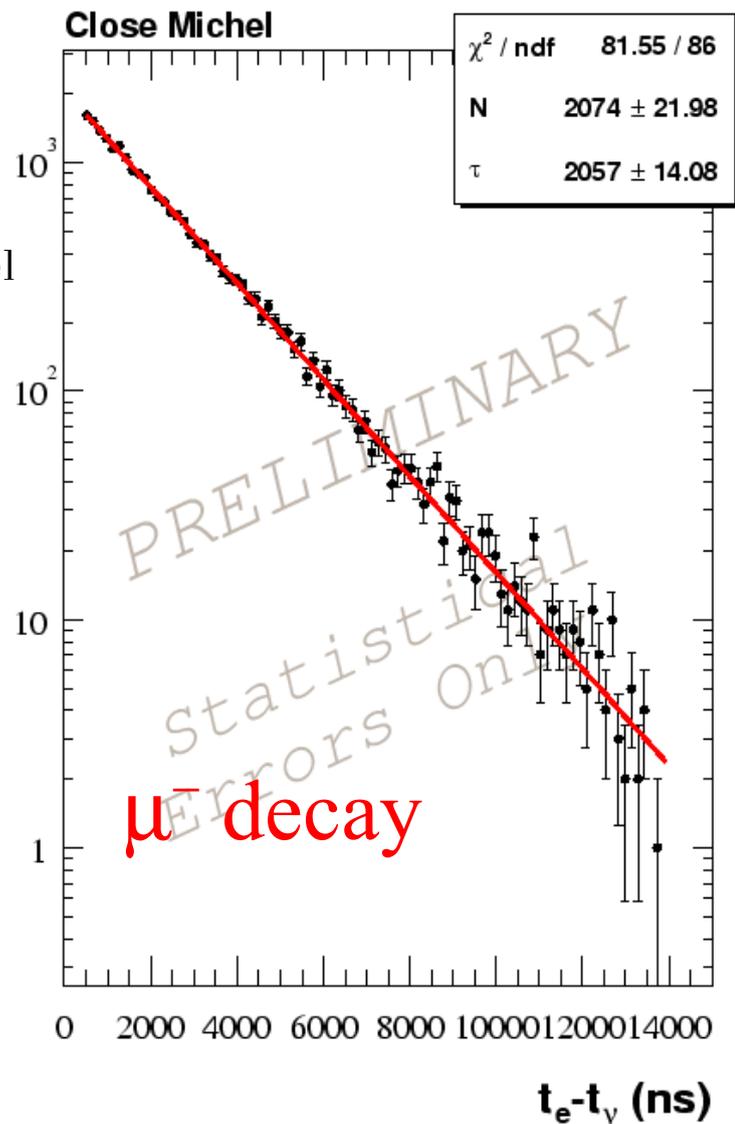
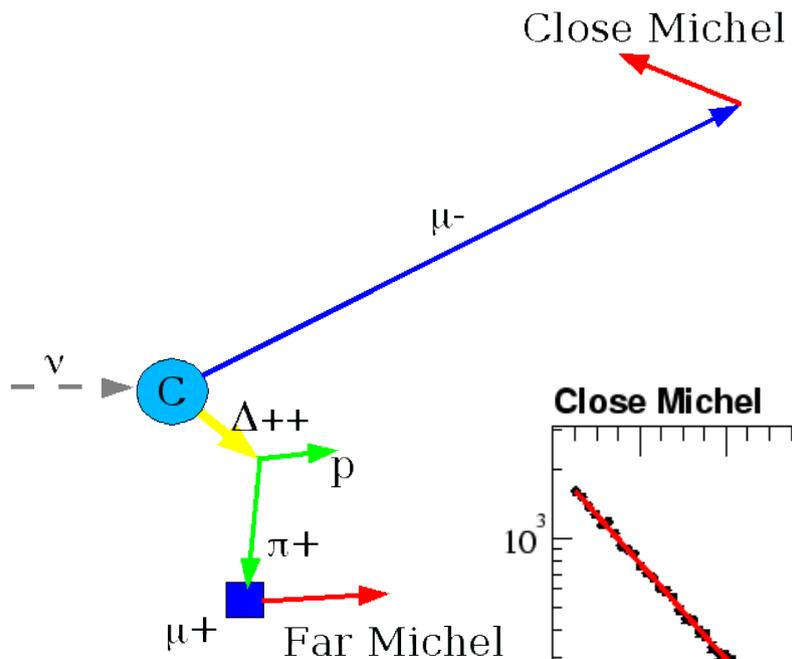
Not Well Measured at Low Energies

About 31% of the MiniBooNE events are **CC  $\pi^{+}$**

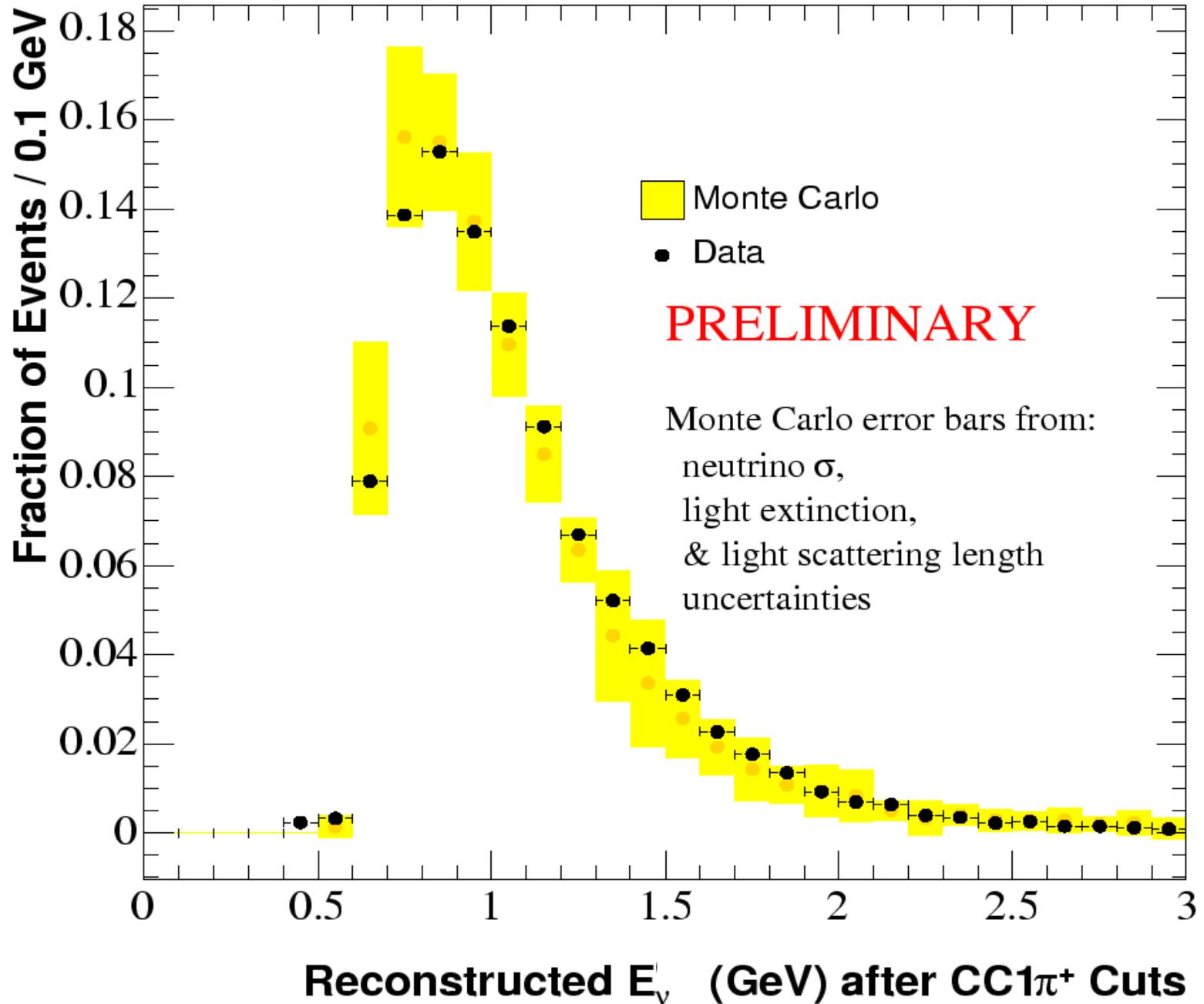
MiniBooNE now has the world's largest sample of **CC  $\pi^{+}$**  events!

Important for  $\nu_{\mu}$  Disappearance &  $\nu_{e}$  Appearance Background

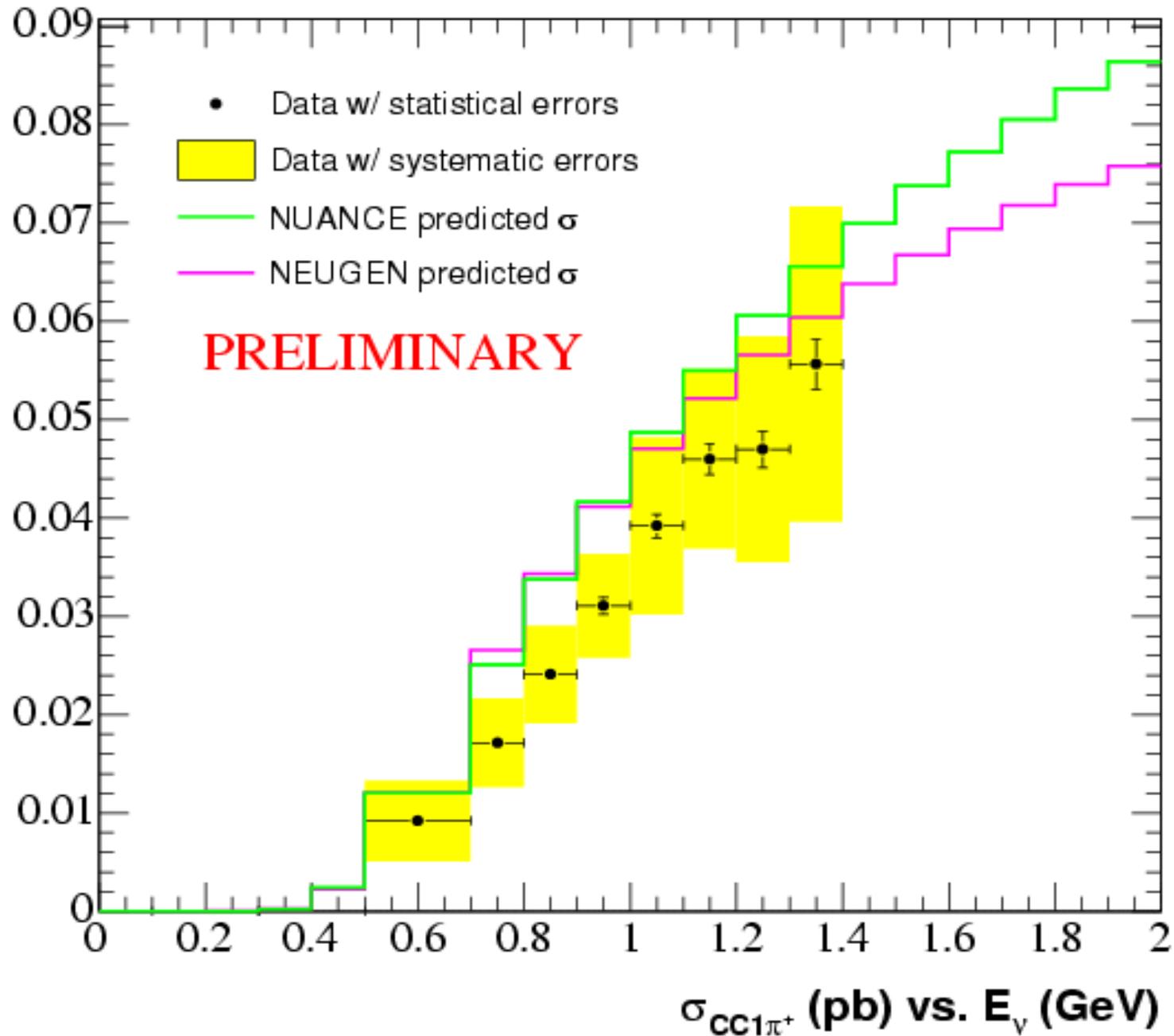
Selection: Double-ring event with 2 Michel-electrons



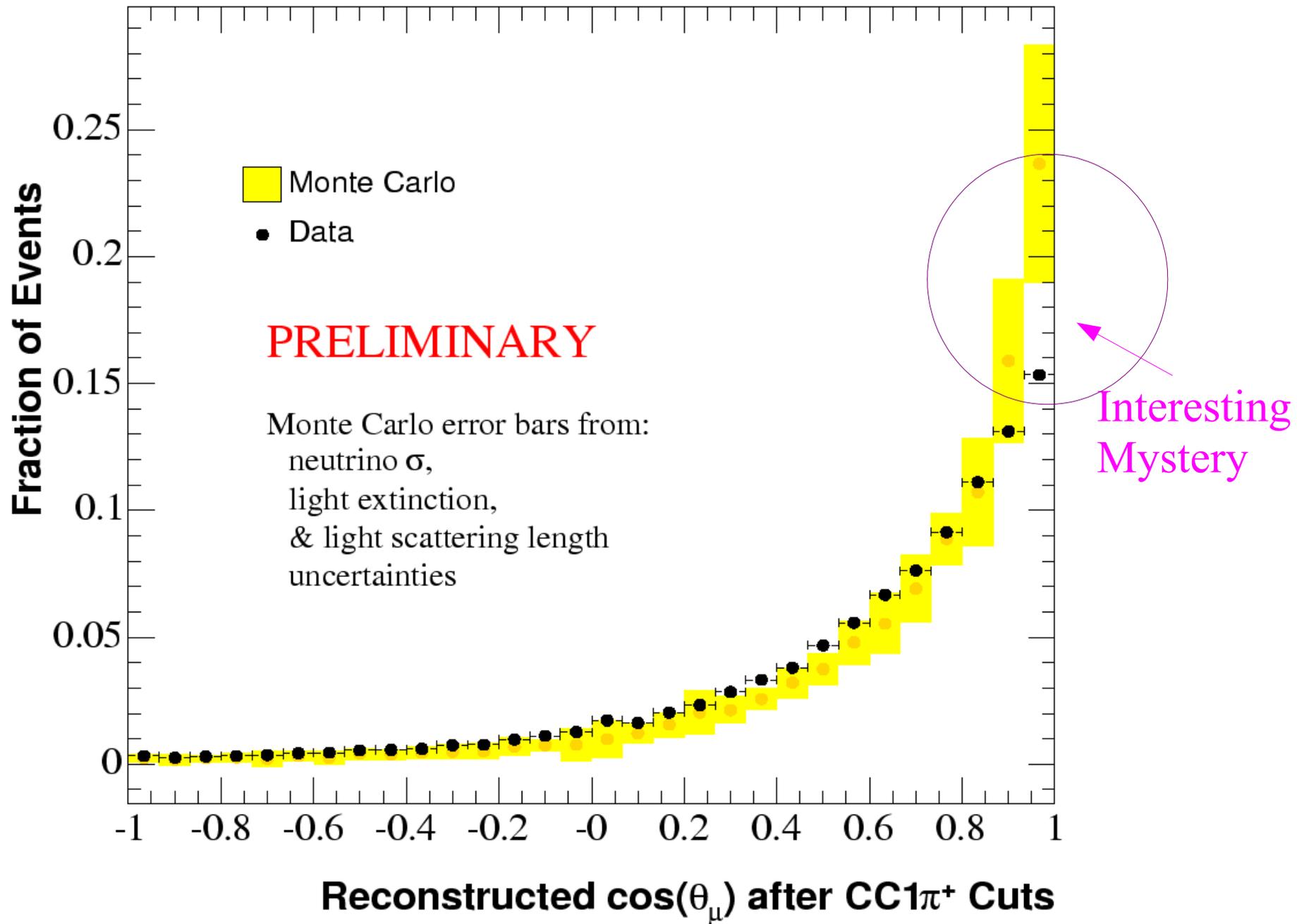
# CC $\pi^+$ Energy Distribution



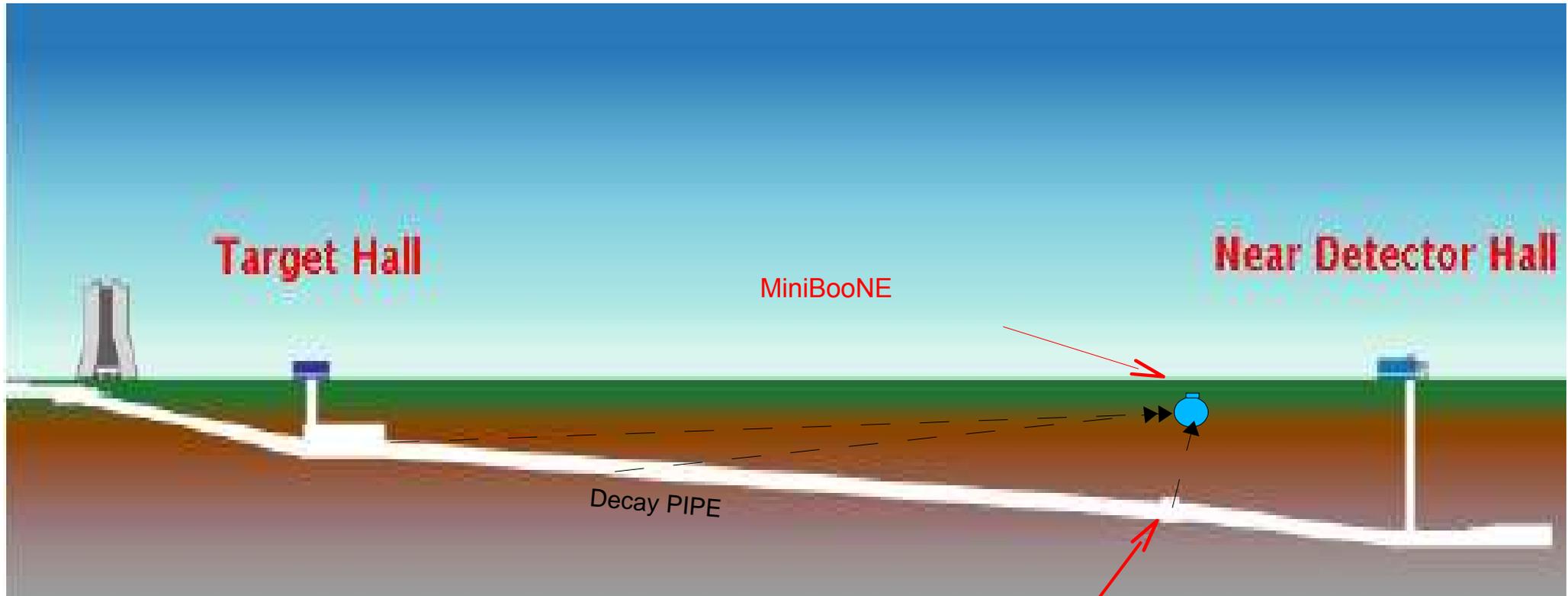
# CC $\pi^+$ Cross Section (from CC $\pi^+$ /CCQE Ratio)



# CC $\pi^+$ Angular Distribution



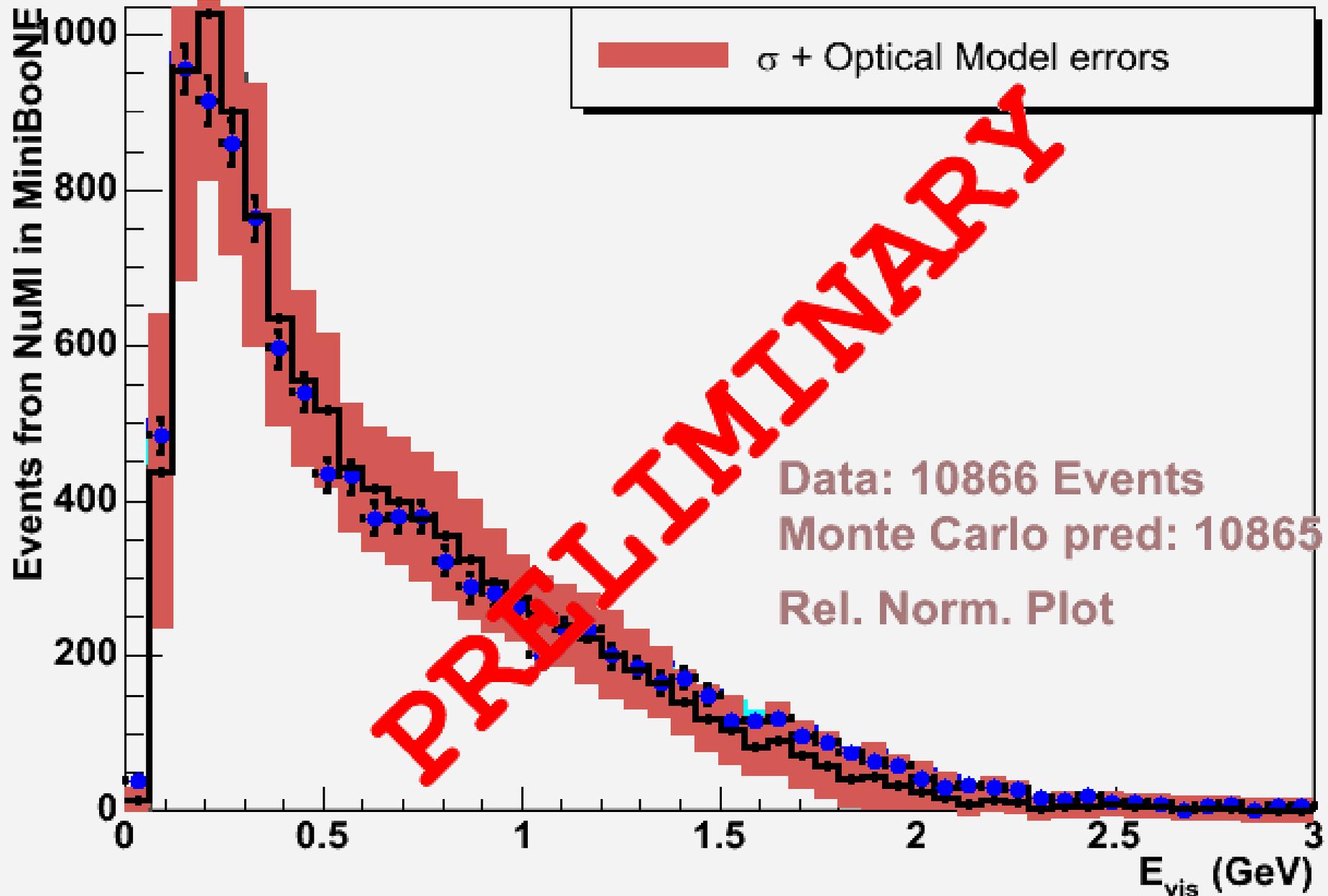
# MiniBooNE Observes Events from both the 8-GeV Beamline & from NuMI



NuMI beam line

Absorber (Beam DUMP)

# NuMI/MiniBooNE Data/MC Comparison

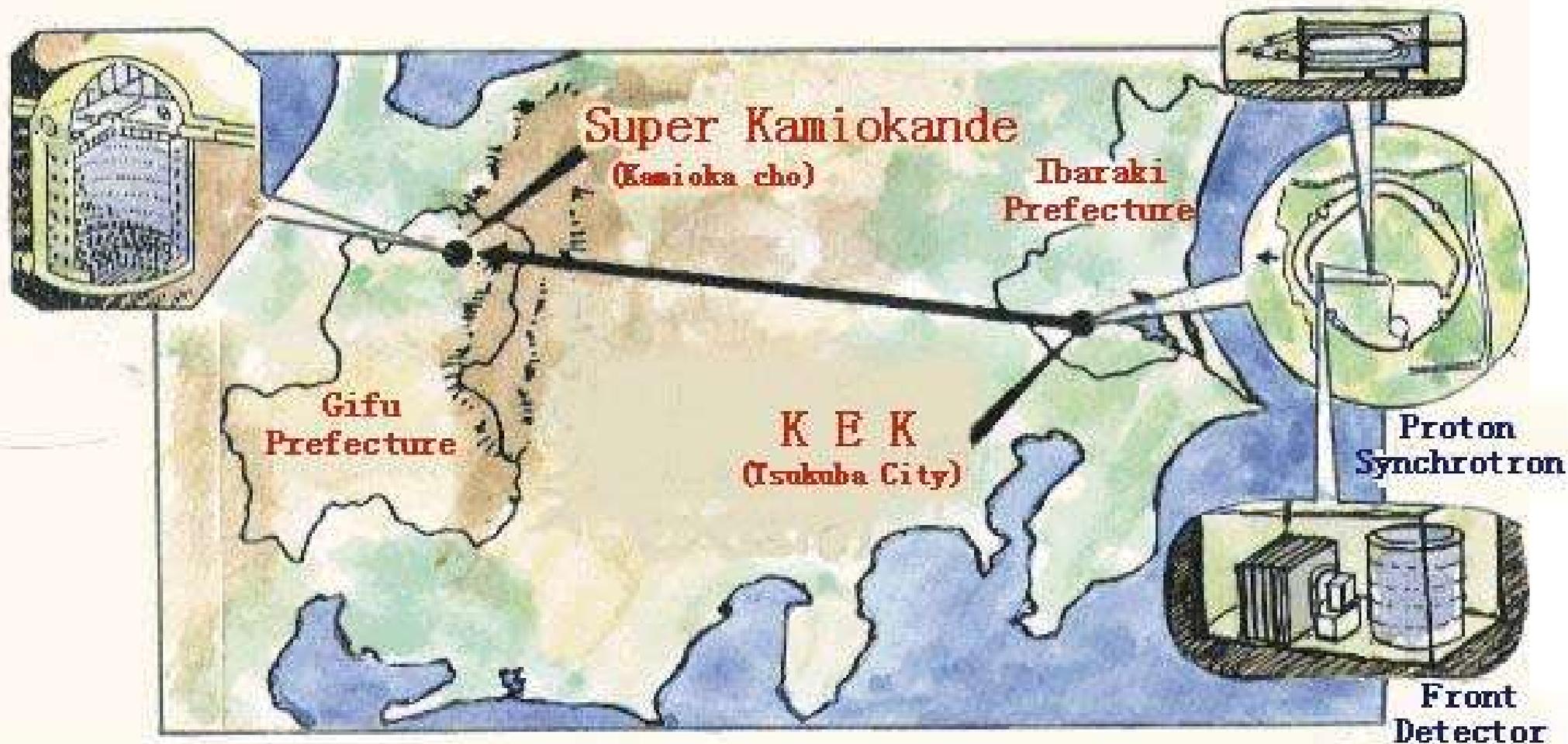


# MiniBooNE Schedule

- Publish CCQE,  $CC\pi^+$ ,  $NC\pi^0$  Cross Section Results
- Complete “Blind” Oscillation Analysis (Good agreement between data & MC) & Present Oscillation Results
- Run Antineutrinos (and/or Neutrinos) in 2006
- If MiniBooNE Confirms LSND  
=> **Physics Beyond the Standard Model!**

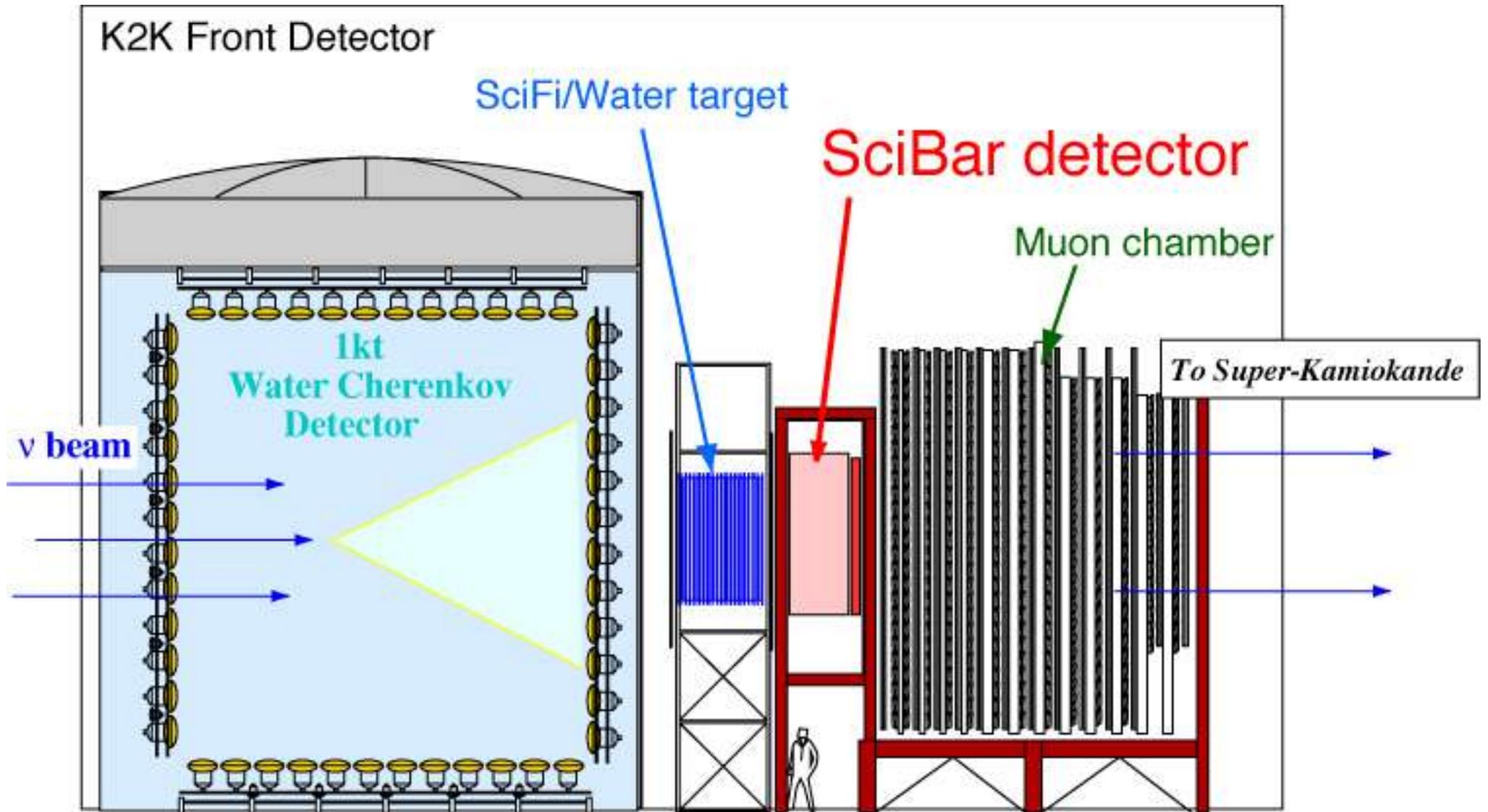
# K2K Experiment

Atmospheric  $\nu$  Oscillations!



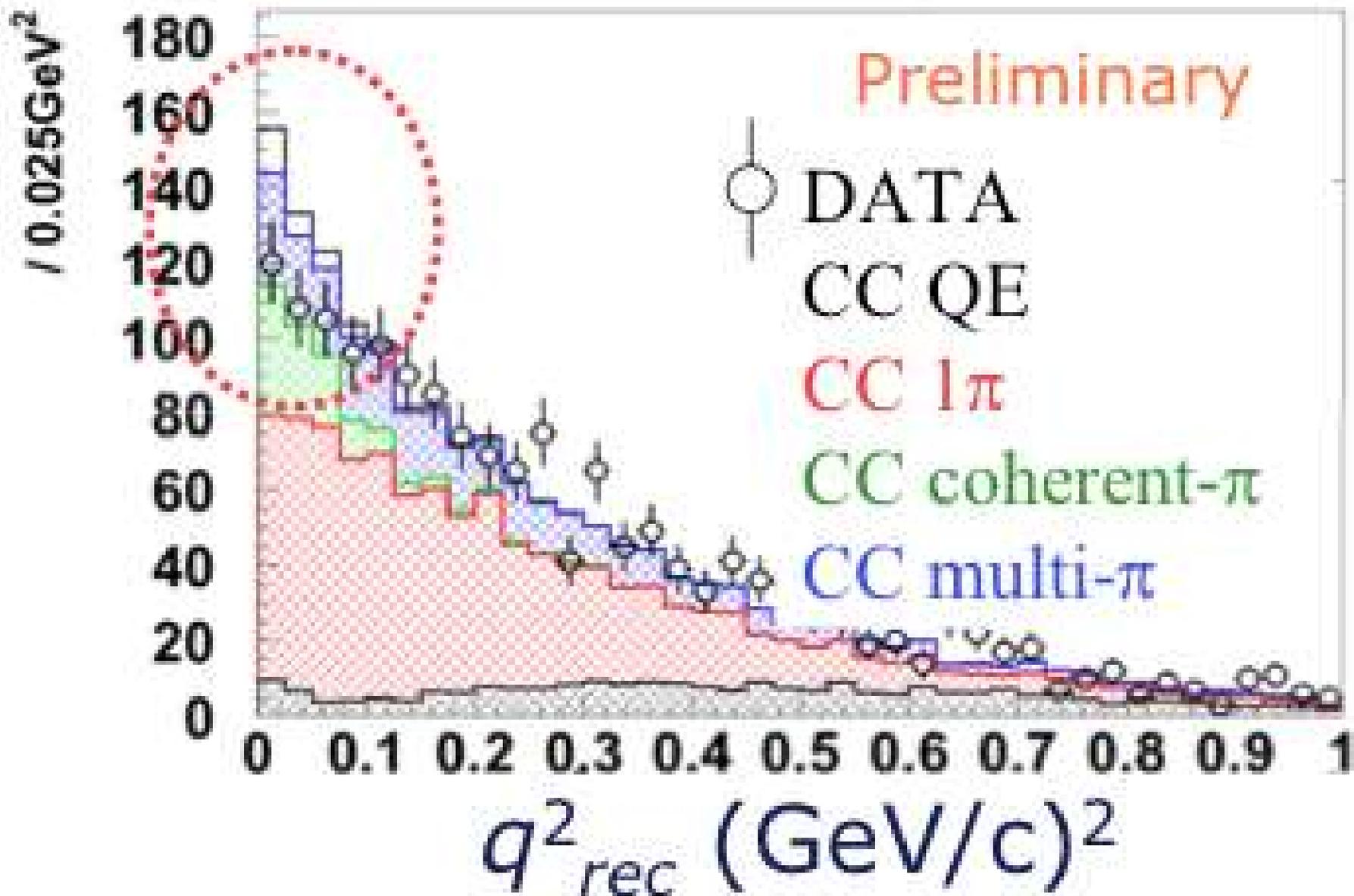
$\nu$  Cross Sections!

# K2K Near Detectors



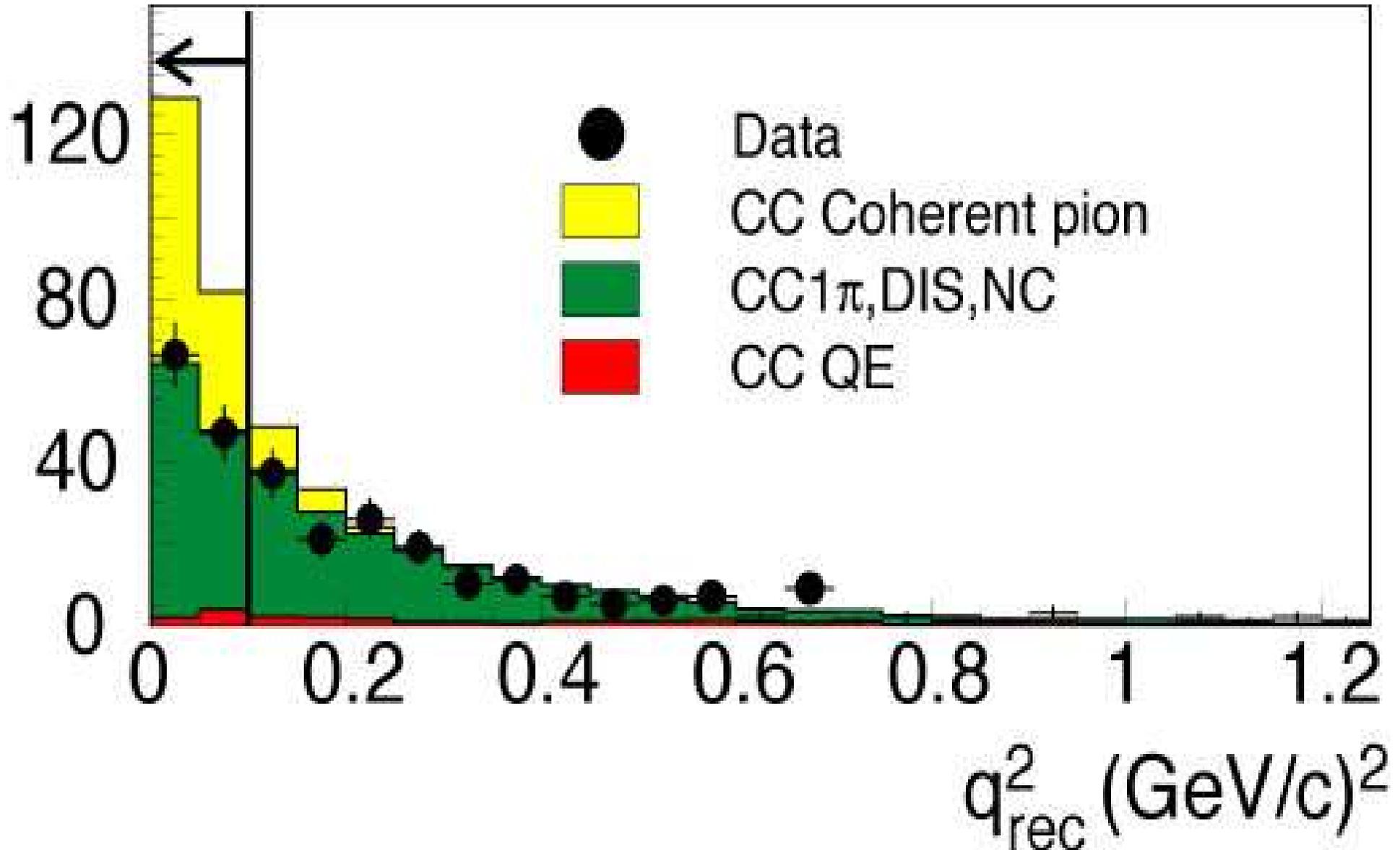
# Low- $q^2$ puzzle in K2K SciBar Detector

1.7E19 POT

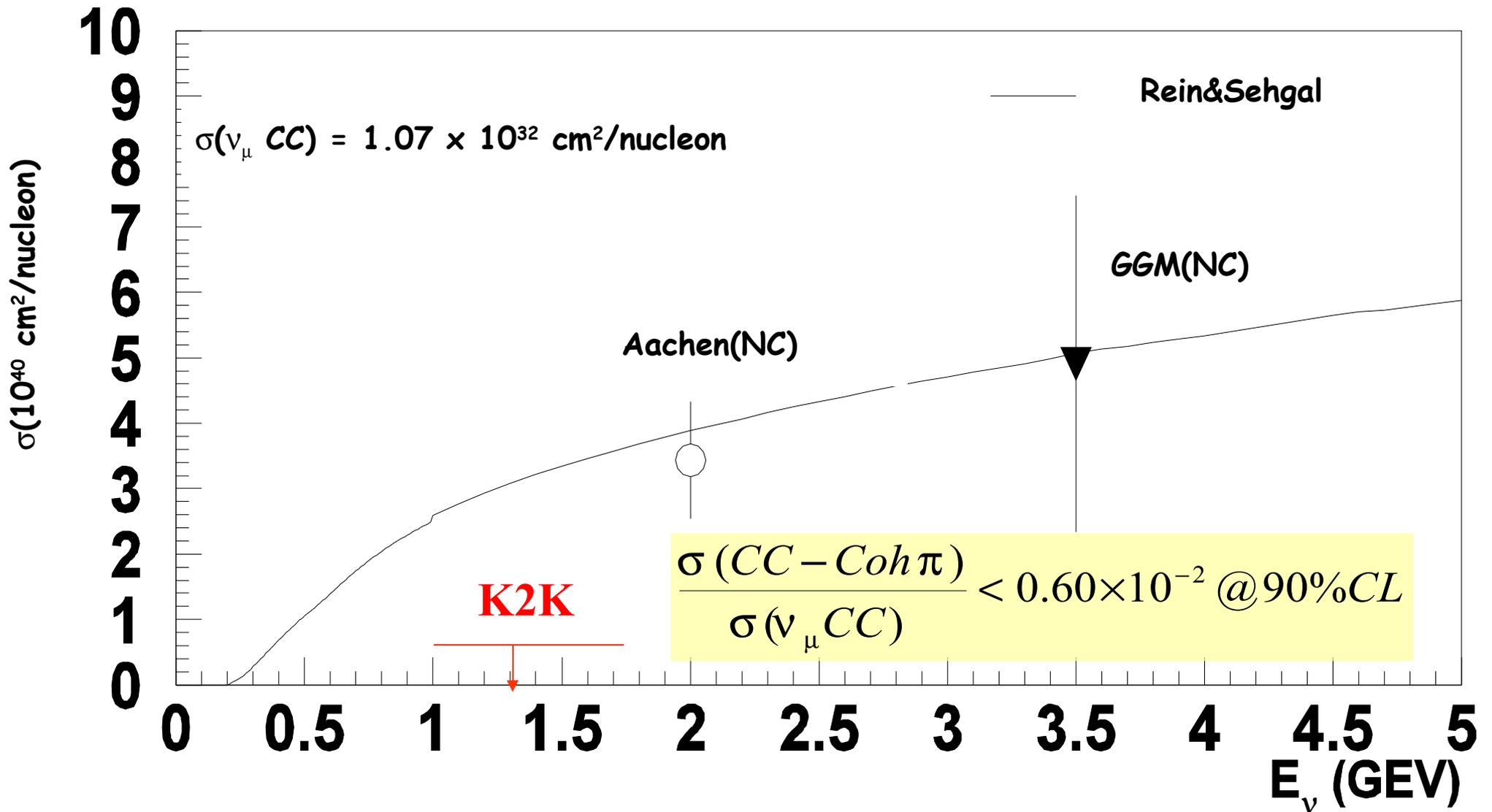


# K2K observes no evidence for $CC\pi^+$ coherent scattering!

hep-ex/0506008:  $CC\pi^+/CC = (0.04 \pm 0.29 + 0.32 - 0.35)\%$   
 $< 0.60\%$  (90% CL)



# Comparison with other Experiments



# Conclusions

- Neutrino Physics has been spectacularly successful in recent years, culminating with the discovery of Neutrino Oscillations and Neutrino Mass!
- Nevertheless, this may be just the “Tip of the Iceberg”!
- The ongoing MiniBooNE and MINOS experiments at Fermilab will make definitive tests of the LSND result and measurements of atmospheric oscillations
- Exciting future neutrino experiments are planned:  
Tritium  $\beta$  Decay, Double  $\beta$  Decay, Reactor  $\theta_{13}$ ,  
Accelerator  $\theta_{13}$ , Solar  $\nu$ , &  $\nu$  Cross Sections