

Chapter 3

The Muon Arms Detector Design

In this chapter the details of the Muon Arms design will be presented. It will also cover the current R&D results and construction status of the various components of the detector.

3.1 Muon Magnets

There are two muon magnets in the PHENIX detector. The North Muon Magnet was in the PHENIX baseline. The South Muon Magnet is the result of the upgrade of PHENIX to include a Spin Physics program. Because of constraints imposed by the existing PHENIX detector hall, the South Muon Magnet could not be identical to the North Muon Magnet. This section will describe both of the magnet designs as they currently exist as well as the muon identifier steel which is behind the magnets in each arm [25].

3.1.1 North Muon Magnet

The design criteria for the North Muon Magnet was to cover from 1.1 to 2.4 unit of rapidity (10° to 37°), have large acceptance in ϕ for low cross-section measurements, have a p_T kick of 200 Mev/c at 15° , and have an $\int B \bullet dl$ equal to 0.715 Tesla-meter at 15° . Figures derived from Vector Fields magnetic simulation software, shown in Figures 3.1, 3.2 and 3.3, illustrate the potential lines, the flux density in gauss and the saturation in the iron, respectively, for the final design of the North Muon Magnet. Table 3.1 shows $\int B \bullet dl$ along a line at 15° , 20° , 25° and 30° in θ , from 1.8 to 6.29 in z .

The design for the North Muon Arm magnet has had changed very little since the writing of the PHENIX CDR Update [2]. What changes have been made are the following:

1. The access door at the top of the "teacup" assembly, and its internal ribs have been removed.
2. The depth of the notch in the piston, where the station 2 tracking chambers will be located, will now only be 9.0 cm deep whereas it was defined as 12.0 cm deep.

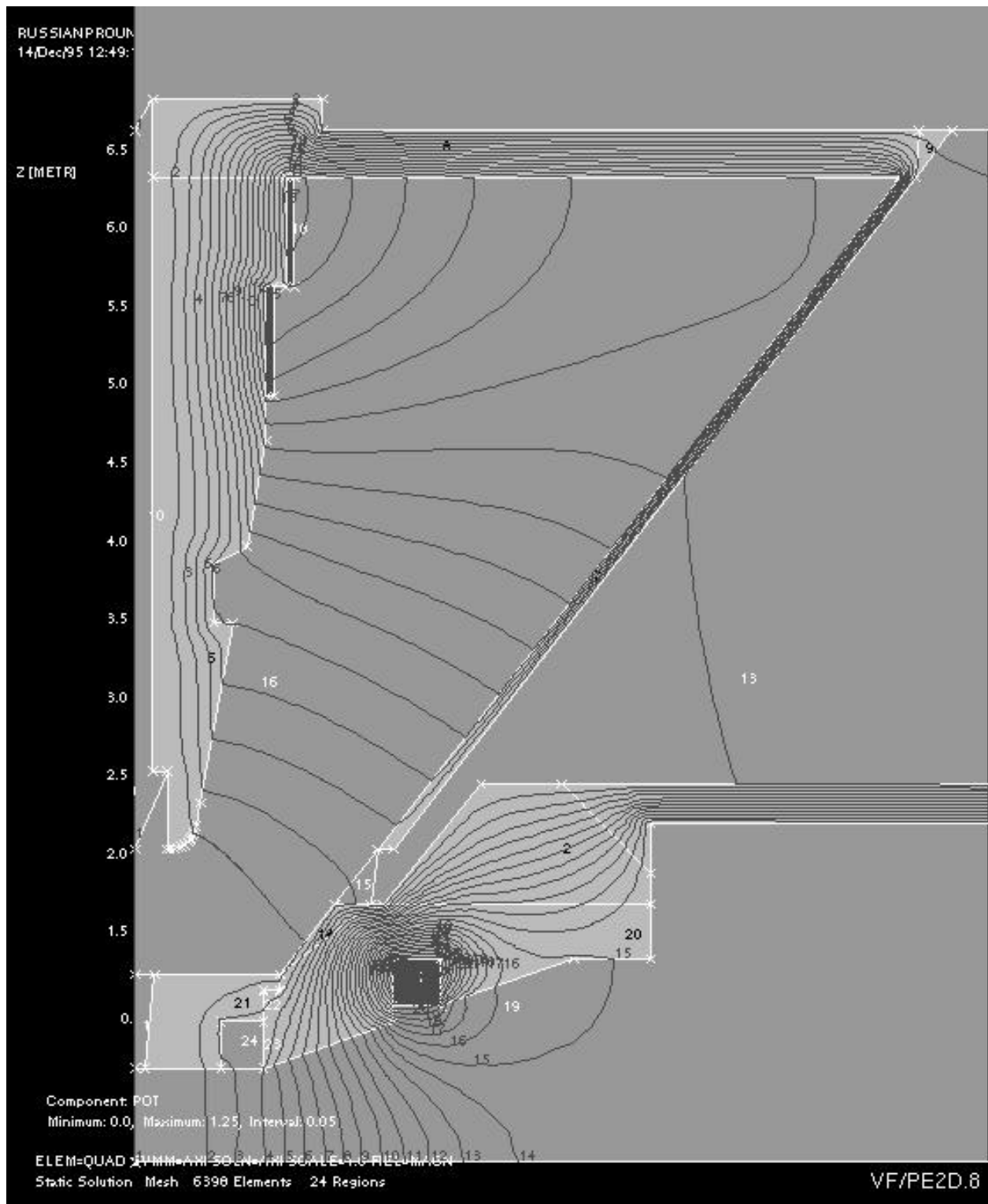


Figure 3.1: A plot of the potential lines of the PHENIX CM and NMM. The radius from the beam axis is shown on the horizontal axis and the distance away from the vertex (z -direction) is shown on the vertical axis. The scale for both axes is meters.

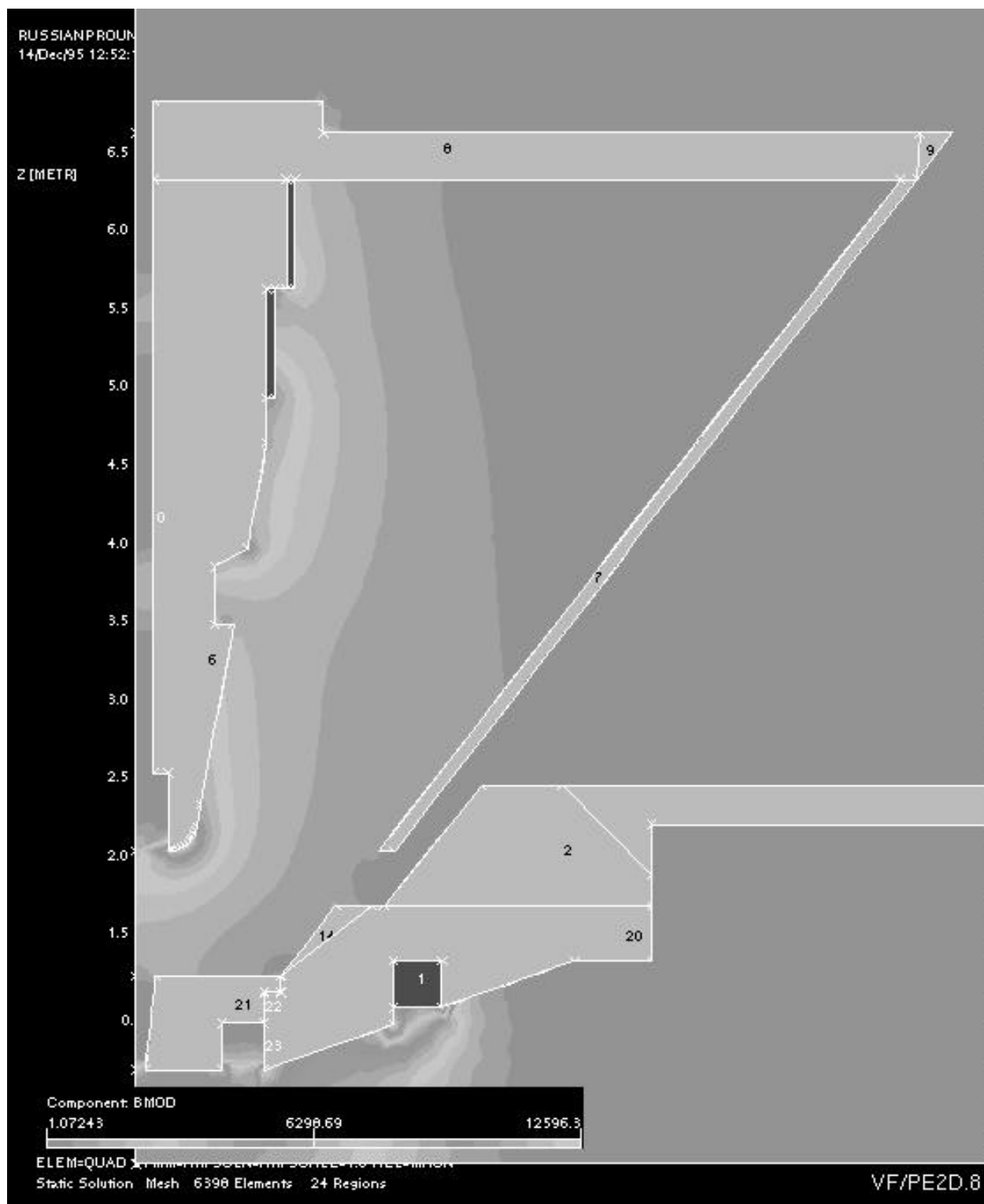


Figure 3.2: Shaded regions on the CM and MM show the levels of flux density within the magnet volume, units in gauss.

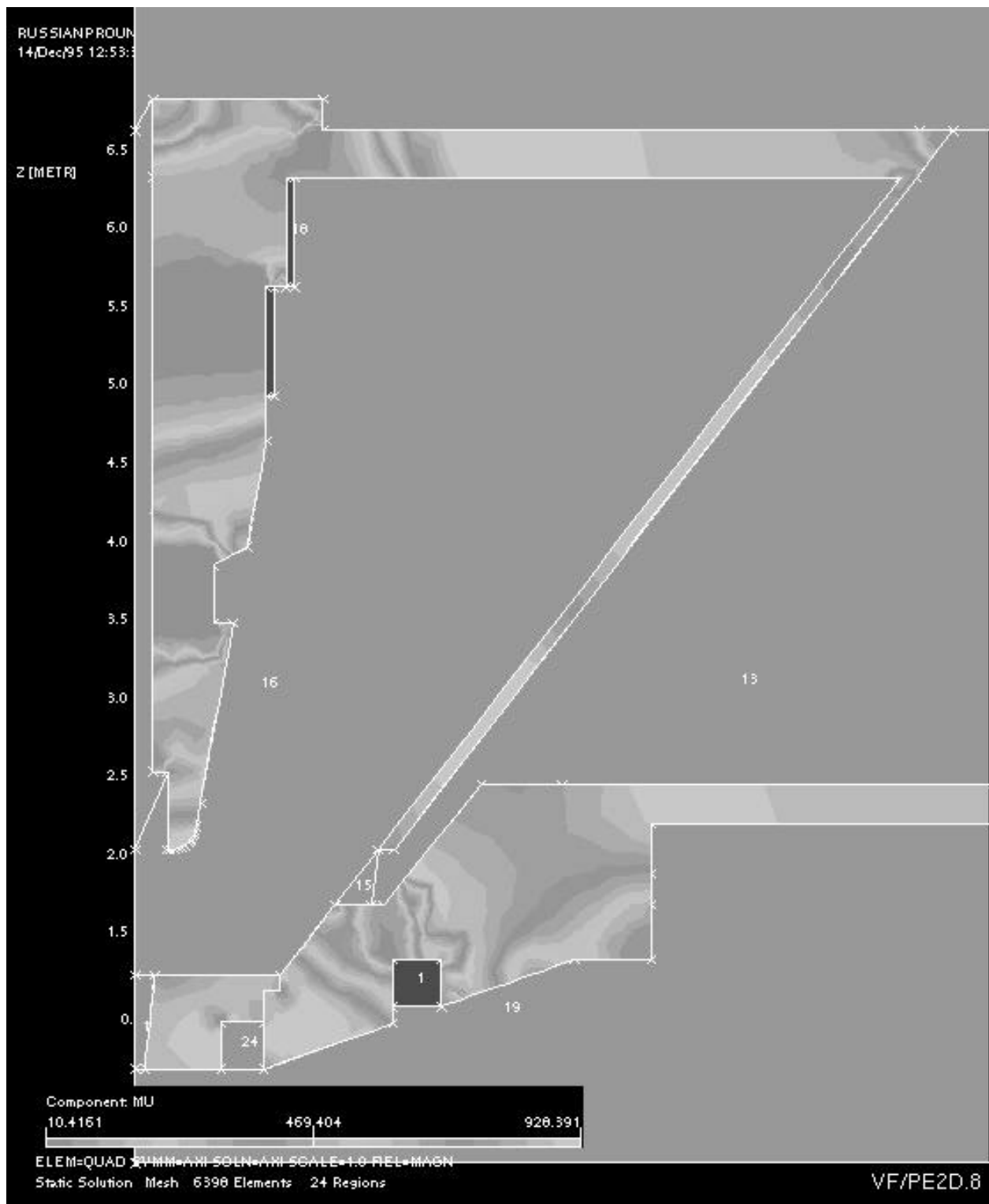


Figure 3.3: The shaded regions on the CM and NMM show the level of saturation of the steel. For example, the NMM piston is fully saturated (dark grey) under the smaller diameter coil. The least saturated regions are very light grey.

Angle	$\int B \bullet dl$ (gauss-meters)
15°	7,143.58
20°	4,802.95
25°	3,498.39
30°	2,687.55

Table 3.1: Integration of field values along a line, range of z values 1.8 m - 6.29 m.

3. The bolt on the low carbon steel iron plug for the front of the piston will have a full radius instead of a sharp corner. It was concluded that the fields at that corner were several kilogauss, which was too high for the proximity of the station 1 tracking chambers that mount off the front of the magnet.
4. After careful engineering analysis, it was concluded that neither the station 1 nor the station 2 tracking chambers should be mounted off the piston. Instead, they will be mounted off the most stable elements of the magnet assembly, the bottom three lampshade panels.

Figure 3.4 shows an isometric view of the central and north muon magnet with the north muon identifier absorber plates in place. Figure 3.5 shows a cross-sectional view of both the central and north muon magnet.

3.1.2 South Muon Magnet

In 1995, a Spin physics program was added to the PHENIX baseline. This has resulted in the addition of a South Muon Arm to the PHENIX detector. Central to the South Muon Arm is the South Muon Arm Magnet. The conceptual design of that magnet has been given in a muon arm technical note [26] and will only be summarized here.

The design criteria for the South muon magnet are the following:

1. To the maximum extent possible, utilize the existing North Muon Magnet design concept in order to take advantage of the proven design details and subsystem interchangeability.
2. The maximum weight of any single component of the South Muon Magnet should not exceed 40 tons in order to allow the existing PHENIX detector hall crane to be used for the magnet's assemble.
3. Provide Central Magnet accessibility by designing for full retraction of the South Muon Magnet up against the South Muon Identifier. That is a distance of approximately 1.5 m.
4. Minimize the number of hall track components by following the same design criteria used for the Central Magnet and the central detector arms.

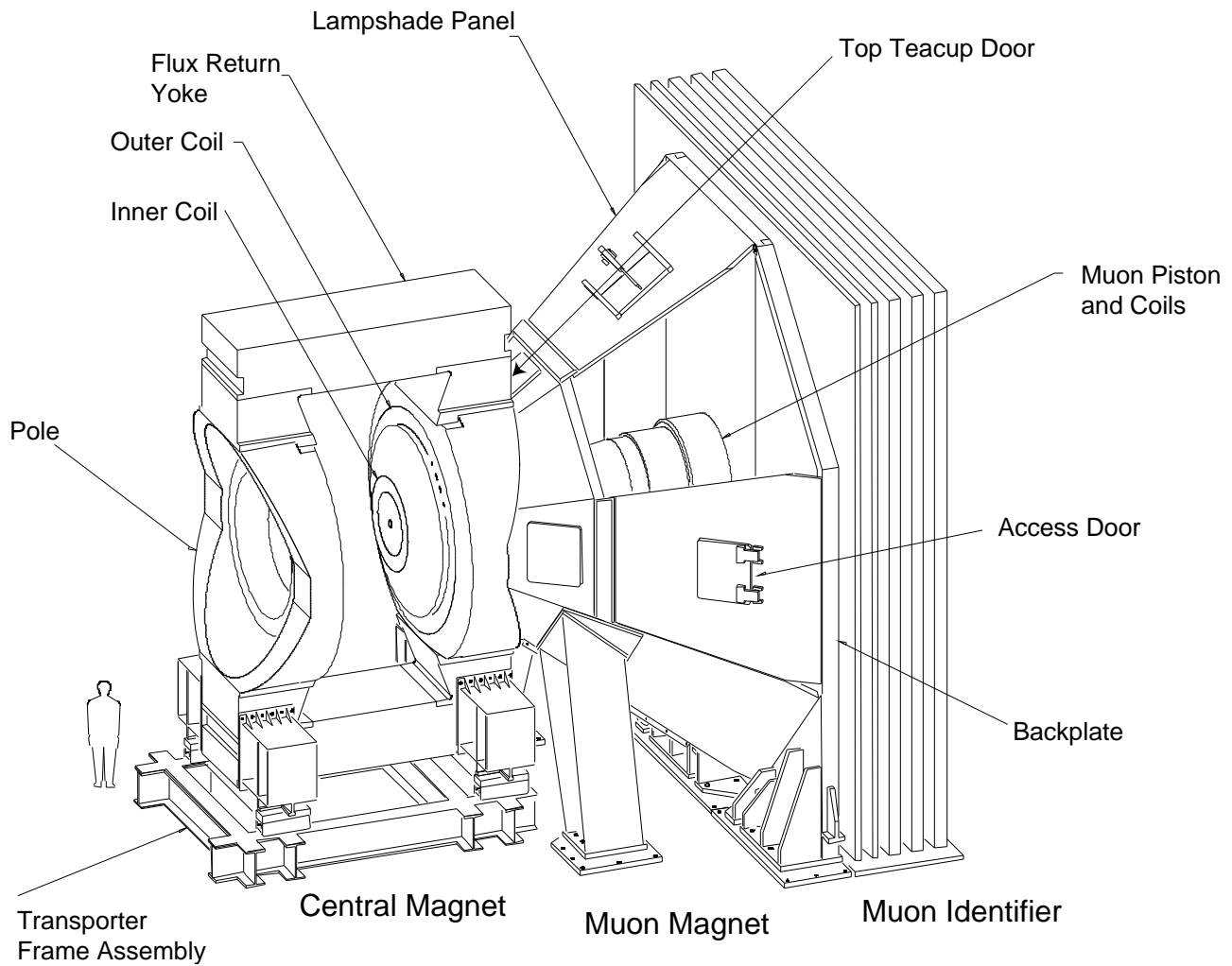


Figure 3.4: Isometric View of the PHENIX Central and North Magnet Subsystems.

5. Try to design for similar a p_T kick of 200 Mev/c at 15° .

The result is a design shown in Figures 3.6 and 3.7.

The two most striking differences between the two magnets is that the South Muon Magnet is 1.5 m shorter, and the piston has to follow a larger angle (12°). This design was developed out of Vector Fields magnetic modeling. Figures 3.8, 3.9 and 3.10 show the potential lines, the flux density in gauss, and the saturation of the iron for the final design of the South Muon Magnet. Table 3.2 shows $\int B \bullet dl$ along a lines at 15° , 20° , 25° and 30° in θ from 1.8 to 4.29 meters in z .

3.1.3 Tracking Chamber Placement In the Magnets

If one studies the two plots of the potential lines for the two muon magnets (Figures 3.1 and 3.8), it will be noticed that the placement of the second tracking station should not

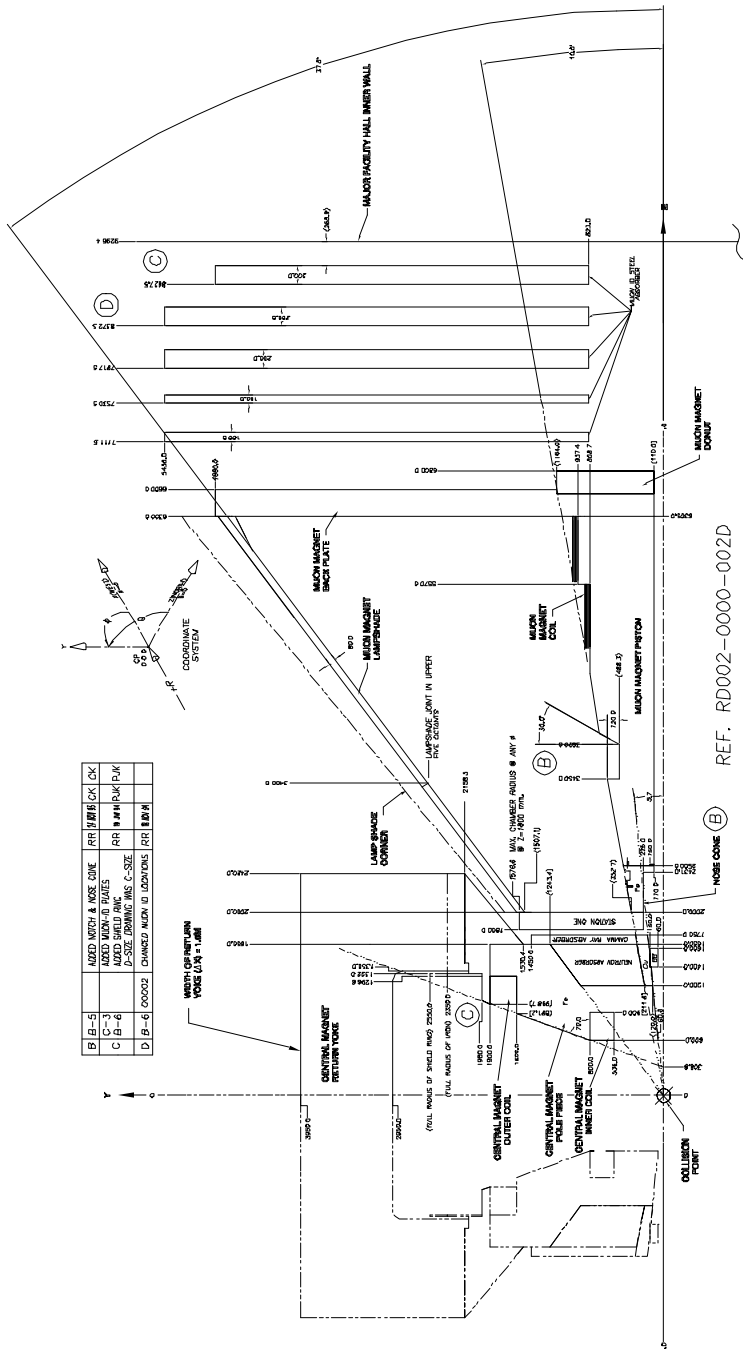


Figure 3.5: Configuration control drawing showing Central and North Muon Magnet in cross-section.

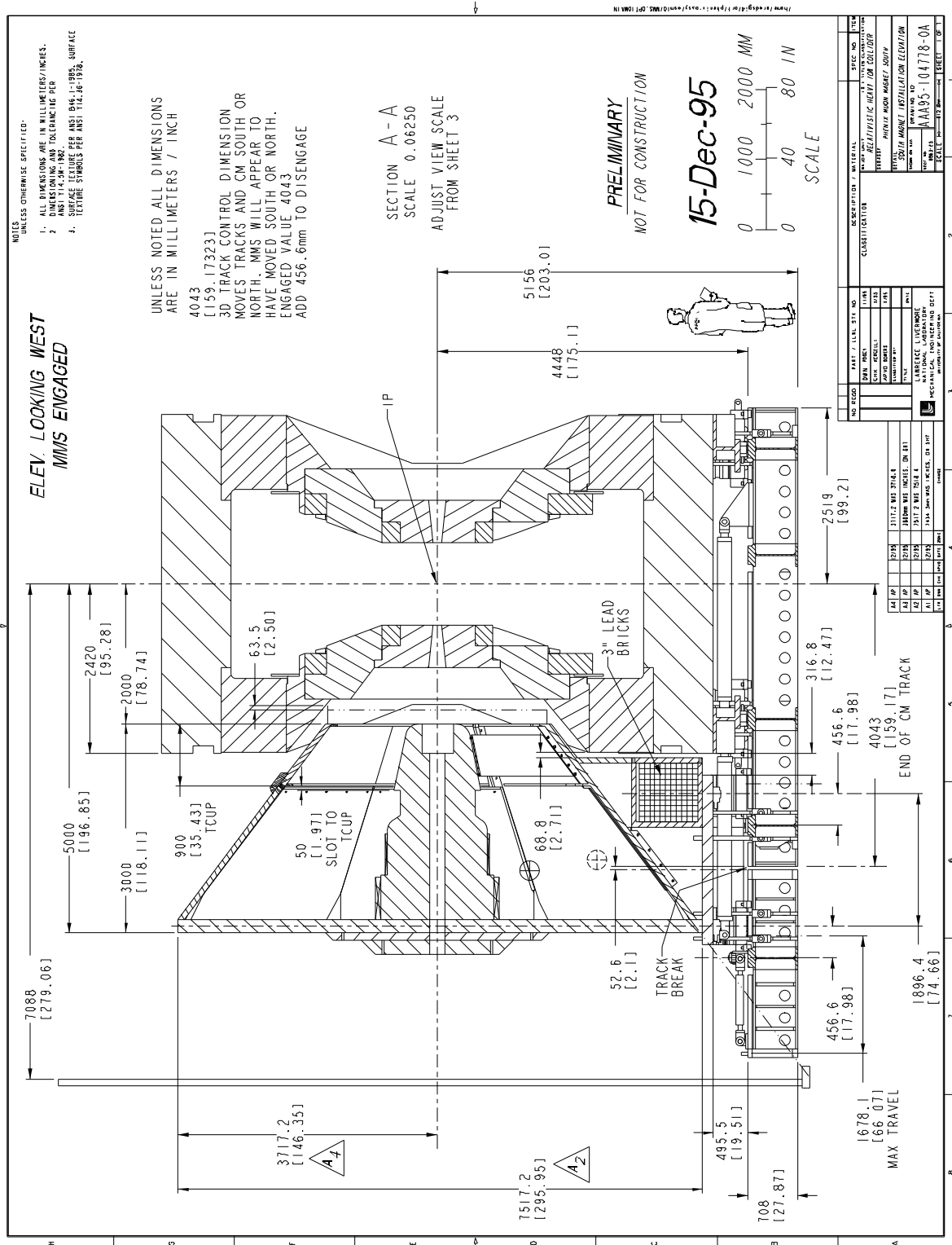


Figure 3.7: Cross-section View of South and Central Magnets.

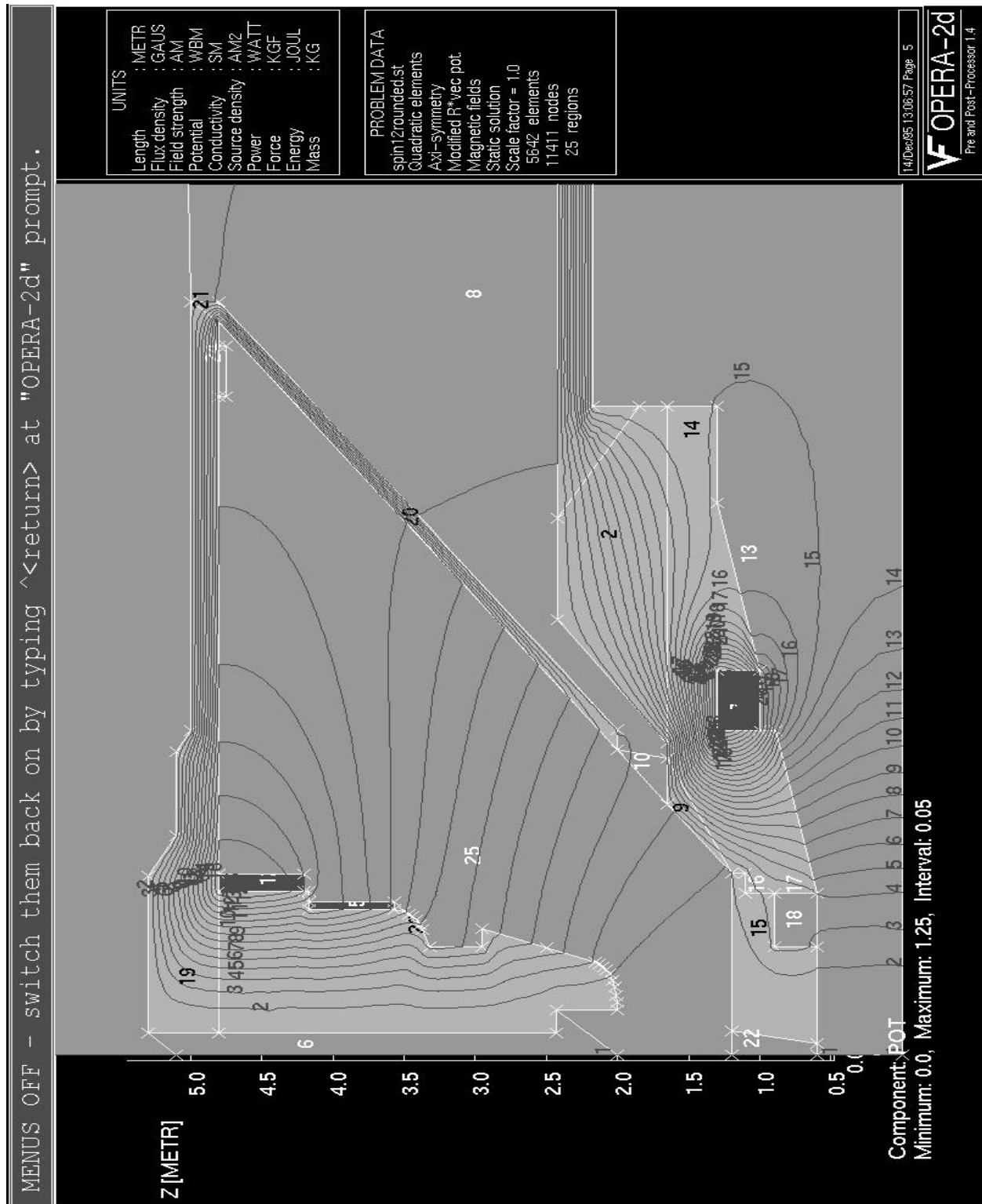


Figure 3.8: A plot of the potential lines of the PHENIX CM and SMM. The radius from the beam axis is shown on the horizontal axis and the distance away from the vertex (z -direction) is shown on the vertical axis. The scale for both axes is meters.

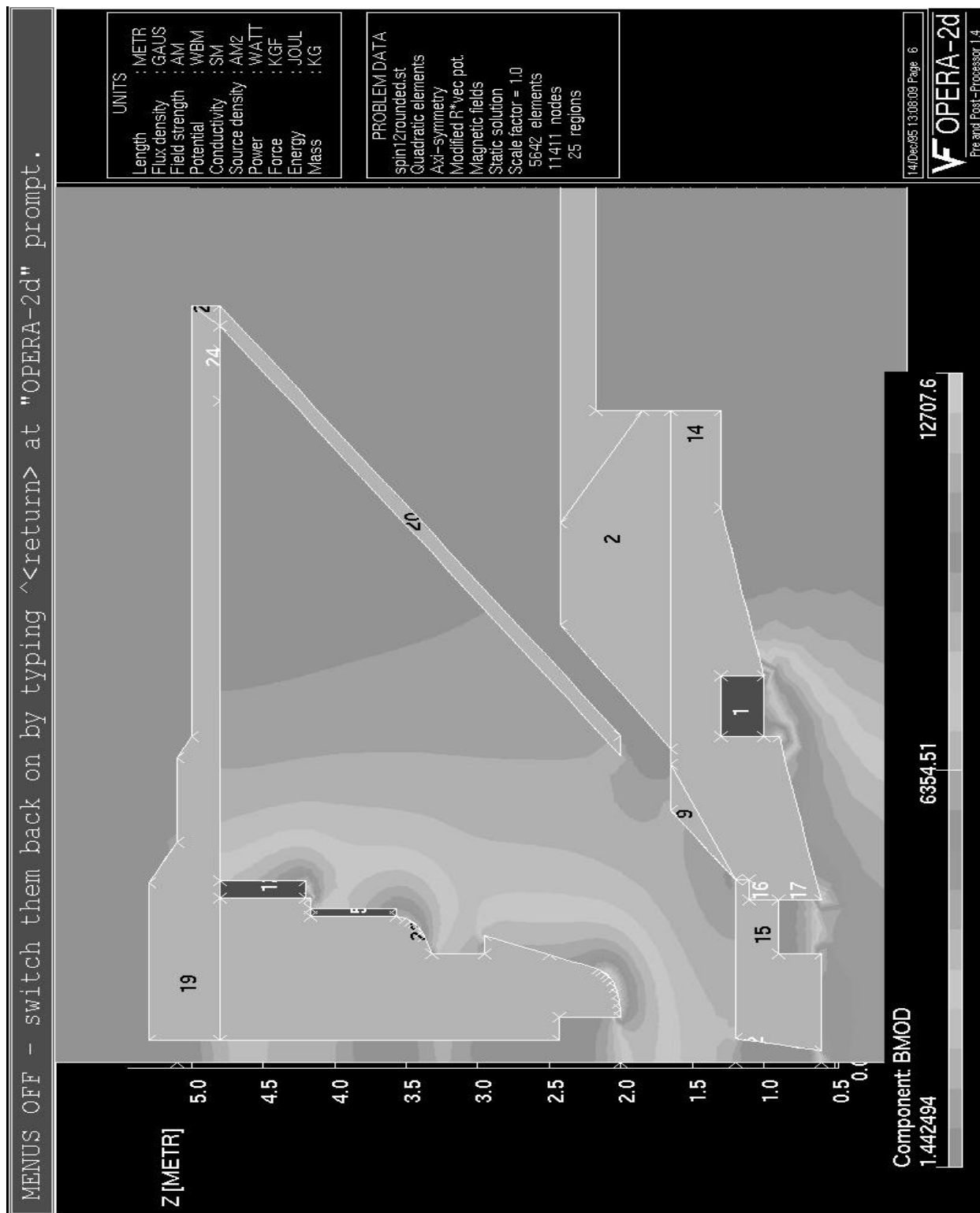


Figure 3.9: Shaded regions on the CM and SM show the levels of flux density within the magnet volume, units in gauss.



Figure 3.10: The shaded regions on the CM and SMM show the level of saturation of the steel. For example, the SMM piston is fully saturated (dark grey) under the smaller diameter coil. The least saturated regions are very light grey.

Angle	$\int B \bullet dl$ (gauss-meters)
15°	7,679.54
20°	4,931.30
25°	3,447.34
30°	2,557.81

Table 3.2: Integration of field values along a line, range of z values 1.8 m - 4.29 m.

be at the mid-point of the length of the piston. Indeed, a careful calculation was done to determine the z -position of the maximum radial deflection of charged particle trajectories in the magnet. The results of these calculations and the z -positions of the chambers for the North and South Muon arms are shown in Figures 3.11 and 3.12.

3.1.4 Muon Identifier Steel

The Muon Identifier Steel which will be used in the detector has been described in detail in the PHENIX CDR Update [2]. In summary, the Muon Identifier Steel consists of five steel plates. The first two are 10 cm thick and the last three are each 20 cm thick. The muon identifier steel for both the North and South Muon Arms are identical.

Since the writing of the CDR Update, the following modifications of the Muon Identifier steel have been made:

1. The muon identifier detector panels are now all of equal thickness. This requires new z locations for the steel absorber rows.
2. The center-plate of the second steel absorber plate is 200 mm narrower than the others. The side plates in this row are the same as in other rows, so the overall width of this row is narrower. This modification helps to hide the cracks between adjacent plates from row to row.
3. The last muon identifier plate and the corresponding detector extend under the experimental hall's bridge crane support. Consequently, the last absorber plane and detector planes 5 and 6 are shorter than the others.

The modifications to the North Muon Identifier can be seen in Figure 3.5 and are detailed in Phenix Technical Notes [27, 28].

3.1.5 Current Status of the Muon Arms Magnets and Identifier Steel

By the time of the upcoming "Muon Arms Review" (January, 1996), representatives from the PHENIX collaboration will have traveled to the Izhora Steel Works in Russia to inspect a trial assembly of the North Muon Magnet. Based upon their inspection of all the steel for

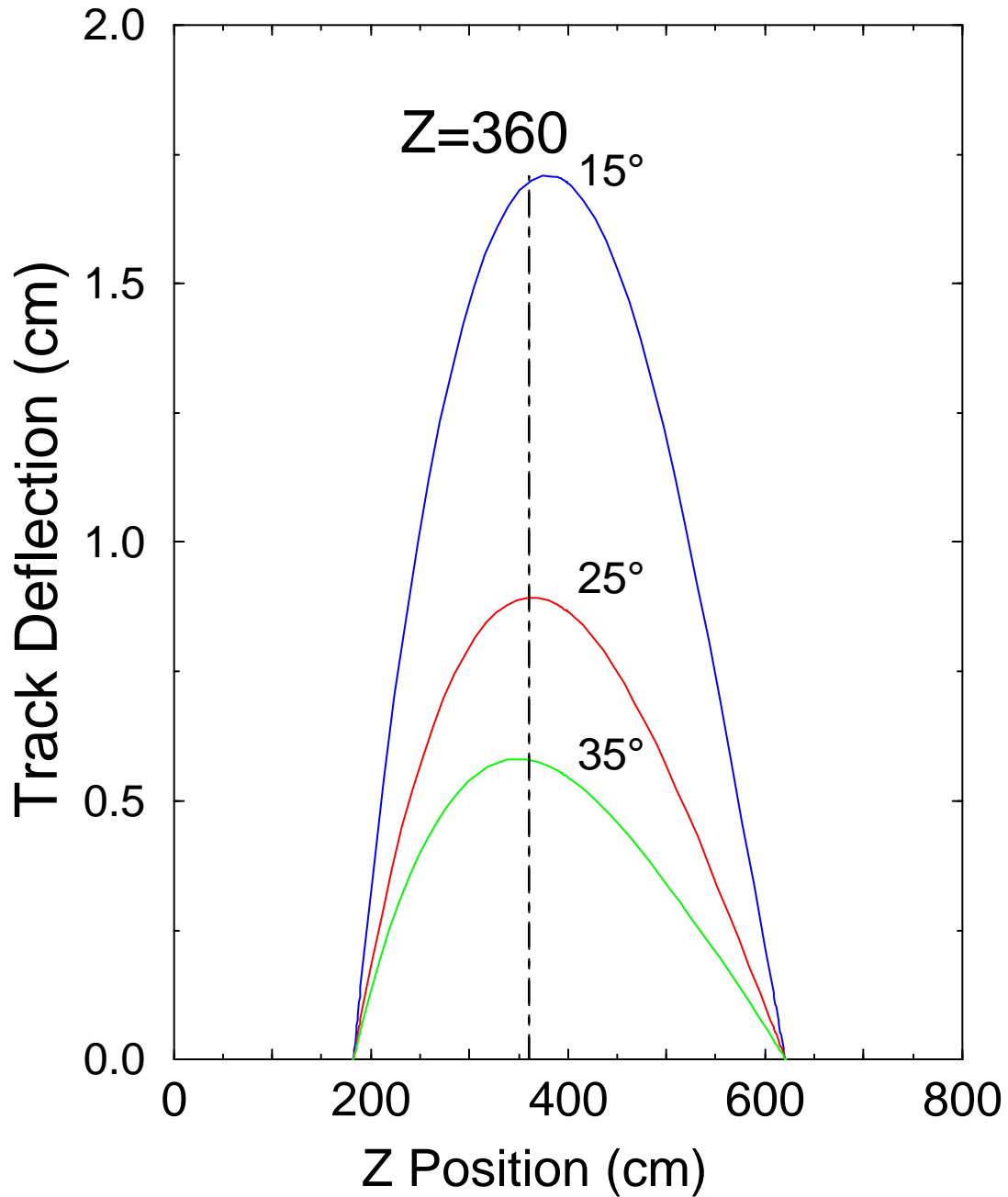


Figure 3.11: The track deflection vs z position for 10 GeV/c μ 's in the North Moun Arm Magnet at various θ angles. The optimum position for the Station #2 chambers is shown on each curve.

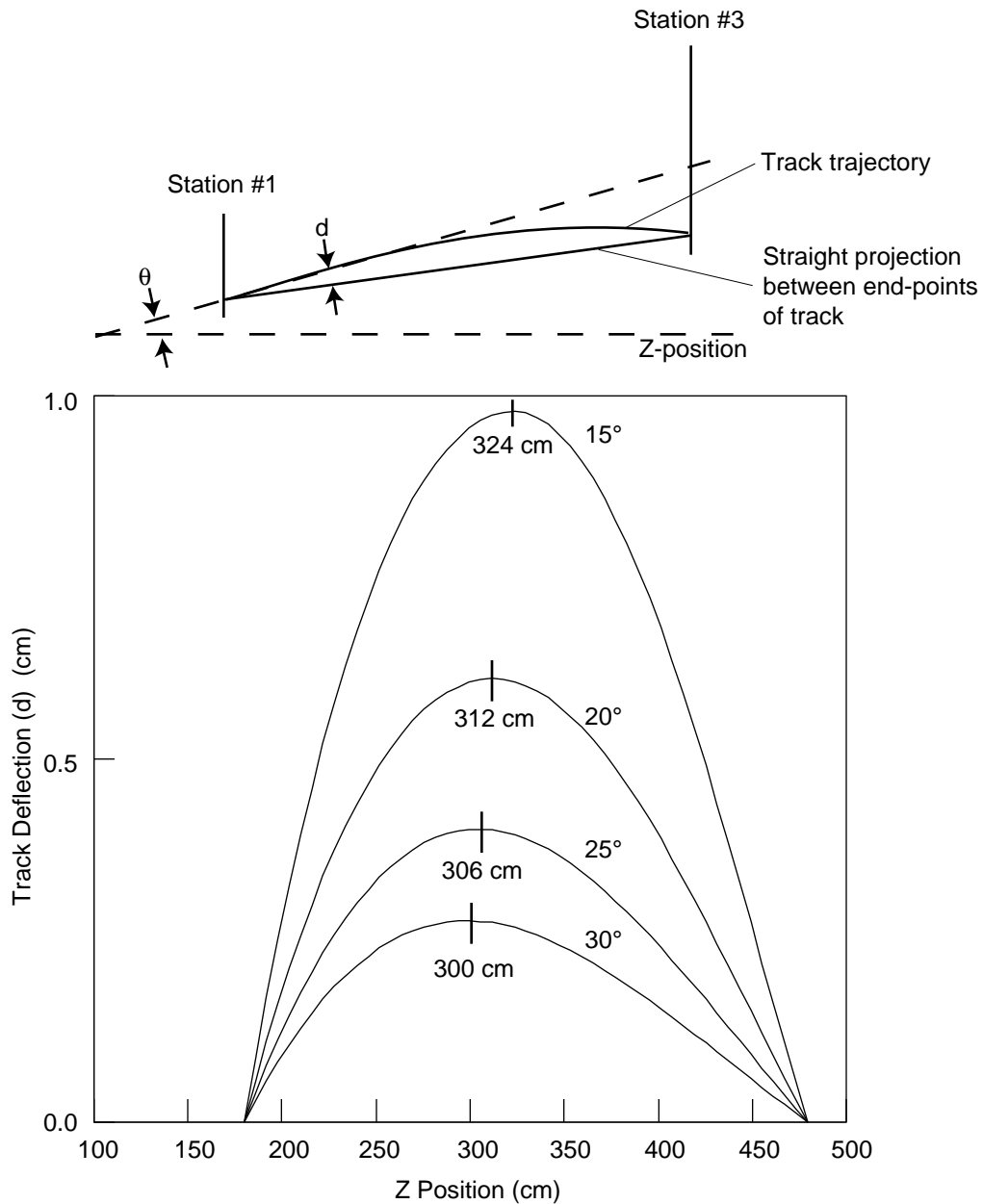


Figure 3.12: The track deflection vs z position for 10 GeV/c μ 's in the South Moun Arm Magnet at various θ angles. The optimum position for the Station #2 chambers is shown on each curve.

that magnet, as well as the Central Magnet and the steel plates for both Muon Identifiers, all steel for the North Muon Magnet, Central Magnet, and both Muon Identifiers will be shipped to a port of entry in the metropolitan New York area.

A contract with our Japanese collaborators (a Japanese industry) to construct the South Muon Magnet and the coils for both the North Muon Arm Magnet and the South Muon Arm Magnet is nearing completion.

A contract has been signed between the PHENIX collaboration and Rigging International, of Oakland, California, to offload the magnet and identifier pieces from their trans-Atlantic ship and transport them to the major facility hall at the RHIC accelerator complex at BNL. Rigging International will also do the necessary rigging for installation of the steel into the experimental hall.

3.2 Absorbers

Several absorbers are a part of the muon spectrometers. These absorbers are continually being developed and are described in some detail in the PHENIX CDR [1], Sections 10.2.2 and 10.4.1, and also in the CDR Update [2], Sections 10.4.2 and 10.6.1. They consist of the copper nose cone, the copper insert (or flowerpot), the lead curtain, the neutron absorber, and the iron plug. The ones for the North Muon Arm are all shown in Figure 3.13. There is a similar set for the South Muon Arm.

The copper nose cone is bolted on the central magnet pole tips. The purpose of the nose cones are to reduce the flight path for hadrons from the collision point to the hadron absorbers (starting with the central magnet pole pieces). This results in a decrease in the number of muons from weak decays that enter the muon arms. Mechanically, the angle of the two nosecones on the North and South muon arms follows the angle of the central magnet pole pieces (70°) from the beam axis.

The copper insert, or flowerpot, is attached to the downstream sides of the central magnet pole pieces. The inserts serve two primary purposes. First, they are the mounts for the beam-beam counters, and second, they stop particles from shower leakage out of the central magnet pole piece from entering the tracking chambers. Additionally, they provide more absorber material for the more energetic primary particles at very forward angles which would strike the front face of the piston. The inserts have an outer tapered edge following the 10° and 12° lines matching the muon magnet piston angles for the North and South arms, respectively. They are both bored out to accomodate the beam-beam counter on each side of the vertex.

There is space available for two other absorbers. One, the so-called lead curtain, would suppress photon and e^-e^+ leakage from the back of the central magnet pole piece, and the other, the neutron absorber, would reduce neutron leakage. The lead curtain is to consist of 7 cm of lead and the neutron absorber of borated polyethylene. The neutron absorber would be placed on the downstream face of the central magnet pole pieces and the lead curtain would be attached to the back of it.

The last absorber is the iron plug. They are the front part of the muon magnet pistons. They are bored out to almost twice the diameter of the piston bore so that particles will not shower off the front of the piston into the first tracking station.

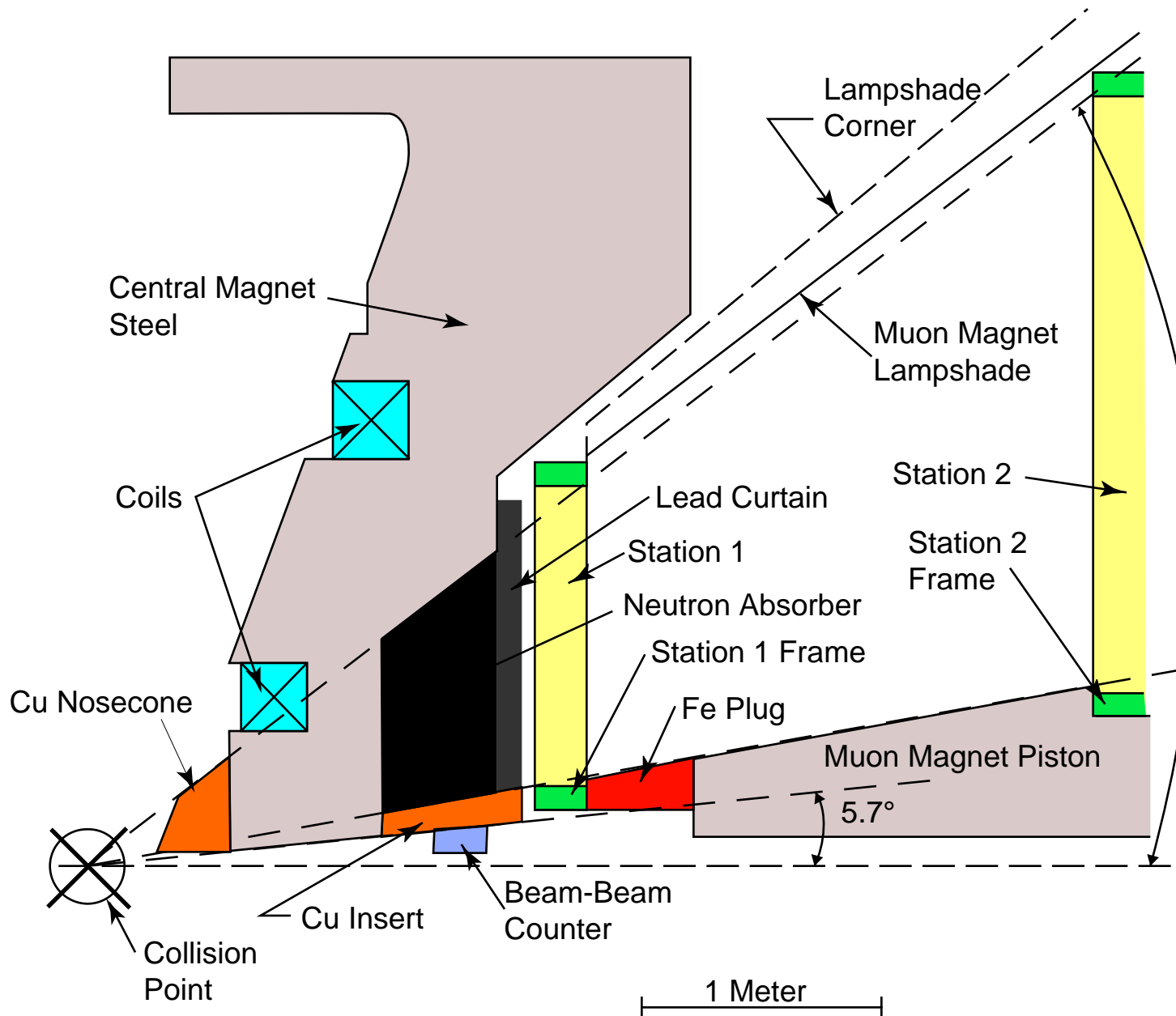


Figure 3.13: A view of the North Muon Arm showing the layout of the specialized absorbers. The absorbers for the South Muon Arm are essentially a mirror image of these with the exception that the plug on the South Arm has a 12° outer angle to match the South Arm piston.

3.3 Tracking Chambers

This section will describe the kind of chambers which will be used for muon tracking in the Muon Arms spectrometer. It will then present the design details for the chambers and some information on how they will be manufactured. Finally, some R&D results on various prototype chambers will be presented.

3.3.1 Introduction to Cathode Strip chambers

The basic mechanical construction of the cathode strip chamber CSC is that of a conventional proportional wire chamber. However, rather than reading out the anode wire to get discreet position resolution, a precise position coordinate is determined from the induced signal on cathode strips perpendicular to the anode wires. Resolutions as small as $50 \mu\text{m}$ have been achieved on chambers as large as 2 m [29]. For the CSC to serve as the muon system tracking chambers, several issues were addressed. These include spatial resolution, identification of a suitable gas with a small Lorentz angle, radiation length, preparation and metrology of the cathode foils, and mechanical design issues related to frame stiffness, frame thickness, and creep. Cost and simplicity in design were additional factors.

CSC signals

The charge on the anode is due to ionization in the gas caused by passage of the muon through the gas volume and the gain of the chambers. In a typical gas volume the ionization process produces about 140 electrons/cm so in the CSC chambers the number of electrons produced will be $140 \times 0.7 \text{ cm}$ electrons or 100 electrons per track. Assuming a gain in the chamber of 2×10^4 , the anode charge will be 2×10^6 electrons. The cathode charge is one half of the anode charge times a reduction factor called the “ballistic deficit factor” that is due to the finite integration times. Taking a ballistic deficit factor of 0.5, the total cathode charge will be 5×10^5 electrons, or 80 fC.

Cathode Strip Chamber Design Issues

The primary design issues addressed have been factors effecting the intrinsic resolution of the chambers, the mechanical and electronic requirements, and the definition of a suitable gas. Where possible we have relied on the work of previous authors and in particular the vast amount of work done by the Muon Group and the Central Tracker Group of the former GEM collaboration at the SSC.

CSC Candidate Gases

It is desirable to have a gas with a small Lorentz angle so resolution degradation due to a large Lorentz angle can be minimized. Fortunately, a number of fast gases have been identified and tested that are suitable for the CSC’s. Most of these gases contain CF_4 in combination with isobutane or CO_2 . The Lorentz angle is similar for each gas and is about

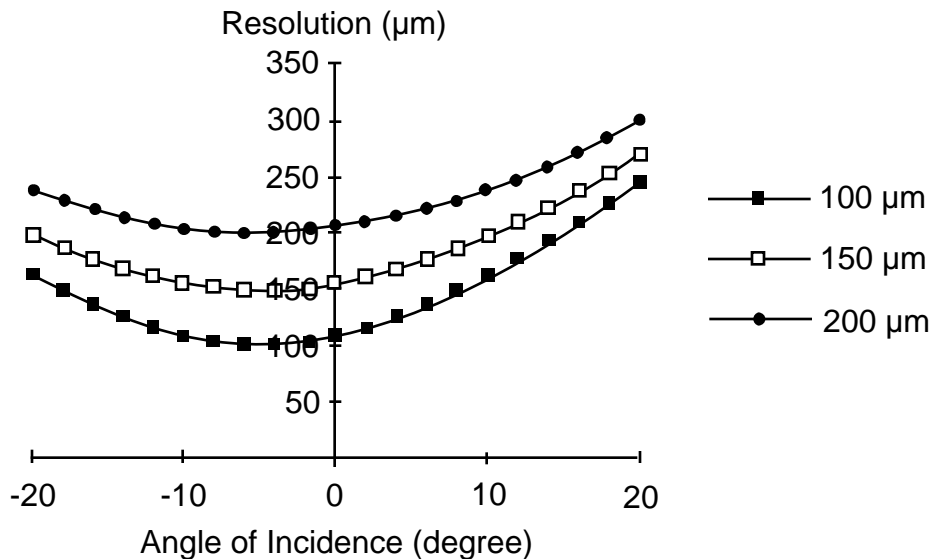


Figure 3.14: Resolution degradation with angle for different intrinsic resolutions.

5° at 0.5 Tesla magnetic field. Most of the prototype work has been done with a 50% CF_4 , 50% isobutane mixture, but CO_2 is also an acceptable replacement for the isobutane.

Resolution Degradation Due to Inclined Tracks and Lorentz Angle

Resolution degradation of the CSC's comes primarily from two causes, tracks inclined from the normal to the face of the chamber, and Lorentz angle. In both cases, the position resolution is degraded because the deposited charge is distributed nonuniformly along the anode wire due to the energy loss fluctuations in the gas. These effects have been extensively studied by the former Muon Group in the GEM collaboration and reported in a GEM note [30]. Additional studies by Musser, *et al.* [31] for the GEM Central Tracker Group on a CSC type chamber very similar to the proposed PHENIX CSC's and for a suitable gas (50% CO_2 , 50% CF_4) have shown that the resolution degradation, d , is equal to

$$d = 0.08l \tan \phi \quad (3.1)$$

where l is the cathode to cathode spacing and ϕ is the angle from the normal to the face of the chamber. Including Lorentz angle smearing is accomplished by modifying the angle of incidence by the Lorentz angle, φ , i.e. $\phi - \varphi$. In Figure 3.14 and Figure 3.15, we show the chamber resolutions when these effects are taken into account. The Lorentz angle was taken to be 5° . In Figure 3.14 the resolution is plotted for chambers with different intrinsic resolutions and the chamber gap = 6 mm. In Figure 3.15 the resolution is plotted for different chamber gaps with the resolutions fixed at $100 \mu\text{m}$. It is clear that we want thin chambers and small angles of incidence and resolutions approaching $100 \mu\text{m}$.

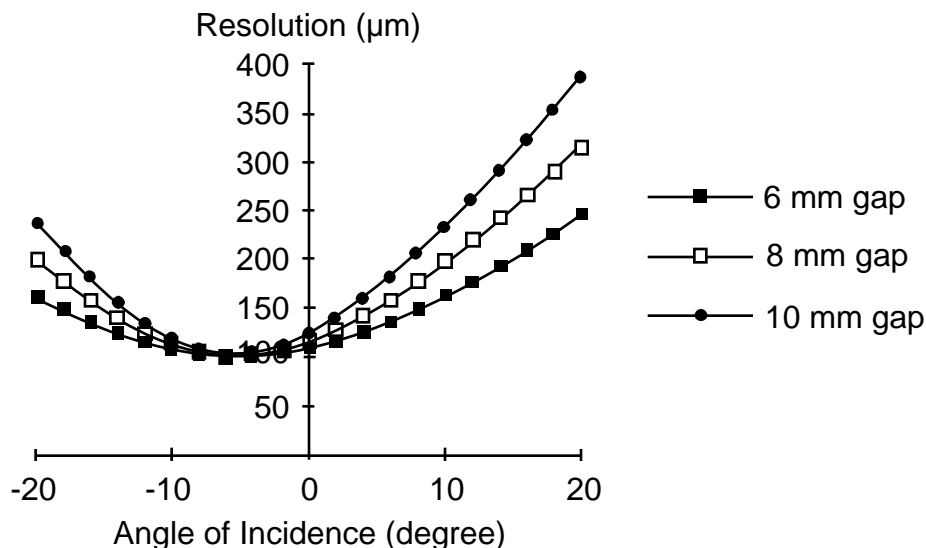


Figure 3.15: Resolution for different chamber gaps at $100 \mu\text{m}$ intrinsic resolution.

3.3.2 Cathode Strip Chambers Design Constraints

The simulation studies have shown that the mass resolution of the Υ continually increases with the resolution at each station until it is limited by the multiple scattering due to the absorber in front of the muon spectrometer. An Υ mass resolution approaching 200 MeV is achieved if the station resolution is $60 \mu\text{m}$. This can be achieved with 3 planes of CSC's with $100 \mu\text{m}$ resolution at each station. We have established this as the baseline design. Achieving $100 \mu\text{m}$ resolution for each plane of CSC's requires careful attention to the design of the cathode strips, the geometry of the drift cell, and signal-to-noise. Generally the dominant factor contributing to the resolution is the electronic noise.

An important aspect to the design of the individual stations is to insure that the detection efficiency is maintained in the face of failure scenarios. The simulation studies have shown that sufficient redundancy exists with three CSC detectors per station to insure 99% tracking efficiency. We have baselined three CSC detectors per station with three coordinate readout for each detector for a total of 9 measurements per station. Each detector will read out a high resolution cathode, a low resolution cathode, and the anodes. By reading out both cathodes and the anodes in each detector complete space point reconstruction is possible for each detector without ghost hit interference. Such a robust system simplifies pattern recognition and maintains high tracking efficiency in the event of failure.

The high resolution cathodes will incorporate the floating strip method to reduce the number of analog readout channels. Each strip will be 5 mm wide with the readout at 1 cm intervals. With this scheme we expect resolutions of less than $100 \mu\text{m}$. The floating strip will be tied to ground through a 500 kohm resistor. The three high resolution cathodes in each station can be oriented with one cathode each at -22.5° , 0.0° , and 22.5° with respect to the radial coordinate. We have not, however, studied whether using three different cathode orientations provides enough additional benefit to override the possible increase in cost.

The low resolution cathodes will also incorporate the floating strip method to reduce the number of readout channels. Each strip will be 1 cm wide with the readout at 2 cm intervals. The readout will be a digital latching scheme where no analog information is used. The resolution will be determined by the readout spacing and will therefore be approximately 6 mm. The strip orientation will be similar to the high resolution cathode but each detector will have a different high and low resolution cathode orientation.

The anode plane will incorporate a standard digital readout scheme with an alternating signal/field wire structure to enhance chamber stability. The anode-field wire spacing is 5 mm and the readout will be irregularly spaced from 5 mm to 2 cm to account for variations in occupancy between 10° and 35° . All anode planes in a station will have the same orientation.

3.3.3 Stations 1 and 3 Designs

We have incorporated honeycomb structures as the baseline design for the CSC chambers at stations 1 and 3. Several reasons led us to decide on the use of honeycomb structures. Since the radiation length budget ($<10\%$) for stations 1 and 3 is less restrictive than for station 2 ($<0.5\%$), we were not under the restrictions of having to use thin foils for the cathode structures. Honeycomb allows us to have chamber designs which have a larger acceptance than if metal frames were used. The cathodes can be made out of etched copper clad in a commercial process than reduces the manpower required and offers a less expensive chamber design than with foils. The designs for station 1 will serve as the prototype for the design of station 3 chambers.

3.3.4 Station 1

The station 1 design is shown in Figure 3.16. Station 1 chambers are designed to be quadrants mechanically but octants electrically. The design calls for four honeycomb structures to form three CSC chamber gaps. The total combined thickness is less than 10 cm. A cross-section of the assembly is shown in Figure 3.17. The pair of cathode surfaces on either side of the chamber gaps have cathode strips etched in copper clad epoxy fiberglass skins. One cathode surface of the pair will be used for the high resolution coordinate, ϕ , and the other cathode surface will be used for a low resolution stereo coordinate. The anodes follow the chord in each 22.5° segment to minimize effects due to the Lorentz angle of the gas. The anodes span octants only and are electrically isolated at 45° .

The support frame for the honeycomb panels are 1.8 cm thick by 2 cm wide lucite on the quadrant radial edges and are part of the inactive area of the chambers. The frame support at the inner radius is inside of 10° and the frame support on the outer radius is outside of 35° . At 45° , the anode support structure and central honeycomb support structure is 1.5 cm wide. This quadrant design maximizes the active area of the chambers.

Mechanical Assembly

We are prototyping two methods for manufacture of the honeycomb structures, a commercial manufacturing process and an “in-house” manufacturing process. The constraints are that

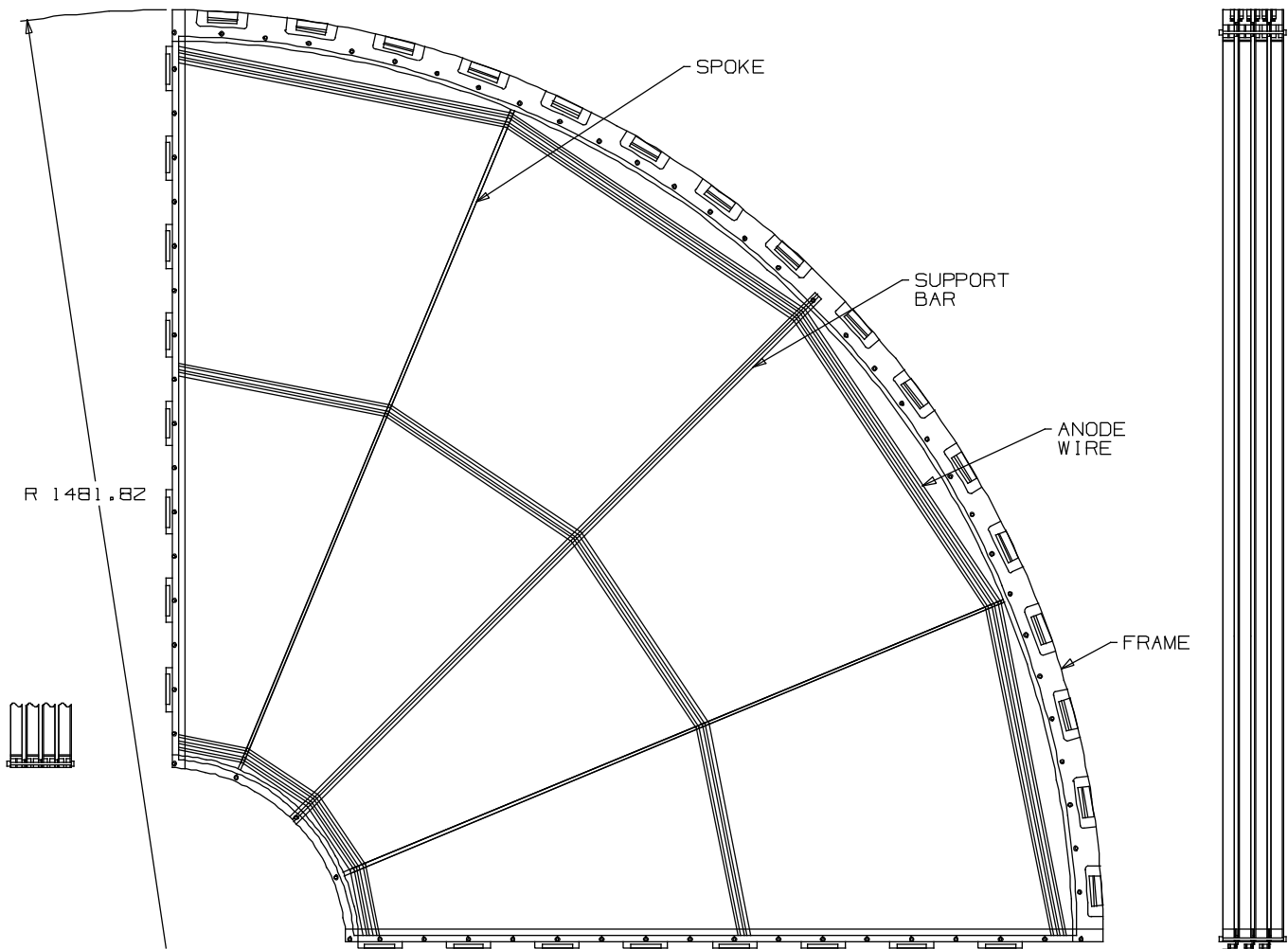


Figure 3.16: Station 1 Quadrant.

the flatness of the panels must be less than $150\ \mu\text{m}$. The advantages of the commercial process is less local manpower but at a higher material cost. In the “in house” method, the cost will be less but more manpower will be required. In both processes, the outer skins will be bare epoxy fiberglass, $250\ \mu\text{m}$ thick. Precision made and drilled etched cathode skins will be laminated to the panels in a subsequent process utilizing precision machined guide plates. The laminations for both will be done locally to control assembly tolerances. The etched cathodes will have alignment fiducials as part of the art work. The goal of the assembly process will be to control the position of the etched cathodes with respect to the alignment pins to better than $25\ \mu\text{m}$. The gap between panels will be controlled by the anode support structure and the edges of the quadrant and by the support structure at 45° . $50\ \mu\text{m}$ tolerance at these locations is expected. The maximum unsupported anode wire length is less than 58 cm. An anode PC board is used on the edges of the quadrants to distribute the HV and to transmit the signals to IDC connectors. Surface mount components will be used for the capacitors and resistors and all components will be external to the active area. The PC

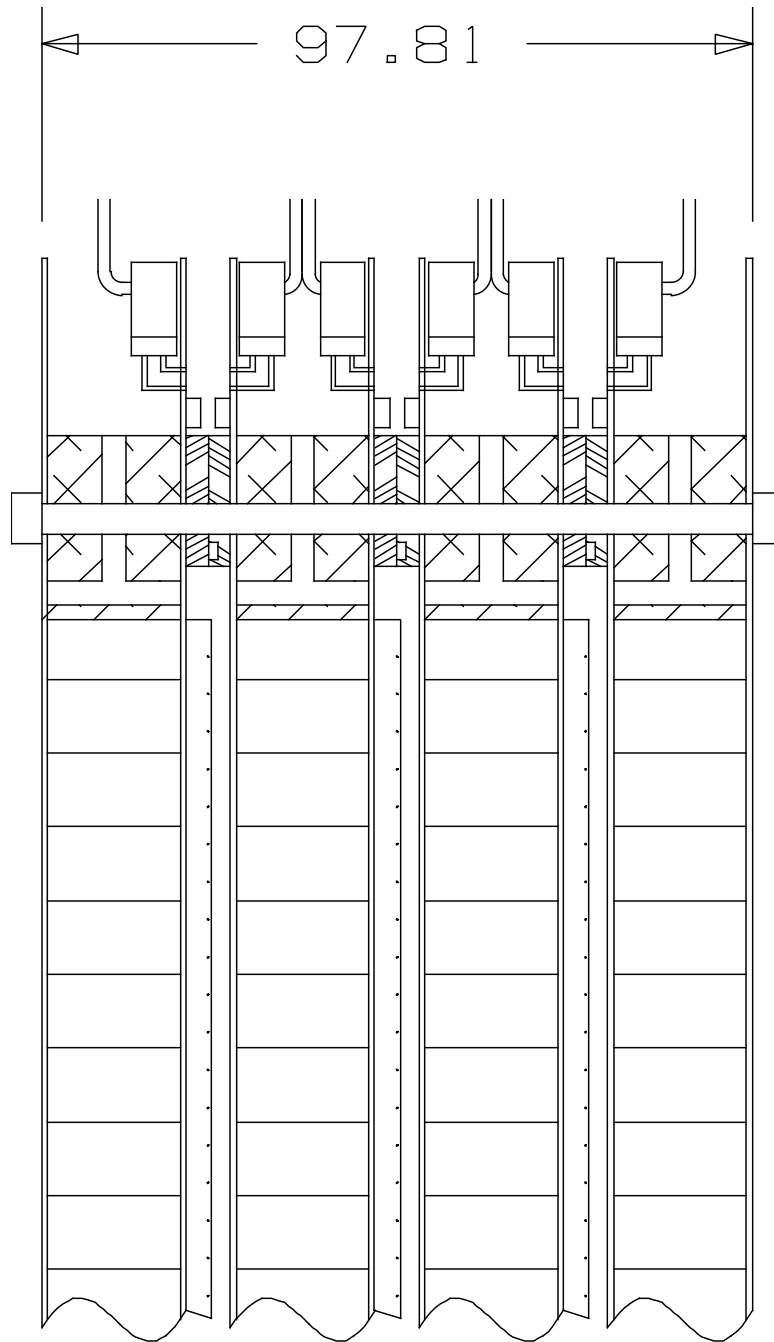


Figure 3.17: Station 1 Quadrant Cross Sectional View.

Component	density	cm/X0	total X0
Hexcel	0.14	485.7	0.0148
Epoxy Fiberglass	1.7	19.4	0.0325
1/2 oz copper	8.96	1.43	0.0095
frames	1.2	34.4	0.209

Table 3.3: Radiation Length of the Materials in the CSC Module.

board will extend beyond the edge of the quadrant by 1.5 cm. The anode wires will be held to the PC board by glue and solder and the positioning of the anode and field wires will be determined by fixturing. The anode supports at 22.5° , 45° , and 67.5° are delrin plastic with precision grooves. At 45° the wires will be glued and soldered while at 22.5° and 67.5° the wires will be glued.

Radiation Lengths

The radiation length of the materials in the CSC module is shown in Table 3.3. Each module is considered to be an active volume surrounded by frames that are inactive for detecting muons.

The total radiation length for the active area of the station 1 module is 5.7%, well within our budget. The frames at the octant boundaries are 20.9% from the lucite and 14.4% from the epoxy fiberglass for a total of 35.3% radiation length.

3.3.5 Station 3

The design for station 3 will be very similar to station 1 except that mechanically and electrically it will be in the shape of an octant. Additional support structure will be added to insure that the mechanical vibrations are kept below $25 \mu\text{m}$. A preliminary design is shown in Figure 3.18. The support in the middle of the chamber following a chord covers the same angular range as the support member for station 2. The cross section of the chamber is the same as station 1 and the radiation length thickness is 5.7%. All other aspects of the station 3 design is identical to station 1.

3.3.6 Station 2 Design

The station 2 chamber design is the most ambitious because of the severe radiation length restrictions. Foil cathode structures without honeycomb were required. In addition, to keep the acceptance as high as possible, the frame material surrounding the active area was kept to a minimum. We were able to minimize the frame material by overlapping the adjacent octants, a possibility for the CSC chambers because the active thickness of the CSC's were very small at 3 cm.

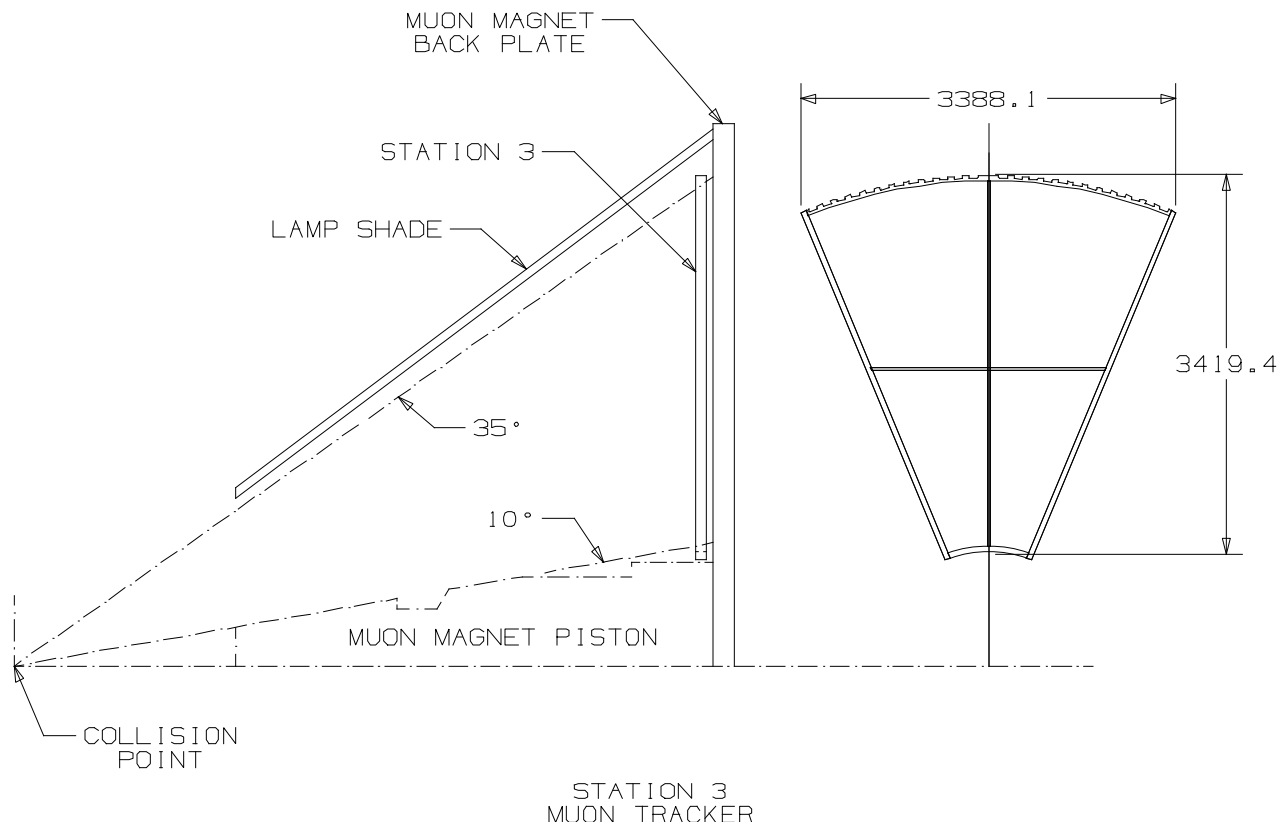


Figure 3.18: Preliminary Station 3 Design.

Station 2 design criteria

The station 2 assembly is shown in Figure 3.19. Each octant is a separate module that is tied together by a support structure that fastens all eight octants rigidly to the magnet. The octant designs are shown in Figure 3.20 and the cross section is shown in Figure 3.21. The module consists of thin frames for the cathode foils and wire planes stacked in a sandwich arrangement to form three CSC chambers. The thin frames are 7.5 cm wide and 3.43 mm thick consisting of a lamination of aluminum and epoxy fiberglass printed circuit board.

Similar to the station 1 design, the anode wires follow the chord in each 22.5° segment. The support spoke at 22.5° is attached to the top of the octant and free to slide at the bottom of the octant but is constrained to maintain the anode to cathode spacing to $50 \mu\text{m}$. The spoke is constructed to form a laminate with one piece grooved to define the anode to field wire spacing and the cathode to anode spacing. A top cap is glued to the grooved piece to define the cathode to cathode spacing and to aid in making the spoke a more rigid structure. We anticipate using a material such as delrin. During the wiring process the anode spoke will be fixtured to maintain its correct position.

The CSC chambers are held rigid and in proper shape and position by two aluminum support structures 3 cm thick and 7.5 cm wide on either side of the thin CSC chamber

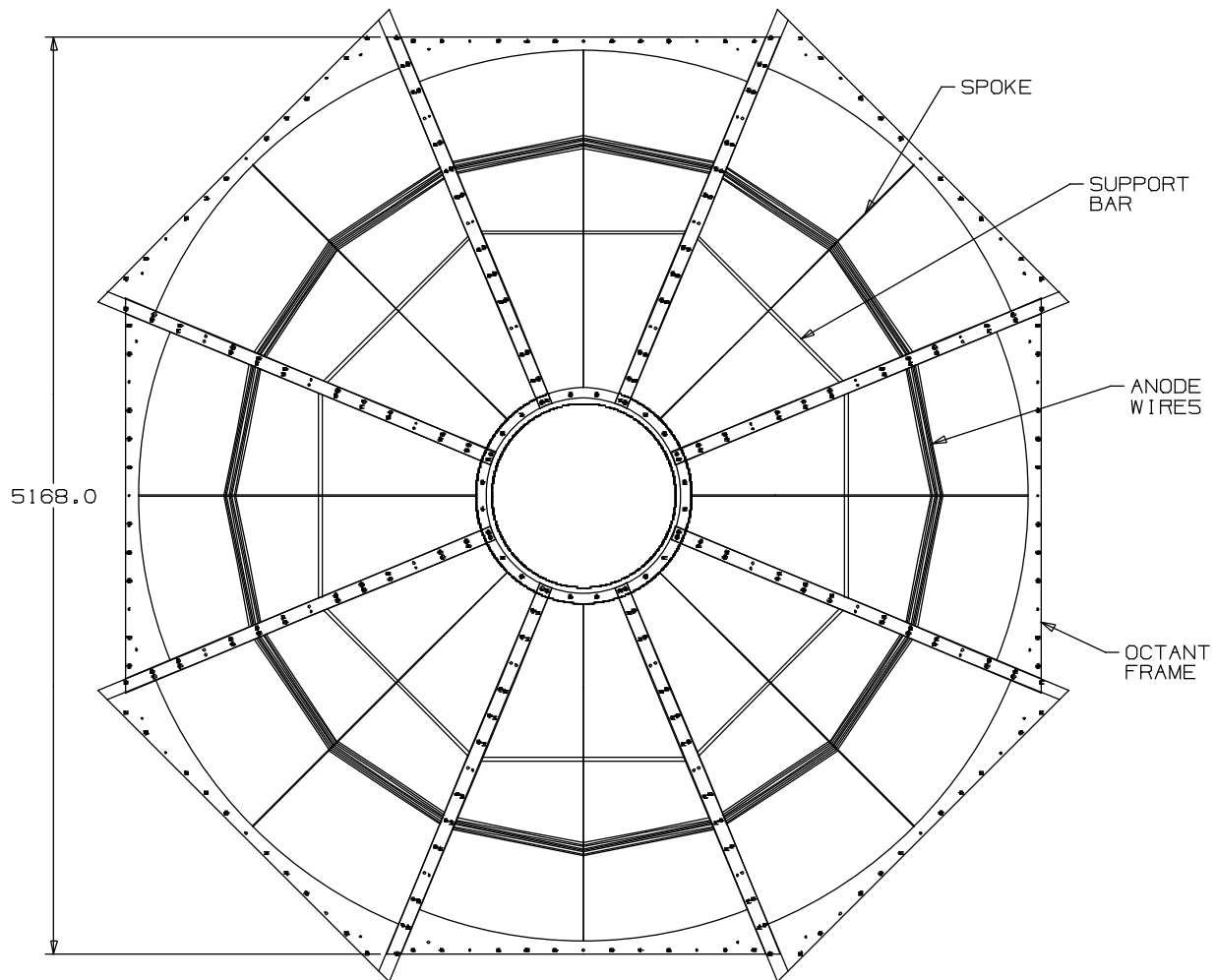


Figure 3.19: Station 2 View with All Quadrants.

frames. Nine alignment pins, three on each long edge, properly locate the thin frames during the assembly process. In addition, a single support structure is used during the etching and wiring process to insure proper location of the planes. A number of detailed calculations were performed to verify this design.

Finite Element Calculations

Finite element calculations have been performed on the frame and frame support structure to understand distortions due to the wire and foil loads. The support frame is shown in Figure 3.20.

The support frame is fabricated from aluminum although aluminum was not our first choice for the support frame. Calculations showed that a carbon composite frame would provide substantial structural rigidity without the need for a cross member support but at an extremely high cost. The cost was projected to be \$75k to \$100k for each octant support thereby completely exhausting the entire station 2 budget. Therefore, efforts were

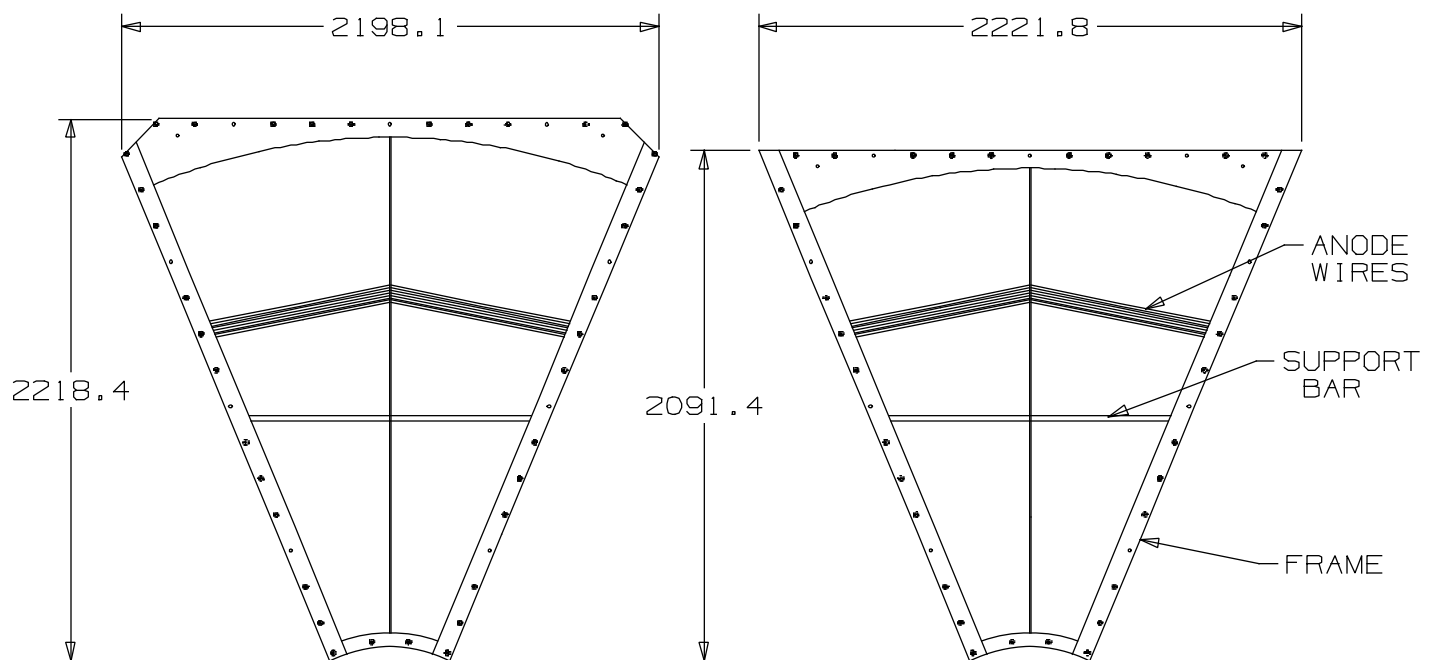


Figure 3.20: Station 2 Octant Design.

undertaken to investigate stainless steel and aluminum as the support frame. The finite element analysis result for aluminum is shown in Figure 3.22. The maximum displacement is 0.21 mm. Stainless steel of a comparable design had less displacement but at a price of three times the weight. We have established the baseline material to be aluminum.

Foil Etching

A technique to etch the thin kapton foil after the foil has been stretched on the support frame has been developed. This process involves electroetching the gold coated kapton with a probe tip attached to a low voltage source. The tip is mounted on a computer controlled $x - y$ table and positioned with an accuracy of $10 \mu\text{m}$ by a linear slide. A drawing of this setup is shown in Figure 3.23. The high resolution cathode will be prestressed during the etching process so that when the frame distortion occurs when all of the foils and wires are in place the cathode strips will be positioned correctly.

Since the kapton foil is available in widths up to only 1.5 m, station 2 must have foils that are glued together to get the required dimensions. A technique for gluing kapton foil was used successfully on cylindrical chambers for the MEGA experiment at LAMPF. They have not experienced any problems in three years. We have adapted this technique to the CSC foils and tests have been completed. Creep of kapton foils occurs in the first few days after stretching. Experience with full size stretched kapton foils have shown no unacceptable loss in foil tension due to creep over six months. Previously stretched 1 m foils have shown no loss in tension over four years. Generally, the foil will not sag under gravity load until the foil tension has almost fully relaxed. Our experience has shown this relaxation appears not to be a problem.

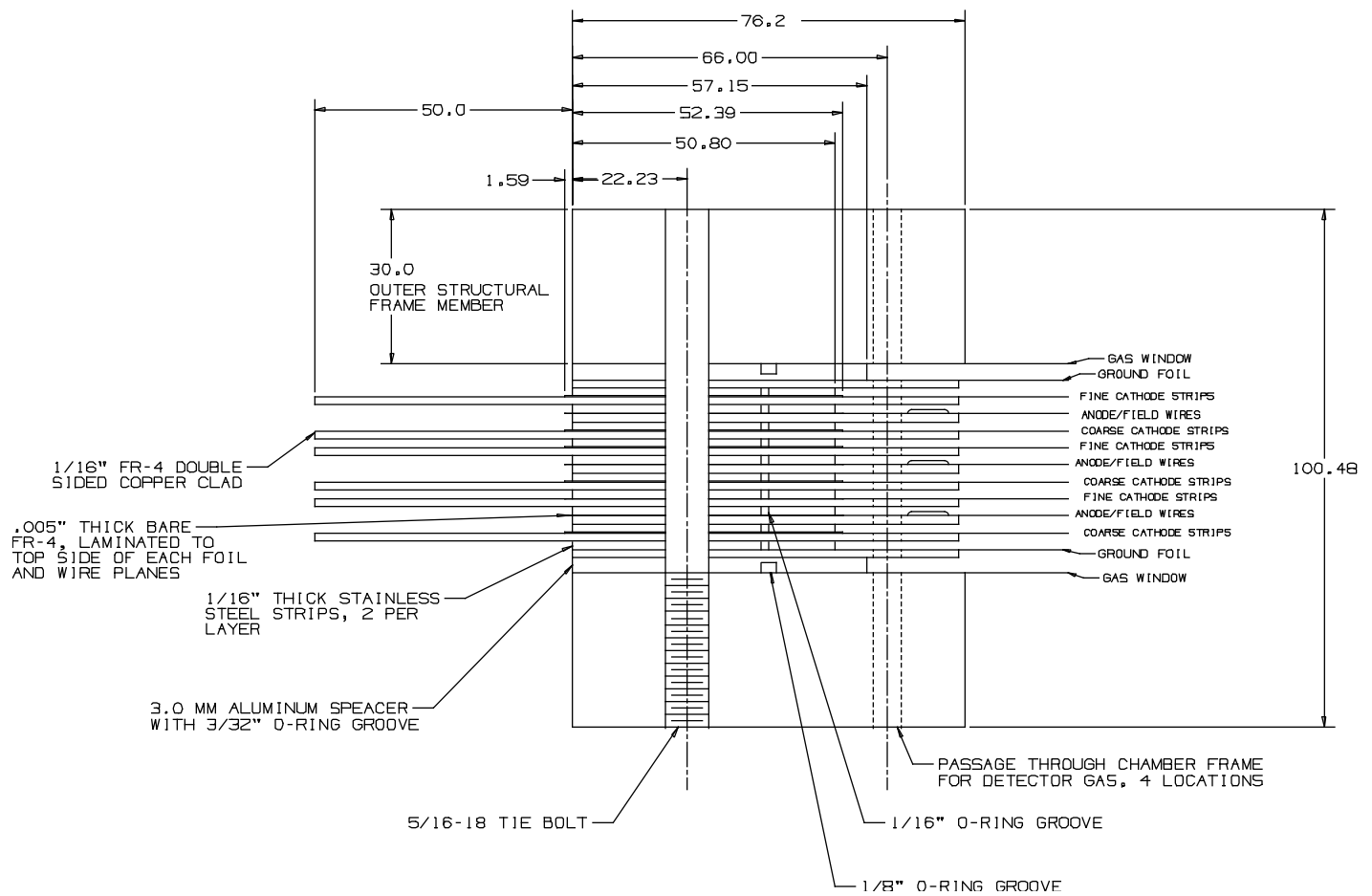


Figure 3.21: Station 2 Octant Cross Section View.

A foil etcher capable of etching the 2.5 m foils for station 2 has been assembled and tested. The positional accuracy of the optical encoders/lead screw has been measured to be $<50 \mu\text{m}$ for the ϕ coordinate but only $200 \mu\text{m}$ for the radial coordinate. The reasons for the discrepancy are not known at this time but suspected to be in part due to uncertain manufacturing tolerances of the space frame and uncertainties in the straightness of the leadscrews. We expect these issues to be resolved in the future.

Station 2 Mounting structure

The mounting structure of the station 2 octants in the magnet must hold the octants rigidly over a short period of time (0.5 hours) when the active alignment system is incapable of defining relative positions of station 2 with respect to stations 1 and 3. The mounting structure is shown in Figure 3.24 as viewed from the front and in Figure 3.25 as viewed from the side. It consists of stainless steel tubing, rectangular in shape, on the front and back of the octants with a truss structure at the top and mounting ring at the inner radius. The mounting ring does not come in contact with the piston. It is mounted to the magnet at

Lin DISP Lc=1

MECHANICAL PARAMETERS
AND CALCULATION RESULT

MATERIAL: AL 6061
RADIAL PRESSURE:
27.06 LBS/IN
TANGENTIAL PRESSURE:
21.80 LBS/IN

PHENIX PROJECT

PROTOTYPE MUON CHAMBER FRAME

STRUCTURE:

TOP BEAM WITH FLAT TOP
AND ARC SHAPE (R=2509MM)
BOTTOM BEAM: 50MMX30MM
SIDE BEAM: 75MMX30MM
SUPPORT BAR: 20MMX30MM

MAX. DEFL: 0.2148MM
AT MIDSPAN OF TOP FRAME

MAX. DEFL: 0.0811MM
AT MIDSPAN OF BOTTOM FRAME

ABSOLUTE DISPLACEMENT AT
BOTH ENDS OF SUPPORT BAR IS
UX=0.0675MM
UY=0.0277MM

FILE NAME: PHENIX_A
DATE: NOV. 6 1994
LANL/ESA-DE Z.CHEN

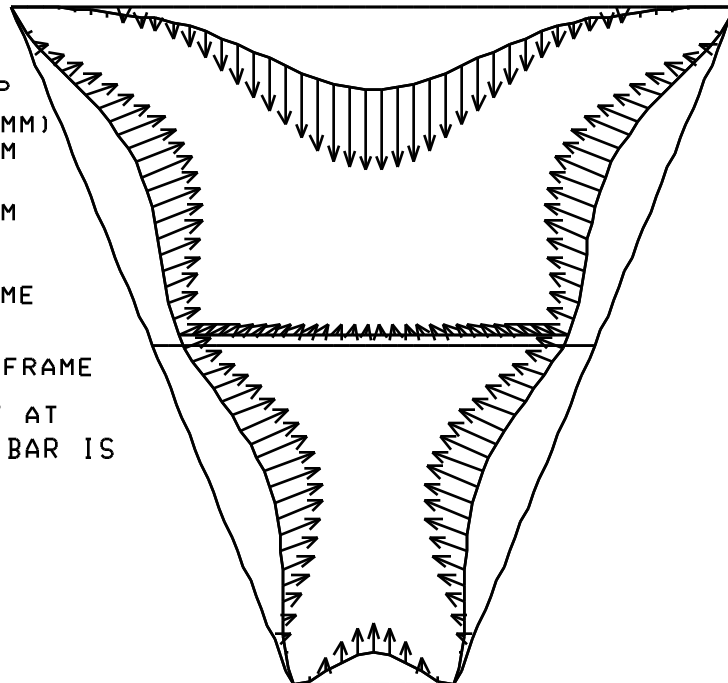


Figure 3.22: Results of finite element analysis of CSC frame.

four points at the bottom of the magnet and at four points to the teacup. The mounting structure, and therefore the chamber octants, are electrically isolated from the magnet.

The mounting structure design criteria was to have vibrations from natural frequencies less than 25 μm . Static distortions due to gravity, thermal, and magnetic effects were minimized but allowed. The active alignment system is designed to account for static distortions.

An analytic calculation of the weight and maximum distortions of the space frame were carried out for a variety of tubing sizes and for stainless steel and aluminum. The results pointed to stainless steel with a cross section of 7.5 cm \times 5.0 cm \times 0.3175 wall as an acceptable material. The maximum displacement under gravity load is 0.399 mm. We were constrained to have the frame and chambers fit in the piston slot of 35 cm. This design is 30 cm. A finite element analysis of the mounting structure with the octants in place has been carried out with the program COSMOS/M to define the natural frequencies of vibration and the static distortions due to gravity. The static distortions due to gravity are shown in Figure 3.26. The maximum distortions were 0.362 mm. The frequency results are shown in Table 3.4.

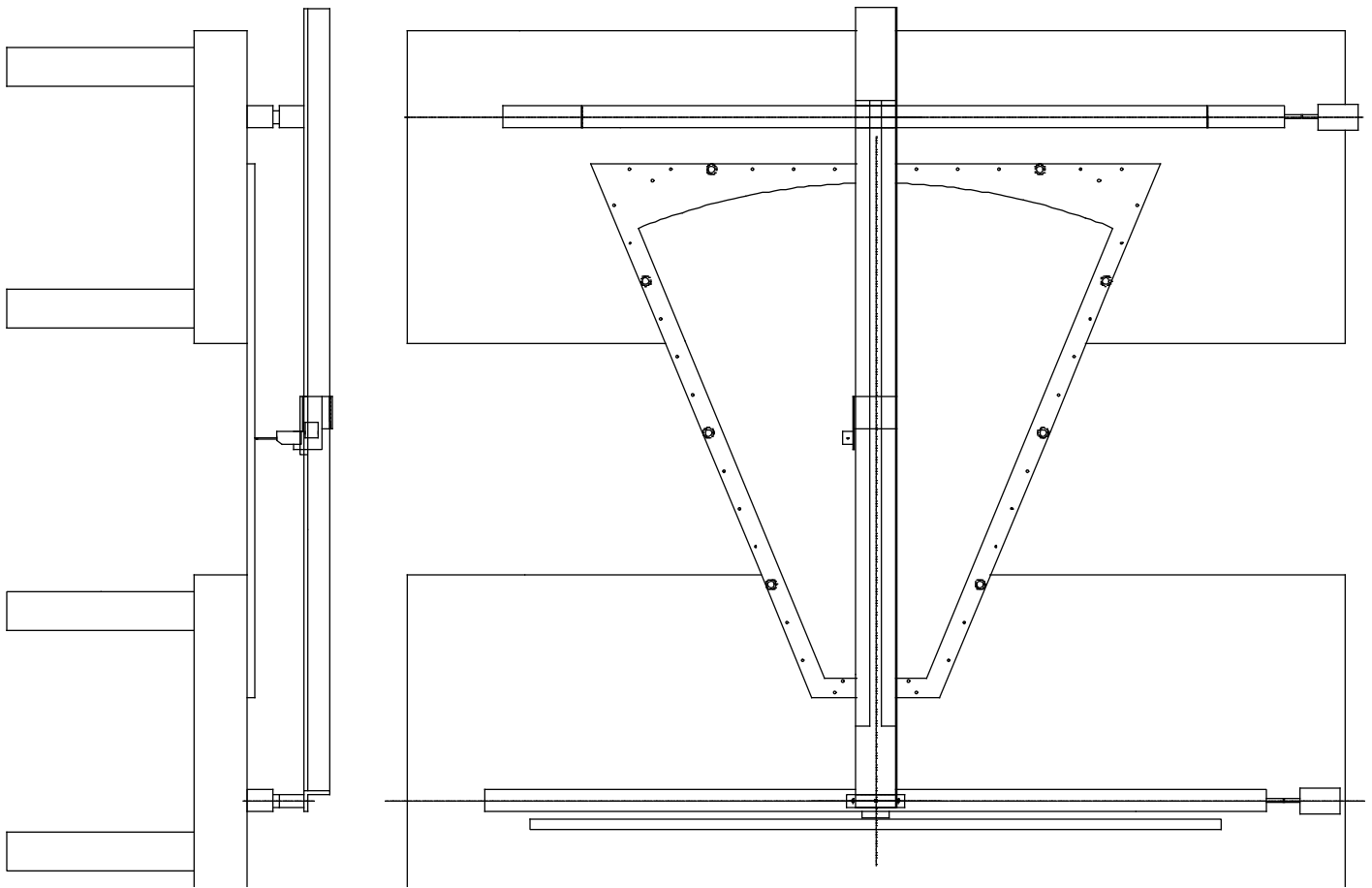


Figure 3.23: Setup for etching stretched cathode foils on frames.

Mode #	1	2	3	4	5
Freq (Hz)	7.9	12.4	16.6	25.5	27.9

Table 3.4: Frequency Results.

The first three modes correspond to displacements in the z direction. The fourth and fifth modes are displacements in the ϕ and r direction and therefore are of greatest concern. To determine the maximum displacements expected from these and all frequency modes, a dynamic analysis was done using the finite element analysis program ABAQUS. The excitation to the station 2 structure was based on measurements made at four laboratory facilities [32]. It was considered roughly typical of what one would expect at the PHENIX facility. No site specific ground motion measurements have been made and the excitation function does not represent vibrational forces that could be generated by equipment (e.g. vacuum pumps, etc.) attached directly to the PHENIX detector. In the calculation the sagitta error was determined under the assumption that stations 1 and 3 are fixed. The frequency

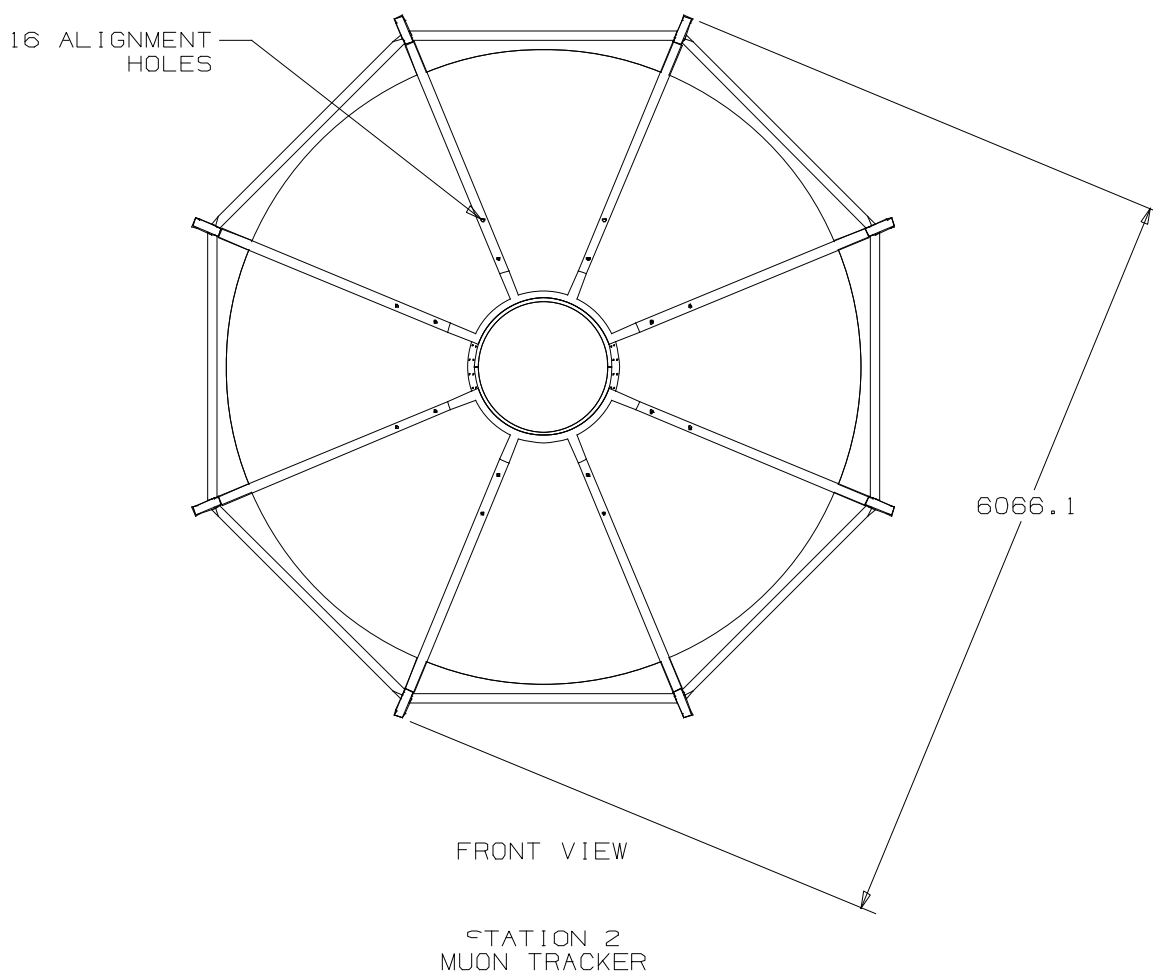


Figure 3.24: Station 2 Mounting Structure - Front view.

analysis for the magnet suggests that motion in the x and y coordinates of stations 1 and 3 is negligible owing to the large massive backplate and teacup to which they are attached. The response power spectral density for x and y components of the sagitta at one particular node (node 40) is shown in Figure 3.27. This node corresponds to the vertical frame of the octant at the top of the magnet at the 12:00 o'clock location. The peaks in the spectrum correspond to the frequencies where the coupling is the largest. The z component is added for reference but does not effect the sagitta. The largest x component of the sagitta rms response is $0.17 \mu\text{m}$ at node 40. The largest y component of the rms response is $0.11 \mu\text{m}$ at node 6. Node 6 corresponds to the horizontal frame of the octant on the side of the magnet at the 3:00 o'clock location. The largest z component of the rms motion relative to the base is $1.1 \mu\text{m}$ at node 260. Node 260 corresponds to the horizontal member at the top of the octant at the top of the magnet. The modes that cause the greatest x direction motion are modes 4 and 5. The mode giving the greatest z direction motion is mode 1. The modes that give the greatest y direction motion are modes 1, 4, and 5. Mode 1 involves primarily z direction motion, but some coupling occurs to y . The results show that the largest sagittal

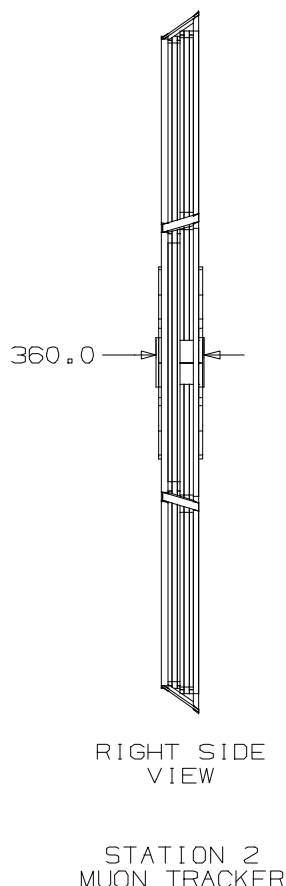


Figure 3.25: Station 2 Mounting Structure - Side View.

error introduced because of ground motion is approximately $0.2 \mu\text{m}$. We expect to measure the ground motion at the PHENIX site and redo this calculation but the results indicate that we have two orders of magnitude margin before ground motion becomes a problem.

3.3.7 Global Tracking System Design Issues

To satisfy the physics requirements of the tracking system, the assembly plan for the detectors, the alignment requirements, and the calibration requirements have to be addressed.

Assembly Plan for the Detectors

The CSC detectors will be assembled precisely primarily by using highly accurate positioning jigs. For stations 1 and 3, the alignment of the high resolution cathode will be accomplished by using a drilling jig to locate the hole for the alignment pins to the cathode strips. The expected accuracy will be the accuracy of the photo graphic etching process, $25 \mu\text{m}$. For station 2, the precision of the strips to the alignment pins will be dominated by the accuracy of the etching table, $25 \mu\text{m}$. In addition, for station 2 we will etch a ground foil with alignment fiducials that are visible external to the chamber. The location of these fiducials

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Lin DISP Lc=1
SPACE FRAME AT STATION 2
WITH MUON CHAMBER STIFFNESS ON
LOADING: GRAVITY
MATERIAL: ST_ST
MAX. DISP: 0.363 MM

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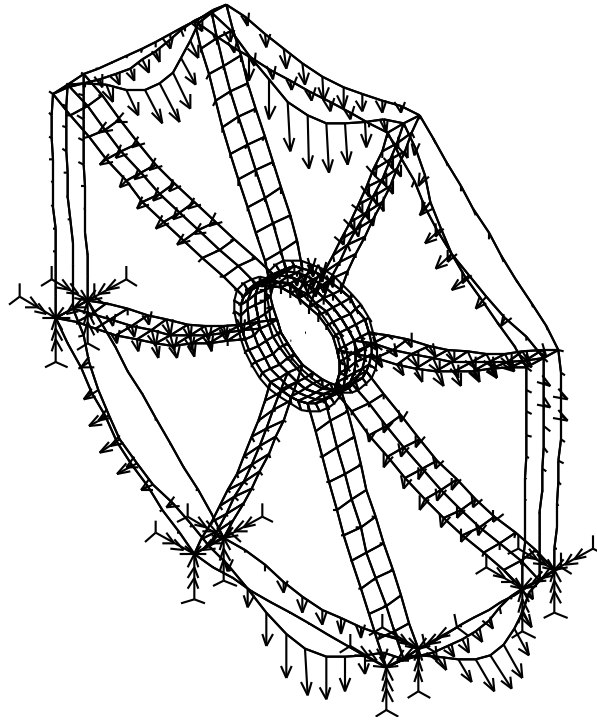


Figure 3.26: Static Distortions of station 2 Due to Gravity.

will be determined during the staging of station 2 in the chamber fabrication building prior to installation in the magnet and again when installed in the magnet. The active alignment system will provide the accurate survey in the magnet and will determine subsequent motion due to thermal effects and field effects.

Alignment Plan for the Detectors

It is our goal for the tracking chambers to define the contribution to the station resolution to be as small as possible consistent with expected alignment and assembly errors. A detailed report on the contributions the errors and alignment and assembly is given Ref. [33]. In summary, we expect the assembly errors to be $25 \mu\text{m}$, the alignment errors to be $10 \mu\text{m}$, and the calibration of the alignment system to be $25 \mu\text{m}$. The combined errors with $100 \mu\text{m}$ intrinsic chamber resolution gives a realizable chamber resolution of $110 \mu\text{m}$. To achieve a realizable chamber resolution of $100 \mu\text{m}$, an intrinsic resolution of $82 \mu\text{m}$ is required. Either resolution is acceptable. Every technique to improve the chamber resolution and the other combined errors will be implemented. The single largest contributor to the resolution is the chamber resolution and this is primarily due to electronic noise and calibration. Our greatest effort will be directed at reducing the electronic noise and insuring that the calibration scheme between channels is very accurate.

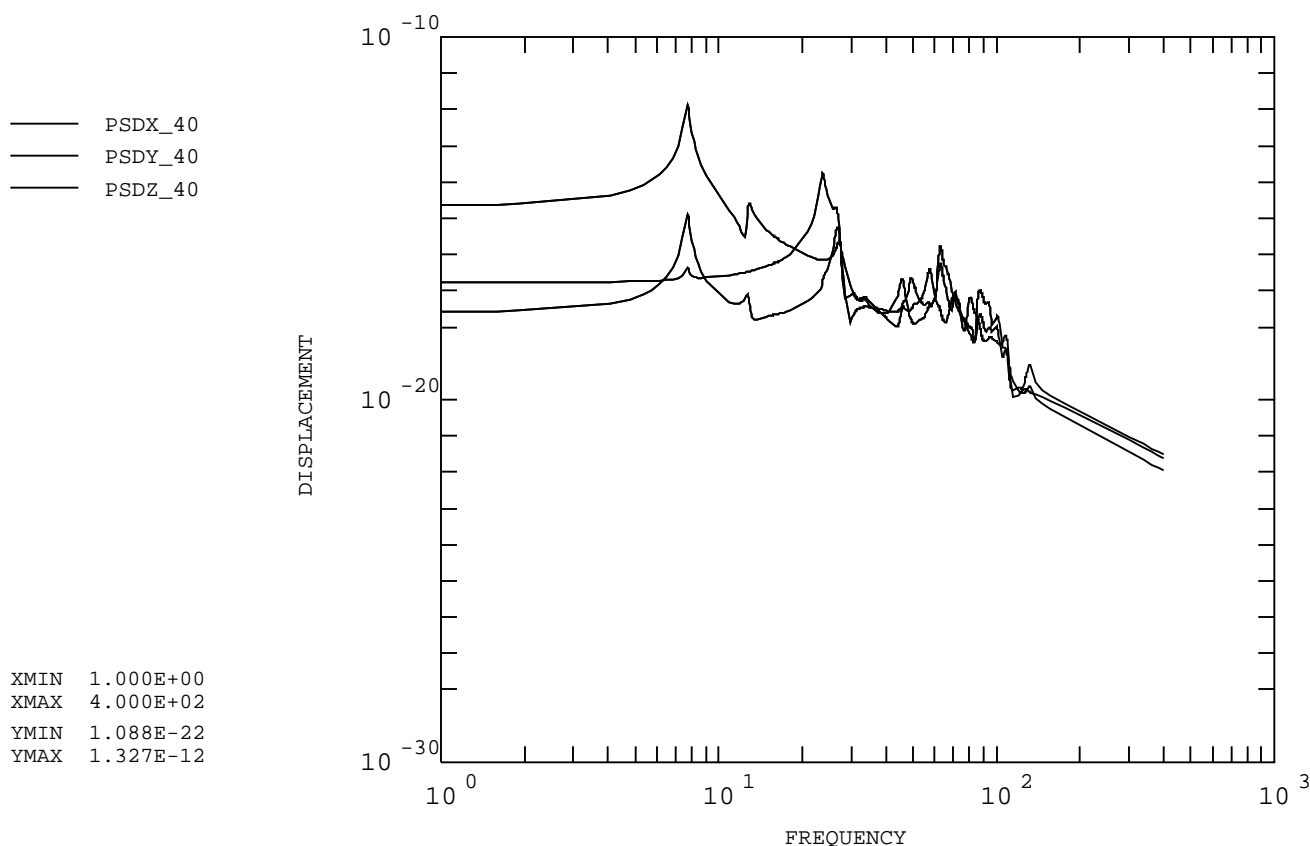


Figure 3.27: Power Spectral Density Sagitta Response for x and y directions at node 40.

3.3.8 CSC R&D Prototypes and Results

The major issues for the CSC chambers involve the mechanical construction of the chambers, the electrostatic stability, and resolution.

Chamber prototypes

The prototype PHENIX chambers will all be full scale prototypes to streamline the R&D. The station 2 prototype consists of two CSC chambers and two drift chambers. The drift chambers are intended to be tested as a backup for the CSC's but will not be discussed here. The structural frame is the same as shown in Figure 3.20 and the cathode foils and anode planes are the same as shown in Figure 3.21. The support frames were built and found to be within the $25\ \mu\text{m}$ tolerance for the alignment pins. Full scale foil frames were constructed and stretched with gold and aluminum coated kapton foils. The technique of using laminated aluminum and epoxy fiberglass frames with glued joints have been successfully tested. The seamed $25\ \mu\text{m}$ aluminumized kapton foils were stretched and etched without problems. Seaming and handling $12.5\ \mu\text{m}$ gold foils was more difficult so we have chosen to go with $25\ \mu\text{m}$ gold foils. We have successfully transferred the alignment pin holes to the thin foil frames by using drill guides in the support frame and the support frame as the template. We

have yet to test the anode support spoke. The maximum unsupported anode wire length in station 2 is approximately 1.15 m. We have just recently completed the construction of drift chambers for experiment E866 at Fermilab with the identical cell design as that for station 2 and with an unsupported anode wire length of 1.25 m. The chambers operated reliably and with a wide efficiency plateau.

Station 1 prototype chamber is full scale with two CSC chamber gaps. Two hexcel panels are being constructed by a commercial vendor and one hexcel panel is being constructed “in house”. The commercial vendor has found no problems with the 150 μm flatness tolerance but the estimated cost of the commercial panels is twice that of the “in house” panels.

Station 3 prototype chamber is now under design and will be a full scale construction.

Resolutions

The resolution of 100 μm has been shown to be readily achievable if attention is paid to the assembly requirements, the electronic noise level, and the calibration.

We believe that our assembly procedures are adequate. The electronic noise level is a concern since the strip capacitance is very large. However, a recent amplifier design [34] has been fabricated that meets or exceeds all of our requirements. The measured noise level at the maximum capacitance of our chamber strips exceeds our noise level requirement by 2.5 and meets our linearity requirements. The remaining concerns are calibration and assembly. The calibration can be accomplished by injection of a precise charge into the amplifier by either a capacitor incorporated into the front end amplifier or an external capacitor etched into the PC board. Incorporating a capacitor into the front end amplifier is estimated to be accurate to 0.5%. Etching a capacitor into the PC board depends on the accuracy of the etching process but we expect to be able to do this to the 0.5% level. We need 1% accuracy so either technique will work. However, we will prototype both since previous experience by these authors has been the most difficult aspect of achieving good resolutions has been uncertainties in the calibration.

A CSC chamber with strips lengths of 1 m and strip widths of 5 mm was tested and found to give resolutions of 110 μm . A plot of the resolution as a function of voltage is shown in Figure 3.28. The limit was determined to be due to limitations in electronic noise levels and dynamic range of the amplifier. We expect to improve upon this by using a better quality amplifier.

Efficiency Plateau

The 1 m CSC prototype exhibited a good efficiency plateau. The plateau is shown in Figure 3.29. The cell design for this chamber is identical to the cell in the PHENIX CSC chambers.

Mechanical R&D

The major concerns for the mechanical design are achieving 150 μm flatness of the honeycomb panels, achieving good electrostatic stability and therefore good efficiency plateau for

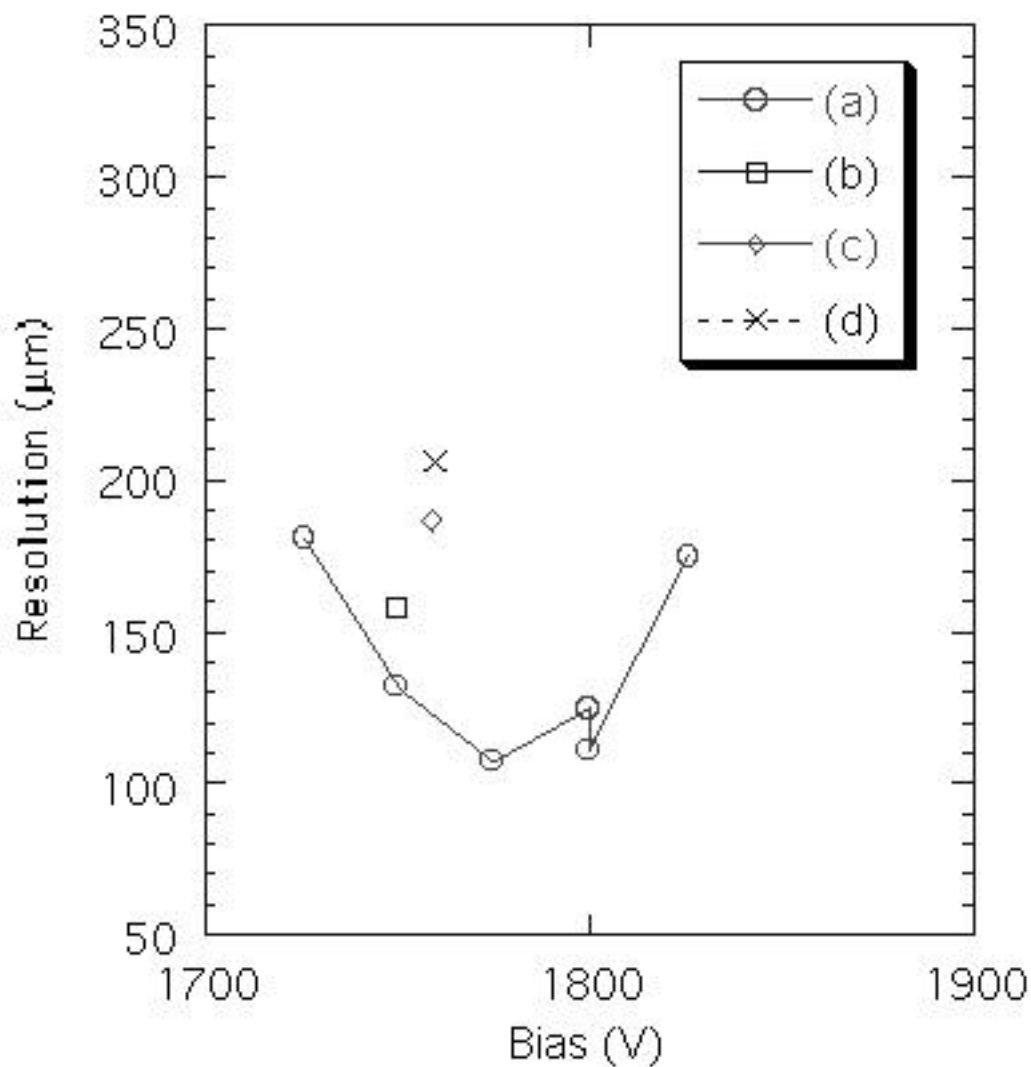


Figure 3.28: Resolutions for a Prototype CSC with different cathode structures.

our choice of mechanical design and unsupported anode wire length, verifying the anode support spoke for the station 2 design, and generally testing that the assembly procedures are adequate to maintaining the resolution requirements.

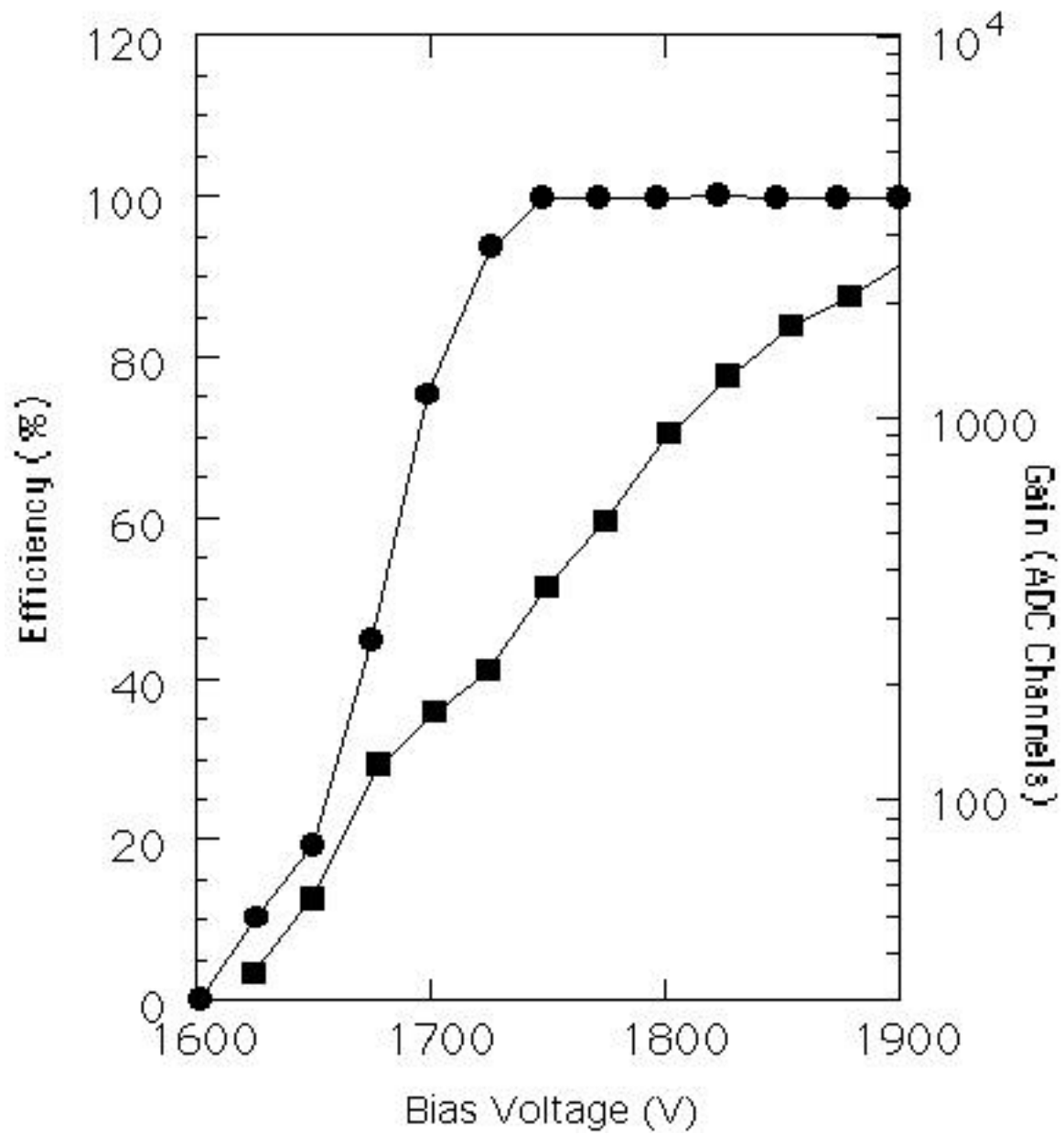


Figure 3.29: High Voltage Plateau for 1 m Prorotype CSC.

3.4 Muon Identifier Detectors

3.4.1 Detector Design Optimization

Steel Thickness

The irreducible μ/π ratio due to weak decay is approximately 1×10^{-3} (primarily determined by the proximity of the nosecone to the vertex). We set a detector design criteria to 1/4 of this, namely 2.5×10^{-4} , for a pion from the vertex to be misidentified as a muon (forming part of a dimuon). The factor of 1/4 provides over an order of magnitude suppression for the *pair* background. Thus, the irreducible background of muons reaching the muon identifier, as opposed to the muon identifier design and the algorithms used to reject the larger hadron background, will set the ultimate physics background level. Of this required net μ/π separation, approximately 10^{-2} is provided by the presence of the absorber preceding the muon identifier which filters out pions. This leaves 3% as the maximum tolerable fraction of the charged pions which may subsequently be misidentified as muons.

The punchthrough probability versus depth of steel for muons of various momenta are plotted in [35] and [36]. In order to set the punchthrough probability for muons of up to 6.5 GeV/c to be 3% or less, a total steel depth of 110 cm (6.6 hadronic interaction lengths) is required beyond the nosecone and central magnet. Subtracting the thickness of the muon magnet backplate, a total depth of 80 cm of steel is required in the muon identifier itself.

A muon at the vertex must have a mean energy of at least 1.9 GeV to reach the muon identifier system. The mean minimum original energy to penetrate completely through the muon identifier is 3.0 GeV [37].

Number of Layers

Ideally, the 80 cm of steel is to be distributed in a sandwich interleaved with detector layers. The detectors themselves require a non-negligible amount of space and the total available space in z is quite limited in the PHENIX detector hall (2.7 m). Also, it is desirable to have the early absorber layers be divided more finely to increase the acceptance for ϕ meson detection. Segmentation also improves the measurement of the trajectory in the identifier. The segmentation chosen is a total of five steel absorbers after the 30 cm thick muon magnet backplate of the north arm of thicknesses 10, 10, 20, 20, and 20 cm. The 6 gaps created by the absorbers are instrumented with the muon identifier panels. The muon identifier for the south arm is identical to that for the north arm (although the muon magnet backplate is only 20 cm thick) and at the same distance from the interaction vertex. Maintaining the same design for both arms helped minimize design costs.

Readout Pitch Requirements

Beam gas studies show that for low polar angles (20° or less), an effective segmentation into logical pads of approximately 13 by 13 cm is required to suppress false roads for tracks in the muon identifier. Rather than develop Iarocci tubes of greater width or use 13 cm wide external strips, the most cost effective solution is simply to use standard Iarocci tubes of

width 8.5 cm with all eight internal wires ganged together. This provides a readout pitch of 8.5 cm along both the x and y directions, thus providing effective 8.5 cm square hodoscopic cells (upon forming the appropriate ANDs). This most cost effective solution exceeds our requirements. This segmentation is fine enough to provide sufficient granularity for matching roads in the identifier to tracks in the muon tracker unambiguously with anticipated occupancies.

3.4.2 Introduction to Iarocci Tubes

What They Are

We use the term Iarocci tubes to refer to planar drift tubes consisting of 100- μ m gold-coated CuBe anode wires at the center of long channels of a graphite-coated plastic cathode, with a gas such as CF₄-isobutane. This same physical detector when operated at higher voltage is a conventional limited streamer tube [38]. We operate them in the proportional mode to increase longevity. They are proportional tubes operated at a voltage such that only a fraction of the signals are not proportional mode pulses.

How They Perform

We apply 4200 V to the anode wires. Incident minimum ionizing particles ionize the gas and create an avalanche. Using a high voltage decoupling capacitor, we read out the signal on the anode wire. The drift time interval (for arrival of 90% of pulses) in the tube is approximately 100 ns for a 25:75 Ar-isobutane gas mixture. This is reduced to only 45 ns if a 50:50 CF₄-isobutane mixture is used.

Why They Were Chosen

Iarocci tubes were chosen because they satisfy all of the detector performance criteria discussed above. They have proven reliability and longevity, compactness, low cost, and are readily available from commercial vendors [39]. Such tubes can be used economically to tile large areas. They have robust wires and seals. They avoid the problems of metal-plastic transitions present at the endcaps for aluminum proportional tubes.

Tubes with 9 mm by 9 mm channels satisfy the count-rate and position localization requirements, but must be staggered by half a cell and ORed in pairs (discussed below) and/or operated with a fast gas (discussed below) to meet the timing requirements of the Local Level 1 (LVL-1) trigger system.

3.4.3 Design of Muon Identifier Detectors

Definition of Terms

A	B	C
D		E
F	G	H

Table 3.5: Panel arrangement viewed from origin.

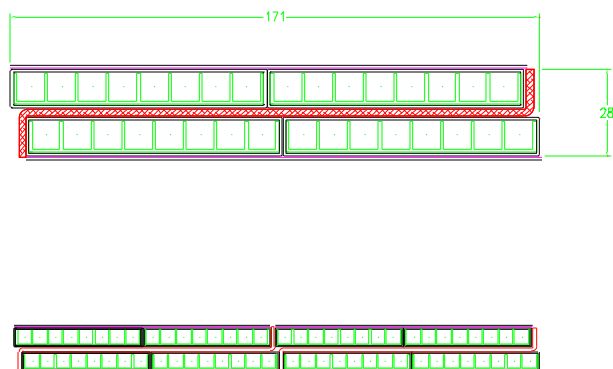


Figure 3.30: Four-pack with its stagger backbone.

A Channel: An individual anode wire and the immediately surrounding gas volume and graphite-coated walls are referred to as one channel of an Iarocci tube. Each wire is held at the center of its channel by means of plastic wire spacers positioned every 50 cm along the tube.

A Comb Profile: The three sides of each of the eight channels in an Iarocci tube are provided by the graphite-coated comb profile.

A Jacket: The (outer) jacket is the surrounding rectangular PVC housing that encloses the comb profile to create the gas-tight box for the eight channels.

A Tube: The term Iarocci tube refers to the ready-to-go detector element consisting of eight channels. Its principal components are the comb profile, the wire, wire spacers, outer jacket, and plastic endcaps.

A Four-Pack: In our muon identifier design, an s-shaped aluminum rail is used to hold two pairs of tubes back-to-back and staggered by half a channel. The rail is referred to as

Panels	Horizontal	Length	Vertical	Length
A, C	98H	5765 mm	128V	4000 mm
F, H	92H	5765 mm	128V	3820 mm
D, E	48H	5265 mm	124V	2286 mm
B	98H	1369 mm	36V	4000 mm
G	92H	1369 mm	36V	3820 mm

Table 3.6: Number and lengths of Iarocci tubes per panel.

Quantity	Value
panels/gap	8
panels/arm	48
tubes/gap	1498
tubes/arm	8988
area/gap	13.1 m \times 10.7 m
LVL-1 trigger roads/arm	749

Table 3.7: Muon identifier specifications.

the stagger backbone and is shown in Figure 3.30. The group of four tubes glued to the rail is called a four-pack. The rail provides stability, alignment, and prevents bowing.

A Panel: Groups of four-packs oriented both horizontally and vertically are held inside an aluminum box. Half are oriented horizontally and half are oriented vertically so that both projections are measured. This total detector element is called a muon identifier panel. There are eight such panels per gap labelled A through H as shown in Table 3.5. Note the square opening in the center for the beam pipe.

Layout of Panels In the Detector

The number and lengths of tubes inside every panel are listed in Table 3.6. Overall specifications are listed in Table 3.7. Adjacent panels overlap along their edges to eliminate the creation of any deadspace by the panel frames. The panel frames are shown in Figure 3.31. The smaller panels (B, D, E, G) lie in one plane which is 10 cm farther from the vertex than the plane of the larger panels (A, C, F, H).

Figure 3.32 shows the active area of the (four-packs) of Iarocci tubes oriented vertically and horizontally in the eight detector panels of one gap. The two vertical cracks between the three “monoliths” of steel, which make up a single absorber plate, are sketched. The 7 cm thick aluminum panel frames and the top support rail are shown. The figure shows where 10° and 37° cones would reach for both detector gaps 1 and 6. As is apparent, the

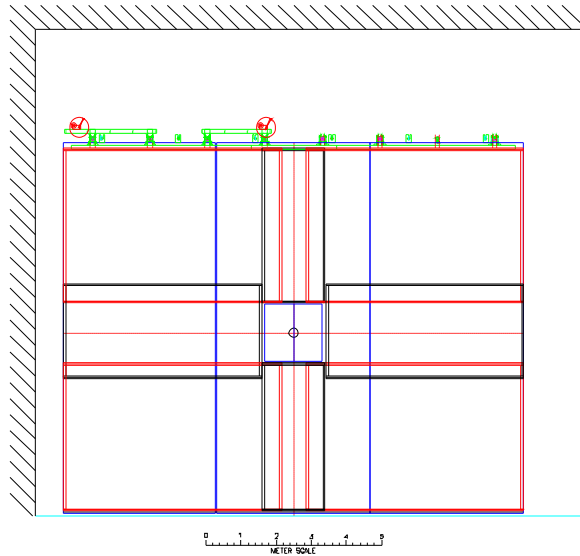


Figure 3.31: Muon identifier frame.

acceptance reaches down to 10° in the first plane (and even farther in subsequent planes) except immediately at the four corners of the square beam hole. The elevation view of the north muon arm is shown in Figure 3.33.

3.4.4 Muon Identifier Design System Issues

Assembly Plan for Muon Identifier

The tubes will be assembled off-site (in China and/or by a commercial vendor) and shipped to BNL after having satisfactorily passed quality assurance tests for both gas leakage and high voltage. The tubes will be installed into panels assembled at a staging area at BNL. The assembled and tested muon identifier panels will be transported by truck a short distance to the Main Facility Hall. Each will be attached to a special exterior handling frame that can be hoisted using the hall crane. The panels will be inserted in appropriate sequence onto the rails of each gap (except the bottom 3 panels which are supported from below) in order to achieve the desired arrangement. In particular, panel “D” will be installed by sliding it across and over the beam pipe along the top rail. It then will be lowered into final position by means of a winch and support rods attached to the top rail. At least part of the shield wall must be removed for panel installation and removal; however, the beam vacuum does not need to be broken at any stage.

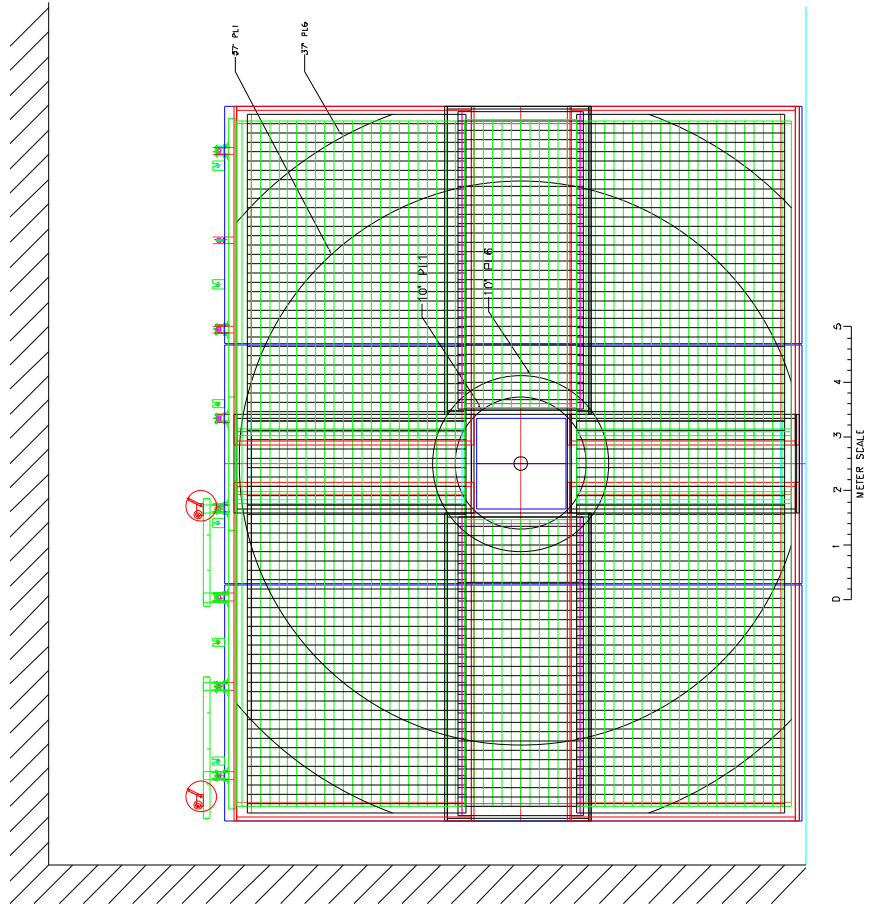


Figure 3.32: Muon identifier panel scheme.

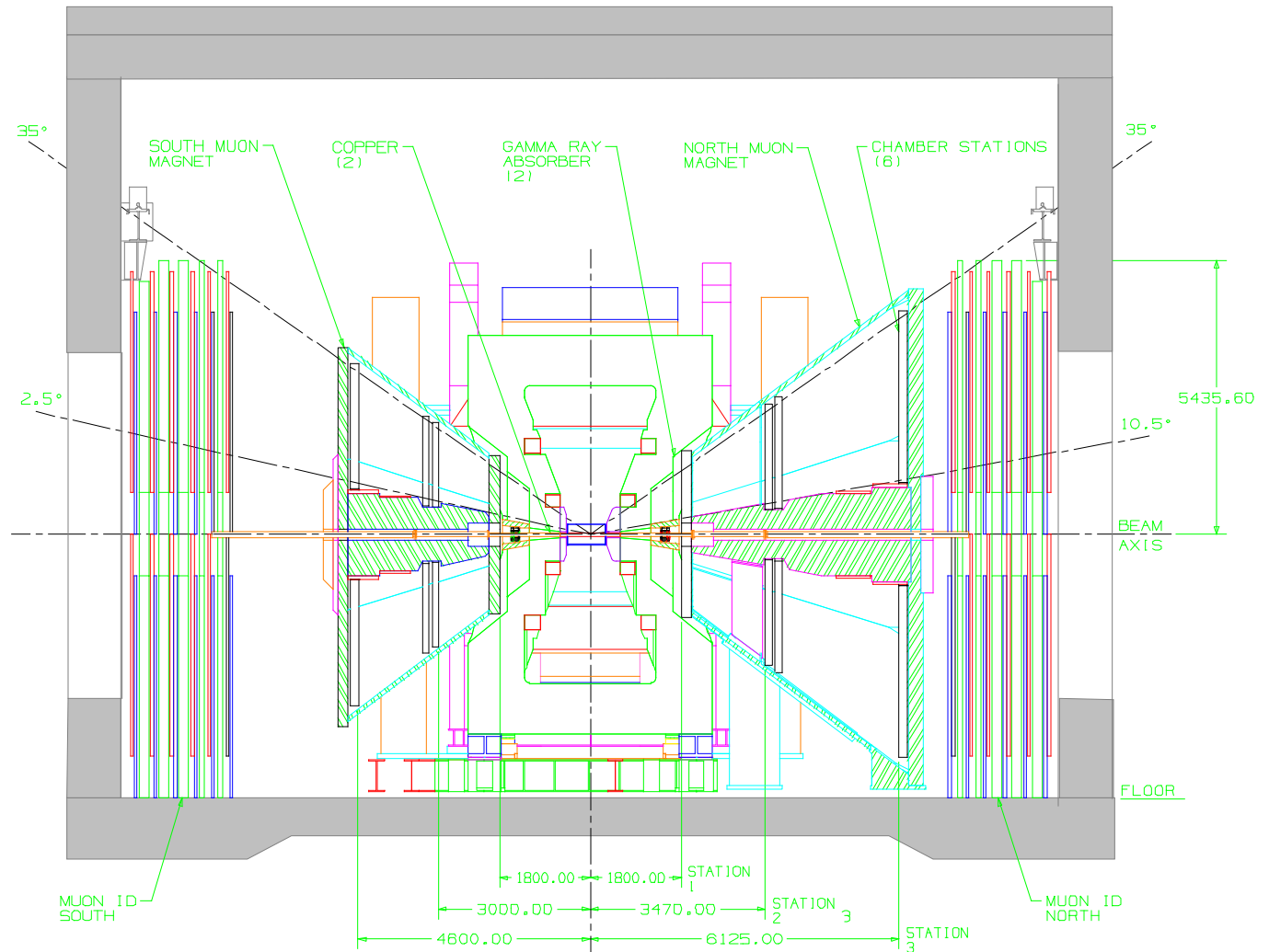


Figure 3.33: Elevation view of Muon Arms.

Timing Requirements for Triggering

We require a drift time interval shorter than the time between beam collisions (106 ns) in order to avoid dead time. Shorter drift times reduce the backgrounds associated with the arrival of out-of-time hits such as upstream beam gas events as well as relaxing constraints on the electronics that feeds the LVL-1 trigger system (by providing margin for variations among tubes, cable delay variation, open/closing of latch, drift from cell edge, and time-of-flight from vertex).

Alignment Requirements

The muon identification system has much less stringent alignment to the vertex requirements than the muon tracker since the typical pointing resolution of the system is measured in centimeters and the muon identifier's main function is to work with the muon tracking system. However, since the muon identification system is used in the trigger to find roads which point back to the vertex, there must be some relative alignment of the identifier panels with the interaction point.

The multiple scattering in the absorber material in front of the muon tracking stations limits the angular resolution of a 10 GeV/c track at the vertex to approximately 3 mrad or more. If this error is projected from the first tracking station to the muon identifier, the position resolution is approximately $(5.4 \text{ m}) \times (3 \times 10^{-3}) = 1.6 \text{ cm}$. An additional 1 to 2 mm or error occurs at each of the absorber layers in the muon identifier. Therefore, the muon identifier planes need not be surveyed with respect to the vertex to better than approximately $\pm 4 \text{ mm}$ in the x and y directions in order to have alignment errors be negligible compared to multiple scattering errors. It is only required that the z positions be measured after they are installed, and this measurement only has to be known to within a few centimeters relative to the vertex.

To Tracking System The effective logical pad sizes represented by the transverse tubes of the muon identifier system are 8.5 cm by 8.5 cm. The position resolution along a muon identifier road cannot be expected to be much better than $8.5/\sqrt{12} = 2.5 \text{ cm}$. If several measurements of resolution 2.5 cm are used, then a projection error of somewhat less than 2.5 cm might be expected. Multiple scattering due to the 30 cm thick muon magnet backplate causes a projection error of 2 mm for a 10 GeV/c track in the tracker to the beginning of a road in the first plane of the muon identifier. We require alignment of the muon identifier planes with station 3 of the tracking system to $\pm 1 \text{ cm}$ in the x and y directions.

Internal The internal alignment requirements of the muon identification system are determined by its ability to form roads when the tubes are misaligned by a given amount. The individual tubes are 8.5 cm wide. We require alignment errors to be a negligible contribution to the projection error of a road from one plane to the subsequent planes in the identifier system. We require a 5 mm panel to panel alignment in order for road finding in the muon identifier to be limited by multiple scattering. Within a four-pack, we require and can easily achieve a stagger by half a channel to within 1 mm.

Method of Alignment

The tubes are installed and surveyed within each panel relative to external fiducial marks on each panel to within an accuracy of 2 mm. The external fiducials of the panels can be viewed by conventional survey theodolites located along the west or east sides of the respective arm and down the the main square hole which contains the beam pipe. Using standard survey techniques it is straightforward to measure the absolute positions of the muon identifier detectors to within 3 mm or less which often exceeds our requirements. Panel positions

can be adjusted by sliding transversely and by means of lowering rods; however, in most instances we just need to *know* the as-built as-installed positions. Further details concerning muon identifier alignment criteria are discussed in [33].

Calibrations for Muon Identifier

We will record calibration data to obtain samples of cosmic muons, samples of beam gas muons, pedestals, to set discriminator thresholds, to monitor pulse heights, monitor timing alignment and AMU timing, monitor HV stabilization, and monitor efficiencies.

3.4.5 R&D Results

We have studied the characteristics of resistive plate counters (RPC) and Iarocci tubes operating in both the proportional and limited streamer modes in order to make a technology choice for the PHENIX muon identifier system. The conclusion of this study is that Iarocci tubes are best suited for this application. The R&D results and technology choice are described in Ref. [39].

Pulse Height Studies

We have measured the pulse heights for signals (from cosmic and radioactive sources) using different gas mixtures, different channel widths, and different operating voltages [39]. We have studied the plateau width versus high voltage, pressure dependence of pulse height, proportional mode running, and limited streamer mode running. The most probable pulse for our baseline design is approximately 0.5 pC.

Timing Studies

The drift time interval for conventional 9 mm by 9 mm LSTs with a standard gas (Ar-isobutane) is approximately 100 ns which is too long for our application because it is of the same order as the RHIC bunch crossing time (106 ns). (The rise time and pulse duration are satisfactory, however.) We studied and achieved faster timing characteristics by ORing pairs of tubes in a four-pack staggered by half a channel. We thus are able to construct a viable muon identifier using a conventional gas. Note that an important additional reason for the stagger is that it reduces the geometrical inefficiency due to the comb profile septa from 10% to near zero (at normal incidence).

We investigated two other means of obtaining shorter drift times by using faster gas mixtures including CF₄ and by developing Iarocci tubes with a 5 mm by 5 mm cross section, instead of the standard 9 mm by 9 mm. Further details on timing studies are available in [39].

We have obtained drift time intervals for 9 mm tubes using CF₄-isobutane for which approximately 90% of the pulses arrive within an interval of 45 ns. For Ar-CO₂ the drift time interval is approximately 85 ns. Our baseline design calls for all layers to have tubes staggered by half a cell (for both vertical and horizontal tubes). The final gas mixture will be chosen in 1996.

Beam Test Results

A series of beam tests (RD93, RD94) were performed at the AGS to study the performance of Iarocci tubes [39]. A prototype muon identifier was prepared for the June 1993 AGS test beam run on the B2 line. The purpose of this prototype was to build a section of muon identifier that was more finely segmented in both longitudinal and transverse directions than planned for PHENIX. Analysis of test beam results were compared with GEANT simulations as well as simulations of performance of the planned PHENIX muon identifier configuration.

The prototype muon identifier was composed of steel plates stacked to form a sandwich of ten layers with gaps between each layer for the detection system to be inserted. The detector was chosen to be Iarocci type gas tubes operated in the limited streamer mode. The signals were read out by x -strips on one side of the tube array and by y -strips on the other side. These detector arrays were interleaved with the layers of steel absorber. Each Iarocci tube had eight $9\text{ mm} \times 9\text{ mm} \times 150\text{ cm}$ long active volumes with an anode wire at the center of each. These sections were separated one from another by 1 mm thick plastic septa. The tubes were packaged in frames that held the readout strips in position. The readout strips were 4 cm wide for both the x and the y -strips and extended the length of the Iarocci tube array. The entire assembly was supported on a large stand used for RD10. Ten 0.5 inch steel plates were stacked together to form layers approximately 12.7 cm thick. The entire assembly corresponded to 7.5 hadronic absorption lengths. Simulation studies with GEANT indicate that a uniform spacing of steel provides good μ/π discrimination. We studied other arrangements with increasing or decreasing thicknesses of steel per layer as a function of muon momenta. This arrangement produced a μ/π separation of approximately 2%. These tests also provided information concerning pulse height and gain stability.

Outstanding R&D

The panel frames have been designed. The mechanical design detailing for services (gas, high voltage, low voltage, data signals) is being completed.

The principal remaining R&D is to construct a full-size muon identifier panel. The goals are to address issues concerning optimization of the mechanical design and mass fabrication. Preparations are in progress. The prototype is to be fully constructed by May 1996 and fully tested by August 1996.

The final gas mixture has not been decided. Use of a fast gas mixture can reduce the drift time interval. We intend to continue studying the performance of different gas mixtures. We anticipate making the final selection in 1996.

Electrical tests using groups of 20 Iarocci tubes (as well as the full prototype) will be performed to study signal characteristics, optimization of the readout, and high voltage issues.

We are pursuing the possibility of mass fabrication of the tubes by CIAE at a cost significantly below commercial rates. We are pursuing tube longevity studies.

3.5 Electronics

This section will present information about the front-end electronics for both the muon tracking chambers and the muon identifier tubes. The remainder of the DAQ electronics chian is described in the CDR Update.

3.5.1 CSC Electronics

The muon tracker has approximately 16000 channels of electronics on the high resolution cathodes, 8000 channels of electronics on the anodes, and 8000 channels of electronics on the low resolution (coarse) cathodes. The three coordinate readout on each cathode strip detector allows for space point reconstruction on each detector without the requirement of using additional detectors to resolve ghost hits. The anode and low resolution cathode electronics requirements are somewhat straight forward because no amplitude or timing information is required. The high resolution cathode, however, has somewhat more stringent requirements because the resolution achievable depends on the noise levels and the linearity of each channel.

Characteristic Counting Rates and Times

To properly design the front end electronics an understanding of the strip counting rates is needed. The updated CDR lists the following parameters,

beam condition	Interaction Rate	dN/d η	time
$p + p$	454 kHz	2.6	2.2us
$Au + Au$ (MB)	13.9 kHz	210	72.2us
$Au + Au$ (Cn)	1.4 kHz	878	721.6us

The multiplicity's quoted are from the target. M.L.Brooks has determined through simulations that the multiplicity's in the muon tracker are 50 for $Au + Au$ (Cn) events. This means that only 0.045 of all tracks enter the tracker. This reduction factor could be applied to the MB events for $Au + Au$ or $p + p$. The occupancies for each strip can be determined in the following manner,

$$\begin{aligned} \text{Readout pitch} &= 1 \text{ cm} \\ \text{number of readouts} &= 6.28 \times R \text{ (R = outer radius of station)} \\ \text{occupancy} &= 3x \text{ (} 50/6.28 \times R \text{)} \end{aligned}$$

where 3x includes charge sharing on more than one strip. Therefore, for station 1 the strip occupancy is,

$$\begin{aligned} Au + Au \text{ (Cn) occupancy} &= 0.135 \text{ counts/strip/crossing event} \\ Au + Au \text{ (MB)occupancy} &= 0.03 \text{ counts/strip/crossing event} \end{aligned}$$

The ability of the detector design to resolve all ghost hits allows us also to always know that a multiple track on the high resolution cathode has occurred. If we know that a multiple track has occurred, than resolving the two tracks is possible. The only requirement is that the charge from both tracks be integrated correctly. This requirement implies that the integration time of the front end amplifier should be much longer than the drift time in the chamber gas so that variations in charge integration between the two tracks is minimized. Integration times of up to 500 ns are therefore desirable.

Another important parameter is the length of time available on average before unacceptable pileup occurs. An integrate and reset front end pileup can naturally occur and subtraction of two samples is necessary. If a front end shaper is used than the signal decay time can be adjusted to decay to a prescribed level so the probability of pileup is small. The average count rate is calculated as follows,

$$\begin{aligned} Au + Au \text{ (Cn)} &= 0.135 \times 1.4 \text{ kHz} = 189 \text{ counts/s} \\ Au + Au \text{ (MB)} &= 0.03 \times 13.9 \text{ kHz} = 417 \text{ counts/s} \\ \text{Total} &= 606 \text{ counts/s} \end{aligned}$$

Given a strip rate of 840 counts/s, the probability of 2 events occurring in 100 us is 0.002, very small. Therefore, a pulse with a long tail can persist for 100 us. The decay time allowed for a pulse to decay to less than 1 bit out of 12 bits is 12 us. The specification for signal processing using a shaper is that the rise time should be greater than 500 ns and the decay time should be less than 12 us.

CSC Detector Characteristics

The high resolution cathodes have 0.5 cm wide strips and a readout pitch at 1 cm, i.e. every other cathode is readout. The intermediate cathode signal is capacitively coupled to the readout strips and the intermediate cathode will have a large value resistor of approximately 0.5 megohm to ground to prevent the strip from floating up to high voltage. The cathode strip will be either gold or aluminum with a resistance of approximately 1 ohm per square. The capacitance of the strip is dominated by the interstrip capacitance and is approximately 1.3 pf/cm. The amplifier will see two capacitances in series and so will see one half of 1.3 pf/cm or 0.65 pf/cm. For chambers 3 m long, the maximum capacitance for the front end electronics is 150 pf. Since the strips on all stations will be variable in length, the front end capacitance will vary from 0 to 150 pf. The high resolution cathodes will be DC coupled into the amplifiers.

The low resolution cathodes will be 1 cm strips and readout at a 2 cm pitch. The strip resistance and capacitance will be 1 ohm per square and 0.65 pf/cm respectively. The strip capacitance will vary from 0 to 150 pf. The low resolution cathode will be DC coupled into the amplifiers.

The anodes are spaced at 1 cm and will be read out at 1 or 2 cm intervals depending on the occupancy. The anodes will be at high voltage, 2500 volts, and therefore the anode signals will be AC coupled to the amplifier via a 100 pf capacitor. We expect to use a fast gas and have a maximum drift time in the gas to the anode of about 50 ns.

The Detector Signals and Noise

The front end electronics on the high resolution cathode should be designed for 80 fC or less input signal. To achieve a resolution of 100 μm due to electronic noise the signal-to-noise must be less than,

$$S/N = \frac{\sigma}{1.65w} \quad (3.2)$$

where σ is the chamber resolution = 100 μm and w is the strip pitch = 1 cm. The required noise level is 0.6% or 0.5 fC. This is equivalent to 3000 electrons of noise.

Electronic Design Issues

The primary design issue for the front-end electronics is to maintain a signal-to-noise of better than 100/1 in the face of a large detector capacitance. For the CSRC chambers, the dominant contribution to the capacitance seen by the preamplifiers is the strip-to-strip capacitance. The capacitance between adjacent strips having a thickness, t , width, w , and separation, s , laying on a dielectric with constant, k , is approximately given by,

$$C\left(\frac{\text{pf}}{\text{cm}}\right) = \frac{0.12t}{w} + 0.09(1+k)\log_{10}\left(1 + \frac{2w}{s} + \frac{w^2}{s^2}\right) \quad (3.3)$$

The second term dominates. Using $k=3.5$ (kapton), $w=10$ mm, $s = 0.5$ mm, $t=2$ μm , the capacitance is 1.1 pf/cm. For the prototype test chamber, the capacitance was measured to be 1.33 pf/cm, in close agreement to the calculated value. We expect the maximum capacitance will be less than 500 pf for all stations so a basic requirement of the front-end electronics is that it must perform to specifications with an input capacitance of 500 pf or less.

An additional concern is the coupling of the signals between adjacent strips. A test of this effect was modeled in the electronic code PSPICE for representative chamber parameters. For this simulation the strips were considered as lumped RC delay lines. The results of the simulation are shown in Figure 3.34. A current pulse similar to signals observed on the prototype chamber was injected into one end of the simulated cathode strip and the pulse response was observed at the far end of the active strip as well as the adjacent strip. The total interstrip capacitance was 300 pf, the resistance was 20 ohms, and the capacitance to the anode plane was 10 pf. The results show significant coupling to the adjacent strip and dispersion of the initial signal. However, if the signal is integrated for > 100 ns the full signal charge is preserved and the signal on the adjacent strip integrates to zero. This implies that the integration time of the electronics should be greater than 100 ns if we want to be insensitive to any distortions in the apparent induced charge distribution due to coupling from one strip to another.

Dynamic Range

The dynamic range of the high resolution cathode electronics depends on the signal to noise level, energy loss fluctuations, and a margin of error required to take into account the

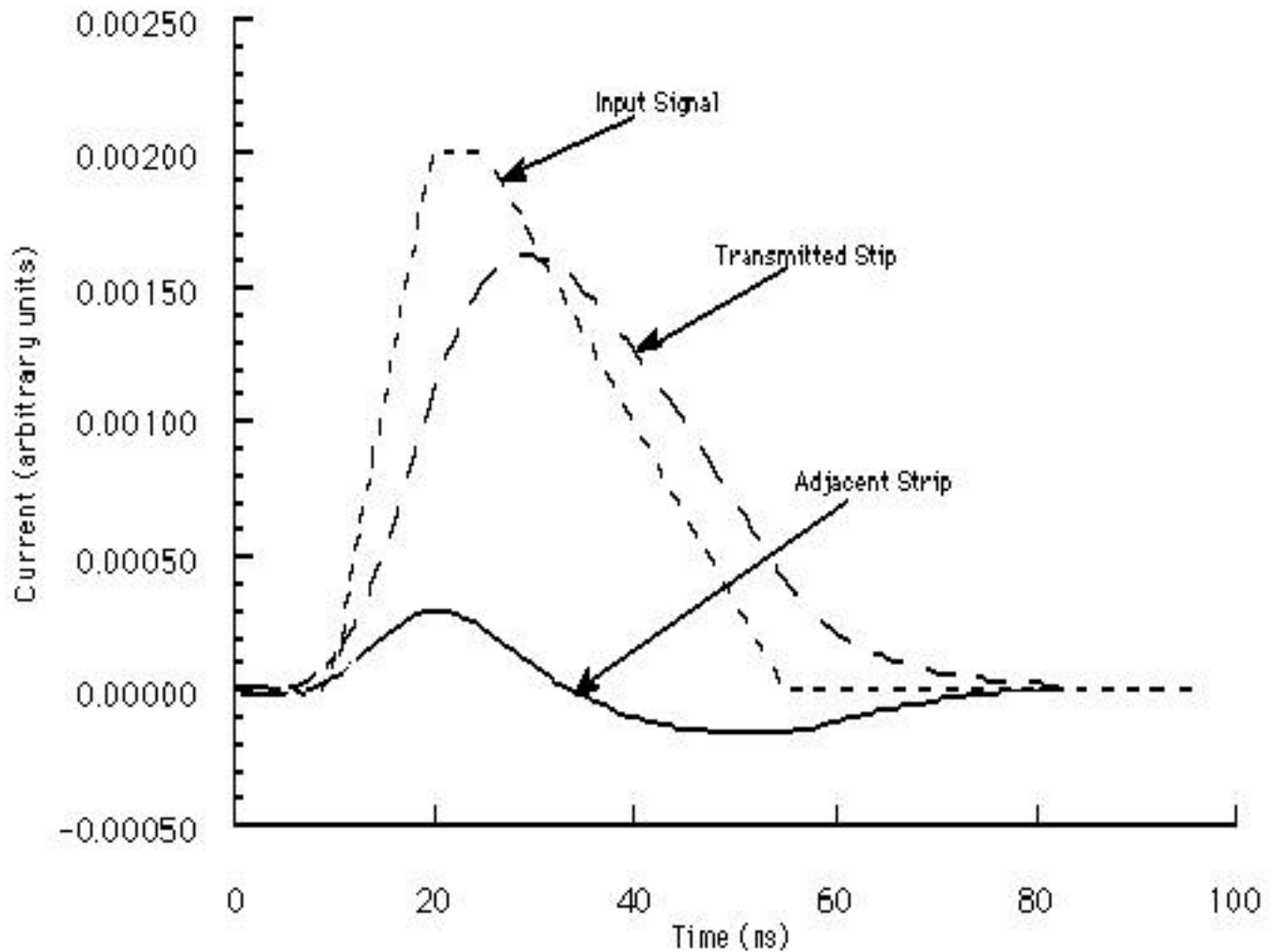


Figure 3.34: SPICE simulations of cathode strip

variations in the gain of the front end amplifiers and the following AMUs. The dynamic range is calculated as follows,

- × 100 signal to noise
- × 10 energy loss fluctuations
- × 2 margin
- 2000 minimum needed dynamic range or 11 bits

A prudent design of 12 bits would be desirable for the dynamic range of the amplifier.

The FEE Scheme

The front end electronics can be designed using either of the following methods:

1. Integrate and reset method similar to that employed by the MVD.
2. An amplifier and shaper with no need for reset.

Method 1 requires a presample to be subtracted from the signal level of interest to get the corrected signal. This subtraction automatically introduces a factor of 1.4 increase in the noise contribution to the signal. For a high resolution system this is undesirable. Additionally, previous experience with a similarly designed integrator (SVX-H) has shown that bad channels that have unreasonably high leakage currents for whatever reason adversely effect adjacent channels by saturating the bad channel and feeding additional leakage currents to the next nearest channels. The longer the integration time, the larger the number of effected channels.

Method 2 requires no presample subtraction except for a pedestal subtraction which can be accurately determined at a calibration time and subsequently stored in a computer.

The preferred FEE scheme is Method 2 because of a better noise level, no reset transients, no saturated channels to interfere with adjacent channels, and the baseline is more solid.

Low Resolution Anode and Cathode FEE

The signal to noise requirement for the anode and cathode front ends are simply that the noise level should be low enough so that at a thresold adjusted to give >99% efficiency, the noise contribution to the strip or wire singles rate should not be significantly greater than the singles rate with beam. The threshold should be a factor of three below the most probable signal. The threshold to noise is determined from the following,

$$\frac{V_t}{\nu} = \sqrt{-2 \ln(4\sqrt{3}f_n\tau)} \quad (3.4)$$

where V_t is the voltage threshold, ν is the noise level, f_n is the noise frequency, and τ is the amplifier time constant. For a time constant of 500 ns and a noise frequency of 85 counts/sec, the threshold to noise should be 4/1. If the signal to threshold is 3/1 than the signal to noise should be 12/1.

The CSC anode front end can be identical to the front end being designed for the Muon ID. The Cathode FEE is similar to the anode except the polarity is for positive signals and the amplitude is between 0.25 and 0.5 of the anode.

3.5.2 Muon Identifier FEE Specifications

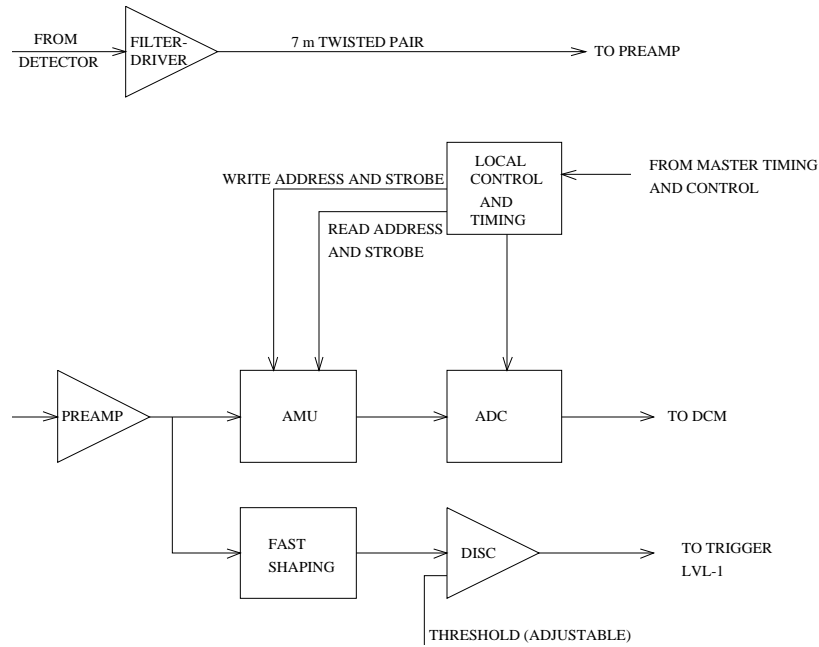
The muon identifier front-end electronic (FEE) specifications are described in Ref. [40]. We intend to operate the tubes in the proportional mode to ensure maximum longevity. There could be significant sources of electronic noise at RHIC. At the time of this writing, it is not known whether an acceptable signal-to-noise can be achieved for readout of eight 5.8 m long wires ganged passively and read out from the far end of a 7-m micro-coax cable. Therefore, the following presumes a readout scheme with an inexpensive active filter/driver located on each “external connector.” This is a custom connector card which plugs into the endcap on one end of each Iarocci tube. The external connector will have a decoupling capacitor, an off-the-shelf filter/driver, and connections to one pair of a 7 m long, twisted pair flat cable for the output.

Filter/driver An off-the-shelf active filter/driver will be located at the tube endcap. It must drive a 7 m twisted pair cable, have modest power dissipation, and be extremely reliable due to minimal access.

Front-end amplifier/integrator The front-end amplifier must be designed for negative going signals (for readout of eight ganged wires). There will be an 8-channel chip implementation. The most probable pulse is 0.5 pC. Most pulses are within three times the mean, but a non-negligible fraction will have a maximum pulse of 20-100 times the mean. The dynamic range should be 8 bits. The amplifier will be designed for a 7 m twisted pair cable between the Iarocci tube and motherboard. The front-end amplifier will be optimized for a 5 ns signal rise time and a 45 ns drift time interval. If possible, this design may also be used for the amplifier for the muon tracker CSC anode wires with minimal modification. The technology is non-radiation hard. There will be a calibration circuit and mask register. There will 8988 channels/arm.

Post amplifier A leading edge discriminator with adjustable threshold will be used to provide a fast trigger on the output current pulse from the amplifier. Its dynamic range is to be 200:1. We intend to re-use an existing PHENIX 8-bit AMU/ADC with minimal development. The AMU will be 64 cells deep to match the LVL-1 trigger delay and pending-sample storage requirements [1] [2]. For each hit, a pre- and post-integration sample will be stored in the AMU.

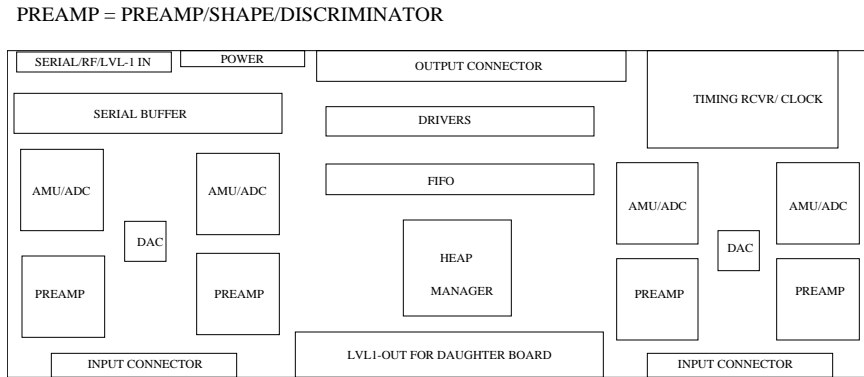
We do not intend to use the measured pulse-height in order to do particle identification. Our physics measurements will be based on digital readout information. However, we will measure and record a pulse height for every hit using an ADC in order to monitor the gain variation of every channel. This will be used for performance monitoring of the efficiency of the muon identifier LVL-1 trigger system. Also, pulse height threshold cuts can be increased offline above the discriminator settings (if warranted after study) to suppress noise. For this limited application, 8-bit accuracy is not required. However, a suitable 8-bit AMU/ADC has already been developed for PHENIX. For cost effectiveness, we will simply re-use this existing version rather than invest time or money to reduce its precision.



BLOCK DIAGRAM OF FRONT END ELECTRONICS FOR MUON IDENTIFIER.

K. Read
10/1/95

Figure 3.35: Muon identifier readout scheme.



MUID “TYPICAL” FRONT END

1 CM
1 CM

64 CHANNELS OUT OF 224
10.6 CM X 28 CM

K. Read
8/10/95

Figure 3.36: Muon identifier front-end electronics motherboard layout.

General Specifications There will be diode protection on all inputs, a total power dissipation of < 200 mW/channel, and fiber optic connections for all signals in or out of motherboard (serial, timing, control, and data output). There will be one motherboard per muon identifier panel, a total of 48 panels per arm, and, thus, 48 motherboards per arm. There will be 128 to 226 channels per motherboard depending on the panel. There will be one heap manager per motherboard. The muon identifier FEE must be able to perform the necessary local LVL-1 logic to provide the appropriate information and matching for the LVL-1 trigger system to construct x and y views. This readout scheme is shown in Figure 3.35. A conceptual layout of the motherboard is shown in Figure 3.36.

3.6 CSC Gas System

The gas system for the CSC tracking chambers will be a closed gas recirculating system designed to maintain the purity while conserving gas and preventing release to the environment. We expect that the choice of gas will be based on CF_4 because of its small Lorentz angle. Possible gas mixtures will be,

CO_2 CF_4 50%:50%
 isobutane CF_4 50%:50%
 Ar CO_2 CF_4

Studies are underway to define the preferred gas.

Specifications and Requirements

The following features define the design of the gas system,

Gas Volume

Station 1 $1.2 \times 10^5 \text{cm}^3$
 Station 2 $4.7 \times 10^5 \text{cm}^3$
 Station 3 $7.2 \times 10^5 \text{cm}^3$

- Flow Rate - 3 gas exchanges per day.

Station 1 255 cc/min
 Station 2 975 cc/min
 Station 3 1500 cc/min

- Gas overpressure < 5 torr across windows
- Gas flow must be measured both entering and exiting the chambers.
- The gas mixture must be monitored to 1% accuracy.
- The gas must eliminate contaminating gases and oil.
- Remote control and monitoring must be provided.

A prototype gas system similar to PHENIX muon needs was designed for the GEM central tracker and will be modified for use with the CSC's. This system is shown in Figure 3.37.

A recirculating system was chosen for two primary reasons, cost and environment. CF_4 is an expensive gas and venting all of the gas to the atmosphere would cost approximately \$130,000 per year. A gas recirculating system can be constructed for much less than this cost so economically it is better to capture the gas than to release it. CF_4 is presently not listed as an ozone depleting chemical or as a noxious gas but may be in the future. Release of flammable gases is controlled so a prudent design should include recirculation.

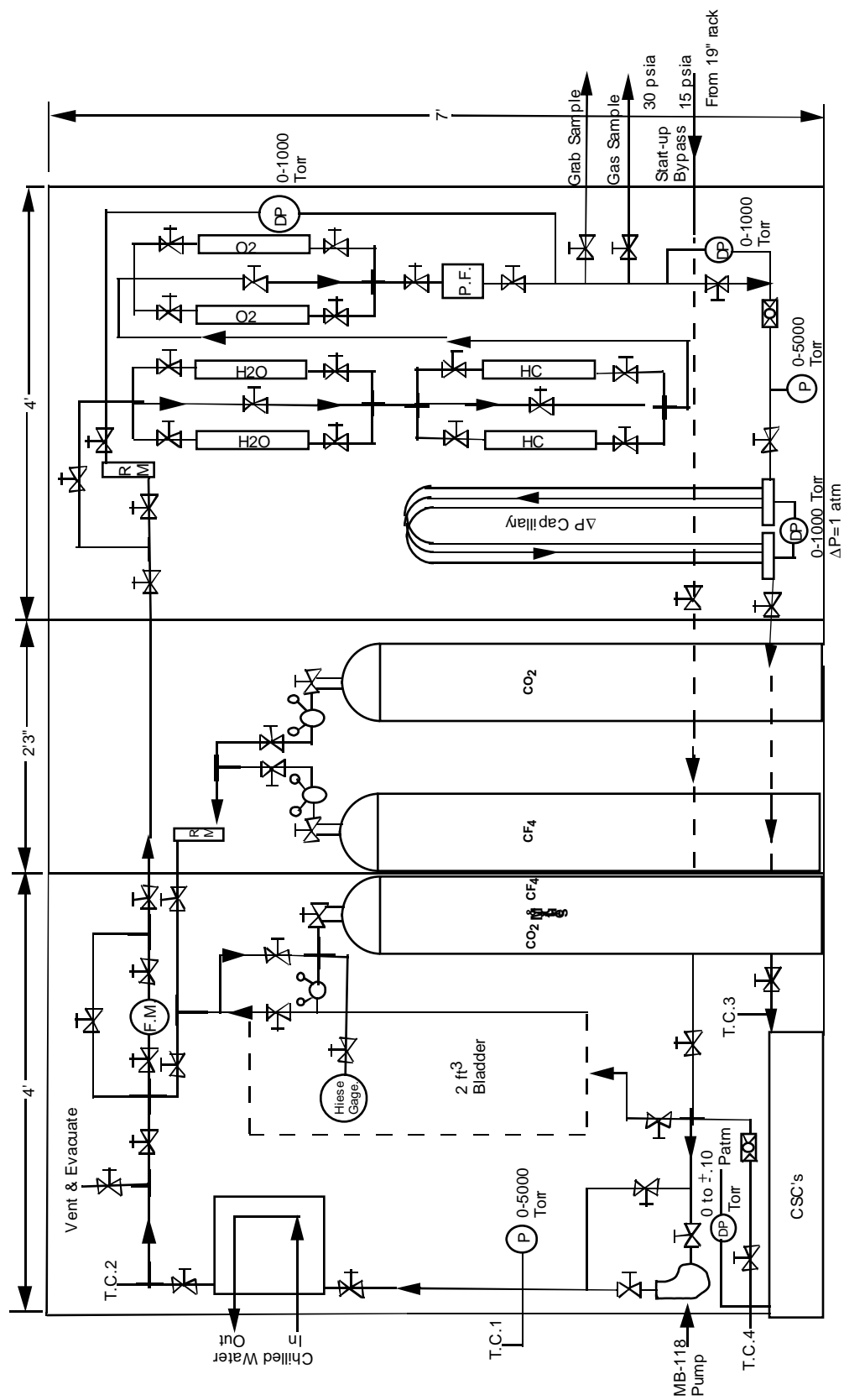


Figure 3.37: Prototype Gas System

The system is designed so that all moving parts and sensors can be located far from PHENIX so inspection and repair can be enabled during accelerator running. Redundant pumps and controllers are installed to permit operation in the event of failure of a critical part. The combination of the pumps and a pressure regulating valve controls the flow emerging from the chambers. The scrubber system will be designed for the specific chamber gas selected. The general theme will be that nitrogen, oxygen, oils, and any hydrocarbons produced in the chambers will be removed.

